

Chapter IV

Pre-heated shock experiments in the molten $\text{CaAl}_2\text{Si}_2\text{O}_8$ - $\text{CaFeSi}_2\text{O}_6$ - $\text{CaMgSi}_2\text{O}_6$ ternary: a test for linear mixing of liquid volumes at high pressure and temperature

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ABSTRACT

We performed seventeen new shock wave experiments on pre-heated (1673 K) hedenbergite liquid ($\text{CaFeSi}_2\text{O}_6$) and two model basalt liquids (an equimolar binary mix of $\text{CaAl}_2\text{Si}_2\text{O}_8 + \text{CaFeSi}_2\text{O}_6$ and an equimolar ternary mix of $\text{CaAl}_2\text{Si}_2\text{O}_8 + \text{CaFeSi}_2\text{O}_6 + \text{CaMgSi}_2\text{O}_6$) in order to determine their equations of state (EOS). Ambient-pressure density measurements on these and other Fe-bearing silicate liquids indicate that FeO has a partial molar volume that is highly dependent on composition, which leads to large errors in estimates of the densities of Fe-bearing liquids at ambient pressure based on an ideal mixing of any fixed set of end-member liquids. We formulated a series of mixing tests using the EOS determined in this study to examine whether ideal mixing of volumes might nevertheless suffice to describe the ternary system $\text{CaAl}_2\text{Si}_2\text{O}_8$ - $\text{CaFeSi}_2\text{O}_6$ - $\text{CaMgSi}_2\text{O}_6$ at high temperature and pressure. The ideal mixing null hypothesis is rejected; compositional variations in partial molar volume of FeO appear to extend to high pressure. Only densities of Fe-bearing liquid mixtures with oxide mole fraction of FeO less than 0.06 can be adequately approximated using an ideal solution.

INTRODUCTION

Silicate liquids play a large role in our understanding of both the early and modern Earth. Most widely-accepted scenarios for the evolution of the early Earth involve at least one stage of extensive or complete melting, for example, in the aftermath of impacts with large proto-planetary sized objects such as the presumed moon-forming event [*Canup*, 2004; 2012; *Canup and Asphaug*, 2001; *Ćuk and Stewart*, 2012]. The chemical and thermal evolution of the early molten mantle and how it affects the chemistry and physical properties of the modern mantle are not well understood. Furthermore, seismic observations of the 410km discontinuity in the upper mantle [*Revenaugh and Sipkin*, 1994; *Song et al.*, 2004] and ultralow velocity zones (ULVZ) in the lowermost mantle [*Garnero and Helmberger*, 1995; *Williams and Garnero*, 1996] have been interpreted to indicate the presence of a neutrally buoyant partial melt, yet there is currently limited knowledge of how the density and compressibility of complex silicate liquids behave at high temperature and pressure. It is therefore essential to have experimental constraints on the equation of state (EOS) of the complex multicomponent silicate liquids that would be actually present in nature. The difficulty lies in that natural silicate liquids comprise a potentially infinite, multi-dimensional continuum of compositions. Understanding their range of volumetric behavior requires either data on an arbitrarily large number of experiments or data on a minimum set of end-member compositions along with a reliable tool of interpolating among them. Choosing the correct method of interpolation is key to producing a predictive model for silicate liquid density and compressibility over a wide range of pressure, temperature, and composition space. Experimental data that provide constraints on end-member volumes as well as a few intermediate compositions offer the opportunity to test whether available methods for such interpolation are reliable.

Ideal mixing of volumes has proven to be a reliable method of interpolation for ambient-pressure experiments within a given composition and temperature range [*Bottinga and Weill*, 1970; *Bottinga et al.*, 1982; *Bottinga et al.*, 1983; *Guo et al.*, 2013a submitted; *Lange*, 1997; *Lange and Carmichael*, 1990; *Mo et al.*, 1982; *Nelson and Carmichael*, 1979], although exceptions from ideality exist and are discussed below. Likewise, ideal mixing silicate liquid volumes at high pressure has been employed as a simplifying assumption, but it in fact remains poorly tested and constrained [*Asimow and Ahrens*, 2010]. In this study, we report seventeen new shock wave experiments to pressures up to 146 GPa on preheated silicate liquids of three separate compositions: hedenbergite (Hd; $\text{CaFeSi}_2\text{O}_6$), a 50-50 mixture of anorthite and hedenbergite (An-Hd; $\text{CaAl}_2\text{Si}_2\text{O}_8\text{-CaFeSi}_2\text{O}_6$), and an equal mixture of anorthite, hedenbergite, and diopside (An-Hd-Di; $\text{CaAl}_2\text{Si}_2\text{O}_8\text{-CaFeSi}_2\text{O}_6\text{-CaMgSi}_2\text{O}_6$). Determination of the EOS of Hd and the two mixtures (or “model basalts”) permits us to test the validity of linear mixing of volumes and ideal configurational mixing of entropies — two assumptions which were both used in *Thomas et al.* [2012] and Chapter III of this text to predict the volumes of multicomponent liquids at high T and P .

METHODS

Sample preparation

The synthesis and analysis for the Hd and model basalt samples used in this study are described in *Guo et al.* [2013a submitted] and *Guo et al.* [2013b in prep], respectively. These studies also describe the 1-bar double-bob Archimedean density measurements and the ultrasonic sound speed measurements that were completed on these materials. The sample materials used for the shock wave experiments were cored directly from the lower buffer rod crucible used in

the sound speed determinations. The cores were sliced into disks, which were then lapped down to under-fill the volume of the molybdenum shock-experiment sample holder. The Mo sample holders are assembled from two machined pieces: 1) the driver plate, which has a hollow cylindrical projection on one side that becomes the sample well and 2) the cap, which is a 1 mm-thick, 8 mm-diameter disk that is welded on after filling (for complete description of the welding process see *Thomas et al.* [2012]). The interior volume of the sample well is 8 mm in diameter and 2.5 mm deep. The samples under-filled this volume by 7-13% at room temperature to prevent the welded cap from bowing during heating [c.f. *Thomas et al.*, 2012]

The surfaces within the sample well including the inside surface of the cap were polished to a mirror finish with 1-micron alumina grit. Smoothing all the surfaces prevents bubbles from clinging to the walls of the sample holder during heating and potentially interfering with the imaged shock wave. The outer surface of the cap was left with the finish quality as-machined (~9-micron finish), in order to match the quality of polish on the inner portion of the rear-face of the driver plate, which was polished on a rotating stage. The speed of the polishing pad decreases towards the center, hence the quality of mirror finish also decreases. Approximately similar finishes on the cap and driver reflecting surfaces ease acquisition of evenly illuminated images of the cap and driver during the shot and greatly increased the precision of the Hd data given below. On the contrary, the model basalt capsules were polished as in the manner as described in Chapter II due to the larger quantity of light needed for the digital recording system (c.f. Chapter III). Hence this dataset is plagued with larger errors derived from determining the driver and cap arrivals with different exposure levels. An improved technique for polishing for driver to a mirror finish would greatly aid data acquisition in the future. Additional sample preparations are detailed in *Asimow and Ahrens* [2010], *Thomas et al.* [2012], and Chapter III.

Experimental setup

The pioneering work for shock studies on molten materials is *Rigden et al.* [1984]. The description of our methods below builds on this and the work of *Rigden et al.* [1988; 1989], *Miller et al.* [1988; 1991a], *Chen and Ahrens* [1998], and *Chen et al.* [2002]. The most recent changes to experimental techniques and data analysis are in *Asimow* [2008], *Asimow and Ahrens* [2010], *Thomas et al.* [2012], and Chapter III.

For this study, seventeen total experiments were performed—twelve in the Caltech 90/25-mm two-stage light-gas gun and five using the Caltech 40-mm propellant gun—eight performed on Hd, five performed on An-Hd, and four performed on An-Hd-Di (see Table 1). All shots were preheated to $1673 \text{ K} \pm 4 \text{ K}$ (1400°C) and used either molybdenum (Mo), tantalum (Ta), or aluminum alloy (Al2024) flyer plates. Flyer plate velocities (u_{fp}) ranged from 1.46 – 6.19 km/s; the methods used for measuring u_{fp} are given in *Thomas and Asimow* [submitted].

During each experiment, the rear face of the target is illuminated by a Specialised Imaging xenon spark lamp and filmed by a Hadland Imacon 790 streak camera through a narrow ($25 \mu\text{m}$) slit focused horizontally across the center of the driver and sample cap. Hd shots were recorded on Polaroid film; the model basalts were recorded using a digital system (c.f. Chapter III). As the shock wave reaches the free surface of the driver or the cap the reflected light from the polished surface is extinguished (or shows a sharp change in intensity). The offset time of these two cutoffs allows for the precise measurement of the shock transit time through the sample and molybdenum cap.

To correct for non-uniform streaking rate, the camera was calibrated using a test streak image modulated by a radio-frequency tuner at 148.9875 MHz for a 1500 ns streak length and by a pulse generator at 1.7870 MHz for a 5000 ns streak (shot 1068). Each pixel line on the shot

streak record can then be converted to a time during the experiment, which permits the shock transit time through the sample to be calculated from the pixel distance between the two cutoffs on the film (or digital image). A detailed method for picking cutoffs and reducing the data is described in *Thomas et al.* [2012].

The final calculated shock state — i.e., shock pressure (P_H), shock density (ρ_H), particle velocity (u_p), and shock velocity (U_s) is solved for iteratively [*Rigden et al.*, 1988] using measured shock wave travel time, impedance matching, the Rankine-Hugoniot equations, Hugoniot data (ρ_o , C_o , s) for the metal flyer plates and hot Mo driver (Table 2), and the initial density (ρ_o) for each of the silicate liquids (see Tables 3, 4, and 5). An initial guess for the bulk sound speed (C_o) and Hugoniot slope s is required to seed the iteration, but the converged result is independent of this assumption.

RESULTS

The shock wave data are reported in Table 1, including shot number, flyer material, the temperature prior to firing, u_{fp} , u_p , U_s , ρ_H , and P_H .

A Hugoniot is a series of peak shock states achieved by progressively stronger shocks in a material from the same initial conditions [*Ahrens*, 1987]. The Hugoniot of a well-behaved material empirically forms a line in U_s - u_p space, given to third order in strain by $U_s = C_o + s u_p$ [*Jeanloz*, 1989]. The bulk sound speed of the material at room pressure, C_o , is given by the intercept; the slope (s) is related to K'_s , the isentropic pressure derivative of the isentropic bulk modulus (K_{os}) by [*Ruoff*, 1967],

$$s = \frac{(K'_s + 1)}{4}. \quad (1)$$

Hedenbergite ($\text{CaFeSi}_2\text{O}_6$) liquid

Figure 1 shows linear fits to the Hd shock data in $U_S - u_p$ space. An unweighted linear fit to all eight preheated hedenbergite liquid data points yields $U_S = 2.587 \pm 0.142 + 1.55 \pm 0.04 u_p$ km/s ($r^2 = 0.996$). The experimentally determined ultrasonic value of the bulk sound speed, 2.613 ± 0.016 km/s [Guo *et al.*, 2013a submitted], lies within the stated uncertainty of the bulk sound speed derived from the intercept of this fit, 2.587 ± 0.142 km/s. This relationship indicates relaxed (liquid-like) — as opposed to un-relaxed (glass-like) — behavior upon shock compression [Rigden *et al.*, 1988]. Therefore, to reduce the error on the linear fit, we fixed the intercept at the mean ultrasonic value yielding an unweighted constrained linear fit of $U_S = 2.613 \pm 0.016 + 1.54 \pm 0.01 u_p$ km/s ($r^2 = 0.999$).

It is necessary to select and apply a thermal equation of state formalism for investigating material properties that lie off the Hugoniot, since the Hugoniot achieves temperatures and energies much higher than those of geophysical interest at lower mantle pressures (even for processes during early Earth differentiation). We have defined the entire P - V - E surface of hedenbergite liquid using various formalisms: the shock wave equation of state (SWEOS) and the 3rd- and 4th-order Birch-Murnaghan Mie-Grüneisen equations of state (3BM/MG and 4BM/MG). For a full description of the thermal EOS equations and fitting procedures, the reader is directed to Asimow and Ahrens [2010] and Thomas *et al.* [2012]. The results and uncertainties for each fit on Hd are given in Table 3, and the Hugoniots are plotted in Figure 2.

The SWEOS is defined by a linear Hugoniot in $U_S - u_p$ space, converted to $P - \rho$ space using the first and second Rankine-Hugoniot equations [e.g., Ahrens, 1987]. States that lie off the Hugoniot are found using the Mie-Grüneisen thermal pressure approximation with a temperature-independent power law expression for the thermodynamic Grüneisen parameter

$$\gamma(\rho) = \gamma_o \left(\frac{\rho_o}{\rho} \right)^q \quad (2)$$

Although a value for q to be used with the SWEOS was not independently determined in this study, experimental results thus far indicate that silicate liquids—including iron-bearing liquids [Thomas *et al.*, 2012]—appear to have q values of -0.88 to -2 for compressions of $1 > \rho_o/\rho > 0.49$. The slope and intercept of the constrained Hugoniot fit given above corresponds to a K'_s value of 5.16 ± 0.04 derived from (1) and a K_{oS} value of 19.98 ± 0.24 GPa derived from the expression

$$C_o = \sqrt{\frac{K_{oS}}{\rho_o}} \quad (3)$$

The 3rd- or 4th-order BM/MG EOS is defined by a 3rd- or 4th-order Birch-Murnaghan isentrope centered at 1 bar and 1673K and a Mie-Grüneisen thermal pressure approximation. The 3BM/MG fit result is $K'_S = 6.22 \pm 0.55$, $q = -1.93 \pm 0.41$, and reduced $\chi^2 = 4.20$. The 4BM/MG fit result is $K'_S = 3.15 \pm 1.67$, $K''_S = 0.78 \pm 0.60$ GPa⁻¹, $q = 0.47 \pm 9.86$ and reduced $\chi^2 = 3.26$. The high precision for most of the points in the Hd dataset propagates into unusually high χ^2 values for both fits. Nevertheless, comparatively the 4BM/MG has very large error bars on and strong correlations among the output parameters indicating a very unstable fitting routine. The 3BM/MG fit, by contrast, has less severe correlation between the parameters and appears justified by the fitting statistics. Hence for this data set we prefer the 3rd-order fit.

An-Hd liquid

The An-Hd Hugoniot in U_S - u_p space is shown in Figure 3. The unweighted linear fit is $U_S = 2.742 \pm 0.282 + 1.40 \pm 0.09 u_p$ km/s ($r^2 = 0.986$). As observed above for Hd liquid, the An-Hd liquid appears to be relaxed upon compression as the experimentally determined bulk sound speed, 2.772 ± 0.011 km/s [Guo *et al.*, 2013b in prep], and the intercept are within error. We can then fix the intercept at the mean ultrasonic value yielding an unweighted constrained linear fit of $U_S = 2.772 \pm 0.011 + 1.39 \pm 0.03 u_p$ km/s ($r^2 = 0.998$). This slope and intercept constrain the SWEOS K'_S and K_{oS} values to be 4.54 ± 0.12 and 21.30 ± 0.25 GPa, respectively.

The 3BM/MG fit result is $K'_S = 5.49 \pm 0.65$, $q = -0.18 \pm 2.25$, and reduced $\chi^2 = 6.37$. The 4BM/MG fit result is $K'_S = 5.10 \pm 1.61$, $K''_S = -0.17 \pm 0.60$ GPa⁻¹, $q = 1.67 \pm 2.79$ and reduced $\chi^2 = 12.53$. Both fits have relatively large error bars and large reduced χ^2 due to the low number of data points and fairly significant scatter in P - ρ space (Figure 4). Yet both fits give similar K'_S to SWEOS, which is well defined by the linear fit. We therefore prefer the 3rd-order fit despite having large errors, since it fits the data sufficiently well, has a lower reduced χ^2 , and exhibits a negative q value, which has been shown to be typical for silicate liquids (see Chapter III). The EOS parameters and their uncertainties for An-Hd liquid are given in Table 4.

An-Hd-Di liquid

The An-Hd Hugoniot in U_S - u_p space is shown in Figure 5. The unweighted linear fit is $U_S = 2.823 \pm 0.197 + 1.52 \pm 0.09 u_p$ km/s ($r^2 = 0.995$). Behaving as the two liquids above, An-Hd-Di liquid is relaxed upon compression [Guo *et al.*, 2013b in prep], which justifies an unweighted constrained linear fit of $U_S = 2.846 \pm 0.014 + 1.54 \pm 0.02 u_p$ km/s ($r^2 = 0.999$). This slope and

intercept constrain the SWEOS K'_S and K_{oS} values to be 5.15 ± 0.06 and 22.03 ± 0.25 GPa, respectively.

The An-Hd-Di liquid was fit only with the 3BM/MG, as fitting the 4BM/MG to so few data points is under-constrained and does not yield a meaningful reduced χ^2 . The 3BM/MG result is $K'_S = 6.20 \pm 0.17$, $q = -1.14 \pm 0.79$, and reduced $\chi^2 = 0.39$. The results and uncertainties for each fit on An-Hd-Di are given in Table 5, and the Hugoniot plots are plotted in Figure 6.

DISCUSSION

Linear mixing assumptions

To make progress on calculating the volumetric properties of the natural multi-component silicate liquids that may occur in the lower mantle, the plan has been to constrain the EOS for a number of important end-member liquids using shock wave studies and then to attempt to reliably interpolate among them. End-member liquids that have been determined previously using shock wave methods are: Fe_2SiO_4 (Fa, fayalite), Mg_2SiO_4 (Fo, forsterite), MgSiO_3 (En, enstatite), $\text{CaAl}_2\text{Si}_2\text{O}_8$ (An, anorthite), and $\text{CaMgSi}_2\text{O}_6$ (Di, diopside). The recommended EOS parameters for each of these end members are given in Chapter III. In *Thomas et al.* [2012], these five end members were used to interpolate within the CaO-MgO- Al_2O_3 - SiO_2 -FeO major element component space to constrain the volume of a liquid along an isentropic temperature-pressure path for a desired composition, yet there was not a clear test of the validity of the two underlying assumptions: (1) linear mixing of volumes and (2) constant entropy of mixing for a given composition.

The assumption of linear or ideal mixing is simplest assumption that can be made for mixtures,

$$V_{tot} = \sum_i X^i V^i + V_{mix} , \quad (4)$$

where the volume of mixing (V_{mix}) is assumed to be zero, and hence the total volume (V_{tot}) is just a linear combination of the end-member partial molar volumes (V^i) multiplied by their respective mole fractions (X^i). The second assumption is that the configurational entropy of a liquid solution is only a function of composition and therefore does not change with variations in temperature and pressure, such that

$$S_{tot} = \sum_i X^i S^i + S_{mix} , \quad (5)$$

where the entropy of mixing (S_{mix}) is assumed to be a constant and hence an isentrope or constant entropy path in T - P space can be found from

$$dS_{tot} = d \sum_i X^i S^i = \sum_i X^i \left(C_p^i d \ln(T) - \frac{\partial V^i}{\partial T} dP \right) = 0, \quad (6)$$

where the T and P derivatives of entropy come from standard thermodynamic identities. Our data do not provide a test of this method of computing isentropes, since we lack experimentally defined isentropes for reference. We therefore focus on evaluating the volume of mixing assumption by comparing isentropes for intermediate compositions constructed with different sets of end members, while acknowledging that this exercise may be compromised to some extent by the configurational entropy issue.

Glass and 1-bar melt data

Both of these assumptions have historically performed fairly well at 1 bar. Linear mixing of temperature-independent partial molar heat capacities has been a successful model for most silicate liquids [Stebbins *et al.*, 1984]. Similarly, linear mixing of volumes has been successful in

determining silicate liquid densities in most major element space ($\text{SiO}_2\text{-Al}_2\text{O}_3\text{-MgO-CaO-Na}_2\text{O-K}_2\text{O}$) [Lange, 1997; Lange and Carmichael, 1990]. However, exceptions to both rules do exist. Heat capacities for the Fe_2O_3 oxide component were shown to vary strongly with composition [Lange and Navrotsky, 1992] and the partial molar volumes of Fe_2O_3 and TiO_2 components are both variable with composition. These exceptions are attributed to the Fe^{3+} and Ti^{4+} cations having multiple coordination states at ambient pressure that are highly dependent on the composition, i.e., fraction of non- SiO_2 components [Liu and Lange, 2001; 2006].

More recent results have found this to be potentially true for Fe^{2+} as well, in that the partial molar volume of FeO (V_{FeO}) in a melt is observed to depend on composition in some cases [Guo *et al.*, 2013a submitted; 2013b in prep]. Such behavior of Fe^{2+} is well documented in the glass literature, which is important insofar as glasses provide a proxy for structural properties of melt at the glass transition temperature. When discussing melt structure, language is typically borrowed from random network theory [Zachariasen, 1932], which distinguishes oxides as network formers, network modifiers, or intermediates. Network-forming cations (e.g., Si^{4+}) occur typically in tetrahedral ([4]) oxygen coordination in various connected units in the melt, whereas network modifiers (e.g., CaO , MgO) typically occupy octahedrally ([6]) coordinated sites and disrupt the tetrahedral network. Intermediates (e.g., Al_2O_3 , TiO_2) can take on either role, may occupy multiple coordination sites, and may either disrupt or stabilize the polymerization of a melt [Bottinga and Weill, 1972; Mysen, 1983; 1988]. This is a simplified view of glasses and melts, but it acts to guide our intuition concerning the behavior of cations and the contributions of Fe^{3+} and Ti^{4+} to the density, heat capacity, and other thermodynamic properties of silicate melts [Burns, 1993; Mysen, 1988].

FeO is an intermediate glass former and has been documented to take on a range of [4] to [6] coordination dependent on composition. Fe^{2+} is predominantly a network modifier in basaltic glasses [Burns, 1993; Jackson *et al.*, 2005; Mysen, 1983; 1988] and also likely in melts with similar SiO_2 mole fraction. This is supported by the relatively low 1-bar V_{FeO} value ($\sim 12.1 \text{ cm}^3/\text{mol}$) for An-Hd, An-Di-Hd, and Hd-Di model basalts at 1732 K. This V_{FeO} has been inferred to indicate an average Fe^{2+} coordination of $\sim[5.7]$ due to its similarity to the molar volume of wüstite, crystalline FeO (in which Fe^{2+} is octahedrally coordinated) [Guo *et al.*, 2013b in prep]. On the contrary, Fe_2SiO_4 glasses contain very low-coordinated Fe^{2+} and are surprisingly polymerized [Cooney and Sharma, 1990]. X-ray absorption spectroscopy studies of both glass and melt of fayalite composition infer an average coordination of $\sim[4]$ [Jackson *et al.*, 1993]. This is again supported by melt densities, where V_{FeO} increases to $\sim 14\text{-}17 \text{ cm}^3/\text{mol}$ [Mo *et al.*, 1982; Shiraishi *et al.*, 1978; Thomas *et al.*, 2012], similar to the molar volume of crystalline CaFeO_2 , a structure in which Fe^{2+} is [4] coordinated [Guo *et al.*, 2013a submitted].

Between these two extremes, Hd and the CaO-FeO- SiO_2 (CFS) melts measured in Guo *et al.* [2013a submitted] display FeO in an average coordination state intermediate between [4] and [6]. The V_{FeO} for Hd melt ($\sim 15.47 \text{ cm}^3/\text{mol}$) is inferred to indicate [4.7] coordination, similar to the values in Hd glasses (~ 4.3) and in molecular dynamics simulation of this melt [Rossano *et al.*, 2000]. The inferred coordination for CFS glasses range from 4.6-5.2. Guo *et al.* [2013b in prep] present an inverse correlation between CaO concentration and average Fe^{2+} coordination number. This is contrary to expectation from the model in Jackson *et al.* [2005] based on valence bond (VB) theory, which suggests a direct correlation between CaO concentration and FeO coordination. In the VB model, the Ca would have a greater “share” of the O valence due to the

higher bond valence of the Ca-O bond and force Fe to find charge balance with other oxygens and hence increase coordination.

The composition dependence of Fe coordination in CFS melts appears to be more strongly related to the degree of polymerization (Si/O ratio) [Burns, 1993]. For example in Hd, where SiO₂ is high (Si/O = 0.32) and there are equal proportions of CaO and FeO, FeO fills both roles of network former and modifier, with an average value of [4.7]. For the family of CFS melts studies by Guo *et al.* [2013b in prep], the Fe content is fixed (40 mol %) and so increasing CaO is coupled to decreasing SiO₂. The Fe²⁺ may be energetically forced to play the role of network former as CaO competes more effectively for octahedral sites. For fayalite liquid, the Si/O value is quite low (0.25), and FeO is again energetically favored as a network former to accommodate the charge balance of the oxygen, creating a highly polymerized glass as observed by Cooney and Sharma [1990]. Qualitatively, one can understand the importance of FeO as an intermediate behavior cation when Fe₂SiO₄ is compared to Mg₂SiO₄ melt, which has the same Si/O ratio. Since MgO is unable to act as a network former, Mg-rich melts typically have fully depolymerized structure and display very low viscosities.

The dependence of FeO coordination and molar volume on other cation concentration invalidates the assumption of linear mixing of volumes at 1 bar for FeO-bearing liquids. It is unclear, however, what effect pressure and temperature may have. For glasses, higher temperatures are believed to decrease the coordination of transition metals [Jackson *et al.*, 2005], and there is some evidence that cations in melts will favor a coordination state lower than that of the glass [Jackson *et al.*, 1993]. At low pressure, network forming SiO₄ tetrahedra predominate, but as P increases, Si⁴⁺ coordination increases (~6.5 at 150 GPa) [Karki *et al.*, 2007]. Less is known about the coordination of network modifiers at high pressure. Preliminary modeling of

Fe_2SiO_4 liquid shows decreasing Fe-O bond length and increasing coordination of Fe with pressure, but this does not constrain the effect of other cations [Muñoz Ramo and Stixrude, in review]. Guo *et al.* [2013a submitted] and Guo *et al.* [Guo *et al.*, 2013b in prep] provide indirect evidence that coordination of Fe^{2+} in basaltic melts may *decrease* towards [5] at moderate pressures ($\sim 5\text{GPa}$).

The only previous test of linear mixing at high pressure and temperature was for Fe-free anorthite-diopside liquids [Asimow and Ahrens, 2010], which found the linear approximation suitable at elevated T - P conditions along the measured Hugoniot. Linear mixing appeared to fail at low temperatures, but this may have been an artifact of uncertainties in the large extrapolation downwards from the Hugoniot temperature. Another partial success for linear mixing was found in MD simulations of An-Di and also in MD simulations of the MgO-SiO_2 binary, each of which showed well-behaved mixing at high pressure which began to break down at lower pressure [de Koker *et al.*, 2013; Martin *et al.*, 2012].

Tests of linear mixing

Our first test examines whether the V_{FeO} difference between Fe_2SiO_4 and $\text{CaFeSi}_2\text{O}_6$ at 1 bar is still significant at high temperature and pressure. That is, can V_{FeO} for either end member substitute for the other, as it would in the case of ideal mixing of volumes? To test this we compare the Hd isentrope from shock wave experiments determined in this study to a modeled linear combination of end-member liquids, where $\text{Hd}_{\text{modelFa}} = \text{Di} + 0.5\text{Fa} - 0.5\text{Fo}$. The equations and details of creating isentropes of mixtures are given in Thomas *et al.* [2012]. The results are shown in Figure 7, where it is apparent that the model does not capture the behavior of the Hd isentrope. The errors for temperature and density are highly correlated such that if the model

predicts a lower temperature than the actual isentrope, it will over predict the density and vice versa. Such is the case for Hd, where the errors at high pressure (~ 120 GPa) and temperature are maximized, and the model underestimates the temperature by 30% and overestimates the density by 20%. In this test, comparing the use of two Fe-rich end members, we find a failure of linear mixing and this is presumably due to complexity in the coordination state of Fe^{2+} .

Our second test examines the question: Which model will perform better in creating an Fe-rich model basalt, An-Hd — using V_{FeO} derived from Fa or V_{FeO} derived from Hd? Our expectation from 1 bar coordination numbers is that since Hd should be a better end member for modeling An-Hd. Two mixing models, $\text{AnHd}_{\text{modelFa}} = 0.5\text{An} + 0.25\text{Fa} - 0.25\text{Fo} + 0.5\text{Di}$ and $\text{AnHd}_{\text{modelHd}} = 0.5\text{An} + 0.5\text{Hd}$, are compared to the An-Hd isentrope determined in this study. It can be seen in Figure 8 that neither model is able to recover the An-Hd isentrope, but that the Fa model does significantly worse. The Fa model also becomes unphysical at high pressure and lower temperature (Fig. 8a) and this breaks the anticorrelation between errors in T and ρ , which are both underestimated, by 15% and 10% at 120 GPa, respectively (Fig. 8b). The Hd model does significantly better with maximum errors at 120 GPa that overestimate T by 6% and underestimate ρ by 7% (Fig. 8d). These errors worsen systematically for isentropes with progressively higher potential temperature (Fig. 8c). These results are in line with our expectations based on 1-bar results: neither Fa or Hd models can fully recover the behavior of the An-Hd isentrope, but closer similarity of V_{FeO} at 1-bar for Hd and An-Hd are significant enough to improve the fit at high pressure and temperature.

Our third test examines the question: As FeO content decreases, does the difference between the end-member V_{FeO} and that of the model basalt still play as significant of a role in the final volume of the mixture at high pressure and temperature? We attempt to recover the

isentrope of the model basalt, An-Hd-Di (which has less FeO than the previous An-Hd basalt), again using Fa and Hd end-member models: $\text{AnHdDi}_{\text{modelFa}} = 0.33\text{An} + 0.667\text{Di} + 0.167\text{Fa} - 0.167\text{Fo}$ and $\text{AnHdDi}_{\text{modelHd}} = 0.33\text{An} + 0.33\text{Hd} + 0.33\text{Di}$. The results in Figure 9 show improvement for both models with the Fa model having high-pressure errors that underestimate T by 12% and underestimate ρ by 4% and the Hd model overestimating T by only 3% and underestimating ρ by 3%. Again, the Hd model does significantly better than the Fa model, likely due to the greater structural similarity of FeO in Hd and in An-Hd-Di at 1 bar. This test also suggests, at least in the studied region of $\text{CaO-FeO-MgO-Al}_2\text{O}_3\text{-SiO}_2$, that the linear mixing approximation becomes better as the amount of FeO in the mixture decreases. That is, despite the variable behavior of V_{FeO} , there may be a threshold concentration of FeO below which this is negligible and the ideal solution approximation is adequate.

It still remains to define best choice for V_{FeO} (i.e., Fe^{2+} coordination state) for modeling a particular liquid composition in the mantle and to define the threshold FeO concentration below which this choice is unimportant. Our final test attempts to answer these questions by comparing the original chondrite model given in Chapter III that uses Fa as the V_{FeO} component ($\text{Ch}_{\text{modelFa}} = 0.62\text{En} + 0.24\text{Fo} + 0.08\text{Fa} + 0.04\text{An} + 0.02\text{Di}$), and a model which uses An-Hd-Di as the V_{FeO} component ($\text{Ch}_{\text{modelAnHdDi}} = 0.62\text{En} + 0.32\text{Fo} + 0.45\text{AnHdDi} - 0.12\text{An} - 0.28\text{Di}$). The results in Fig. 10 show that the An-Hd-Di model predicts a slightly higher temperature but that models are within $\pm 1.5\%$ for both T and ρ . Hence, the models are effectively indistinguishable, indicating that liquids with $X_{\text{FeO}} \leq 0.06$ (chondrite model [Andrault *et al.*, 2011]) can be suitably approximated with linear mixing. Mole fractions greater than this amount, such as An-Hd-Di ($X_{\text{FeO}} = 0.08$), can be approximated within reasonable error ($\pm 3\%$), but only when using an FeO component that has a similar V_{FeO} at 1 bar.

Direct knowledge of the coordination state of Fe^{2+} in a given liquid thus is an important constraint on the volume behavior of that liquid, especially for high FeO contents. Conversely, precise data on the equation of state of an Fe-bearing liquid can be interpreted primarily as a constraint on the coordination state of Fe^{2+} in that liquid. However, liquids containing Fe_2O_3 , Na_2O , K_2O , varying amounts of SiO_2 , and other oxide species are far more complex than the five-component system studied so far. Other cations may influence the Fe^{2+} coordination state and may have variable coordination states themselves. *In situ* experiments analyzing cation coordination of complex glasses and melts at high pressure and temperature would greatly aid in elucidating these consequences and complementing macroscopic constraints from shock wave or other equation of state measurements.

CONCLUSIONS

We completed seventeen new pre-heated shock wave experiments on Hd, An-Hd, and An-Hd-Di liquids to determine the EOS of each composition. Having data on more compositions than are necessary to span the multicomponent oxide system, we applied these data to testing the validity of ideal mixing of volumes as a method of interpolation among end-member liquids. Ambient-pressure density measurements on these and other Fe-bearing silicate liquids indicate that FeO has a molar volume (and therefore coordination state) that is highly dependent on composition [Guo *et al.*, 2013a submitted; 2013b in prep], and our results show that this behavior extends to high pressure. An Fe-bearing end member close in composition to the liquid of interest is necessary for adequately predicting temperature and density along its isentrope. We find that deviations from linear mixing of volumes are significant at the level of several percent for liquid mixtures with molar $X_{\text{FeO}} > 0.06$. On the other hand, densities and isentropic gradients

for liquids in the system $\text{CaO-FeO-MgO-Al}_2\text{O}_3\text{-SiO}_2$ with less than this threshold concentration of FeO can be approximated both at ambient and at high pressure using ideal volume of mixing.

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TABLES

Table 1. Shock compression data

		T (K)	u_{fp}		u_p		U_s		ρ_H		P_H		
Shot #	Flyer material		km s ⁻¹		km s ⁻¹		km s ⁻¹		Mg m ⁻³		GPa		
Hedenbergite													
1068 ^a	Mo	1673	1.976	0.002	1.58	0.01	5.12	0.11	4.27	0.05	23.8	0.4	
413	Al	1673	4.471	0.002	2.006	0.005	5.56	0.07	4.62	0.04	32.9	0.3	
419	Mo	1674	3.497	0.001	2.68	0.02	6.59	0.24	4.99	0.17	52.2	1.5	
418	Mo	1673	4.066	0.148	3.05	0.01	7.50	0.06	4.98	0.03	67.6	0.4	
417	Ta	1677	4.535	0.005	3.65	0.01	8.29	0.10	5.29	0.05	89.1	0.6	
416	Ta	1672	5.144	0.012	4.10	0.01	9.01	0.09	5.42	0.06	109.1	0.9	
415	Ta	1673	5.613	0.001	4.46	0.01	9.46	0.05	5.59	0.03	124.5	0.5	
414	Ta	1675	6.187	0.020	4.88	0.01	10.10	0.05	5.72	0.03	145.6	0.9	
An-Hd													
1070 ^a	Mo	1673	2.006	0.003	1.62	0.01	5.17	0.12	4.04	0.05	23.2	0.4	
1074 ^a	Mo	1673	1.456	0.002	1.211	0.014	4.27	0.28	3.87	0.13	14.3	0.8	
466	Mo	1675	3.618	0.003	2.82	0.01	6.56	0.15	4.87	0.10	51.3	0.9	
469	Mo	1673	4.700	0.003	3.55	0.01	7.99	0.10	4.99	0.06	78.7	0.8	
471	Mo	1674	5.815	0.016	4.37	0.01	8.66	0.11	5.59	0.08	104.8	1.0	
An-Hd-Di													
1069 ^a	Mo	1673	2.016	0.003	1.63	0.03	5.48	0.47	3.88	0.10	24.4	2.1	
1071 ^a	Mo	1673	1.470	0.005	1.20	0.01	4.77	0.30	3.64	0.09	15.6	0.8	
468	Mo	1673	3.503	0.008	2.70	0.01	7.00	0.07	4.43	0.03	51.5	0.4	
470	Mo	1674	4.796	0.006	3.61	0.02	8.31	0.16	4.81	0.09	81.7	1.3	

^a Caltech 40mm propellant

Table 2. Parameters used

		Molybdenum ^a (2273 K)	Molybdenum ^a (300 K)	Tantalum ^b (300 K)	Aluminum 2024 ^c (300 K)
ρ_0	Mg m ⁻³	9.785	10.21	16.65	2.78
C_0	km s ⁻¹	4.858	5.033	3.293	5.330
s		1.288	1.289	1.307	1.34

^aAsimow *et al.* [2008], Chase[1998]; ^b Mitchell and Nellis [1981b]; ^c Marsh [1980]

Table 3. Equation of State fits for molten CaFeSi₂O₆

	Units	SWEOS	BM3	BM4	Source
T_0	K	1673	1673	1673	
ρ_0	kg m ⁻³	2913 ± 11	2913 ± 11	2913 ± 11	G13a
C_0	m s ⁻¹	2613 ± 16	-	-	G13a
s		1.54 ± 0.01	-	-	fitted
γ_0		0.300	0.300	0.300	derived
q		-	-1.93 ± 0.41	0.47 ± 9.86	fitted
K_{S0}	GPa	19.89 ± 0.24	19.89 ± 0.24	19.89 ± 0.24	derived
K_S'		5.16 ± 0.04	6.22 ± 0.55*	3.15 ± 1.61*	derived/*fitted
K_S''	GPa ⁻¹	-	-	0.78 ± 0.60	fitted
χ^2		-	4.20	3.26	
K_T	GPa	19.32	19.32	19.32	derived
α	K ⁻¹	5.91-05	5.91-05	5.91-05	G13a
C_p	J kg ⁻¹	1345.48	1345.48	1345.48	LN92
C_v	J kg ⁻¹	1306.70	1306.70	1306.70	derived

SWEOS= shock wave equation of state. 3BM/MG = 3rd-order Birch-Murnaghan/Mie-Grüneisen EOS 4BM/MG = 4th-order Birch-Murnaghan/Mie-Grüneisen EOS Sources: *Fitted* indicates an adjustable parameter, L97 is Lange [1997], LN92 is Lange and Navrotsky [1992], G13a is Guo *et al.* [2013a]

Table 4. Equation of State fits for molten An-Hd

	Units	SWEOS	BM3	BM4	Source
T_o	K	1673	1673	1673	
ρ_o	kg m ⁻³	2772 ± 3	2772 ± 3	2772 ± 3	G13b
C_o	m s ⁻¹	2772 ± 11	-	-	G13b
s		1.39 ± 0.03	-	-	fitted
γ_o		0.310	0.310	0.310	derived
q		-	-0.18 ± 2.25	1.67 ± 2.79	fitted
K_{So}	GPa	21.30 ± 0.25	21.30 ± 0.25	21.30 ± 0.25	derived
K'_S		4.54 ± 0.12	$5.49 \pm 0.65^*$	$5.10 \pm 1.61^*$	derived/*fitted
K''_S	GPa ⁻¹	-	-	-0.17 ± 0.82	fitted
χ^2		-	6.37	12.53	
K_T	GPa	20.68	20.68	20.68	derived
α	K ⁻¹	5.81-05	5.81-05	5.81-05	G13b
C_p	J kg ⁻¹	1442.34	1442.34	1442.34	LN92
C_v	J kg ⁻¹	1400.19	1400.19	1400.19	derived

Sources: *Fitted* indicates an adjustable parameter, L97 is *Lange* [1997], LN92 is *Lange and Navrotsky* [1992], G13b is *Guo et al.* [2013b]

Table 5. Equation of State fits for molten An-Hd-Di

	Units	SWEOS	BM3	Source
T_o	K	1673	1673	
ρ_o	kg m ⁻³	2722 ± 2	2722 ± 2	G13b
C_o	m s ⁻¹	2846 ± 14	-	G13b
s		1.54 ± 0.02	-	fitted
γ_o		0.361	0.361	derived
q		-	-1.14 ± 0.79	fitted
K_{So}	GPa	22.03 ± 0.25	22.03 ± 0.25	derived
K'_S		5.15 ± 0.06	$6.20 \pm 0.17^*$	derived/*fitted
K''_S	Gpa ⁻¹	-	-	fitted
χ^2		-	0.39	
K_T	Gpa	21.19	21.19	derived
α	K ⁻¹	6.64-05	6.64-05	G13b
C_p	J kg ⁻¹	1491.96	1491.96	LN92
C_v	J kg ⁻¹	1434.47	1434.47	derived

For abbreviations see Tables 3 and 4

FIGURE CAPTIONS

Figure 1. Preheated (1673K) $\text{CaFeSi}_2\text{O}_6$ liquid Hugoniot in shock velocity (U_s)-particle velocity (u_p) space. The dotted line represents the unconstrained linear Hugoniot, G13a is *Guo et al.*[2013a submitted].

Figure 2. $\text{CaFeSi}_2\text{O}_6$ liquid Hugoniot plotted in pressure-density space with thermal EOS fits. Data symbols are the same as Figure 1. Abbreviations: SWEOS- shock wave equation of state; BM/MG – Birch-Murnaghan/Mie-Grüneisen EOS, G13a is *Guo et al.* [2013a submitted].

Figure 3. Preheated (1673K) An-Hd liquid Hugoniot in shock velocity (U_s)-particle velocity (u_p) space. The dotted line represents the unconstrained linear Hugoniot. G13b is *Guo et al.*[2013b in prep]

Figure 4. An-Hd liquid Hugoniot plotted in pressure-density space with thermal EOS fits. Data symbols are the same as Figure 3. The abbreviations are the same as Figure 2 and 3.

Figure 5. Preheated (1673K) An-Hd-Di liquid Hugoniot in shock velocity (U_s)-particle velocity (u_p) space. The dotted line represents the unconstrained linear Hugoniot G13b is *Guo et al.* [2013b in prep]

Figure 6. An-Hd-Di liquid Hugoniot plotted in pressure-density space with thermal EOS fits. Data symbols are the same as Figure 6. The abbreviations are the same as Figure 2 and 5.

Figure 7. A comparison of the mixing model ($Hd_{modelFa} = Di + 0.5Fa - 0.5Fo$) and the Hd isentrope at various potential temperatures (T_P).

Figure 8. a) A comparison of the mixing model using a Fa end member ($AnHd_{modelFa} = 0.5An + 0.25Fa - 0.25Fo + 0.5Di$) and the An-Hd isentrope at various potential temperatures (T_P) and b) its % errors in temperature and density in estimating the An-Hd isentrope. c) A comparison of the mixing model using a Hd end member ($AnHd_{modelHd} = 0.5An + 0.5Hd$) and the An-Hd isentrope at various potential temperatures (T_P) d) its % errors in temperature and density in estimating the An-Hd isentrope. Errors shown in text are the maximum at 120 GPa.

Figure 9. a) A comparison of the mixing model using a Fa end member ($AnHdDi_{modelFa} = 0.33An + 0.667Di + 0.167Fa - 0.167Fo$) and the An-Hd-Di isentrope at various potential temperatures (T_P) and b) its % errors in temperature and density in estimating the An-Hd-Di isentrope. c) A comparison of the mixing model using a Hd end member ($AnHdDi_{modelHd} = 0.33An + 0.33Hd + 0.33Di$) and the An-Hd-Di isentrope at various potential temperatures (T_P) d) its % errors in temperature and density in estimating the An-Hd-Di isentrope. Errors shown in text are the maximum at 120 GPa.

Figure 10. a) A comparison of the two mixing models for chondrite liquid ($Ch_{modelFa} = 0.62En + 0.24Fo + 0.08Fa + 0.04An + 0.02Di$ and $Ch_{modelAnHdDi} = 0.62En + 0.32Fo + 0.45AnHdDi -$

0.12An -0.28Di) and b) their % error difference in temperature and density. Errors shown in text are the maximum at 120 GPa.

FIGURES

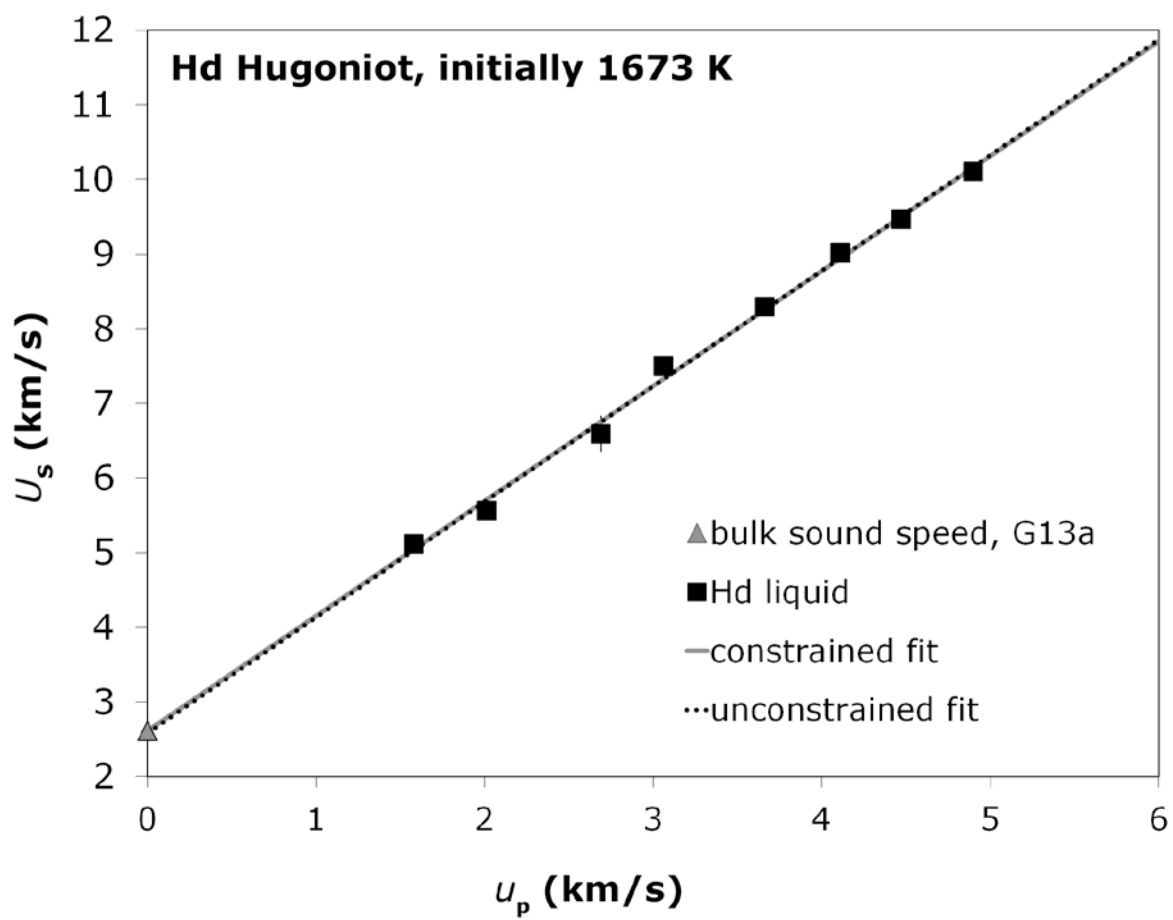


Figure 1

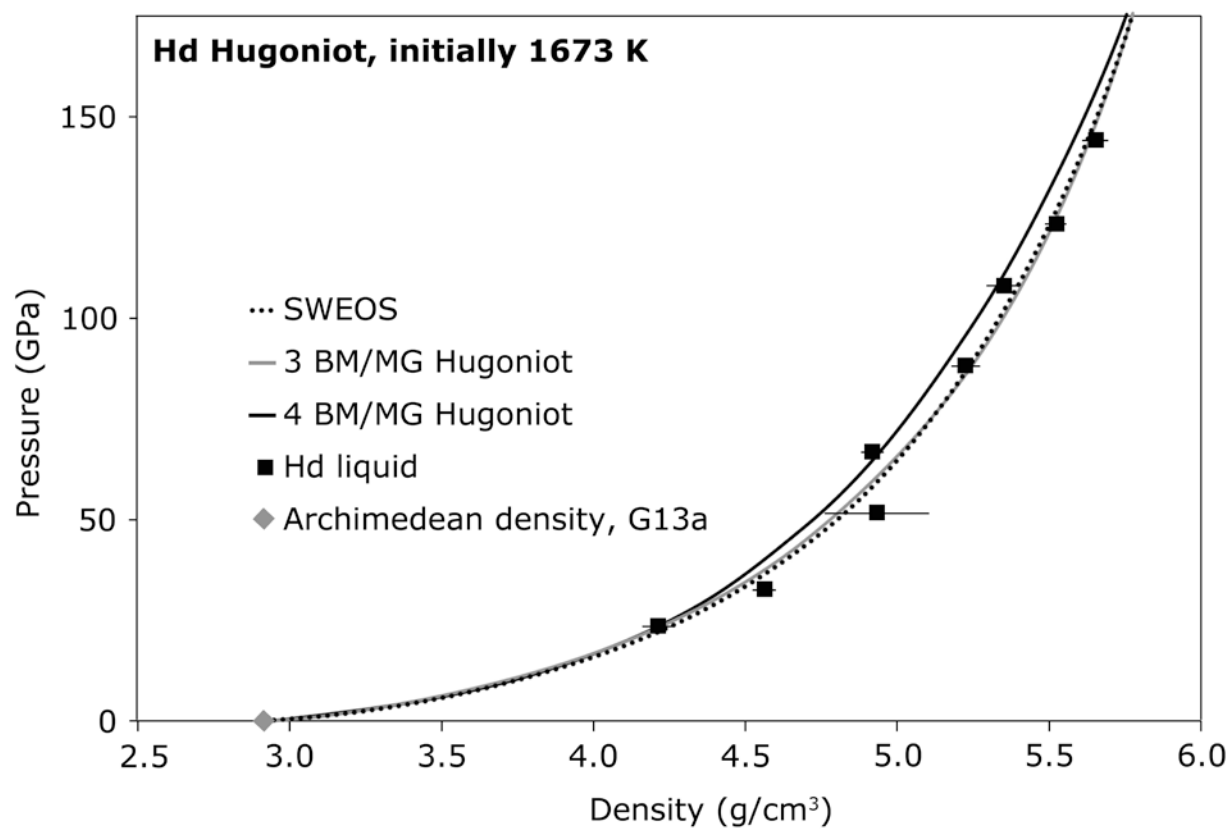


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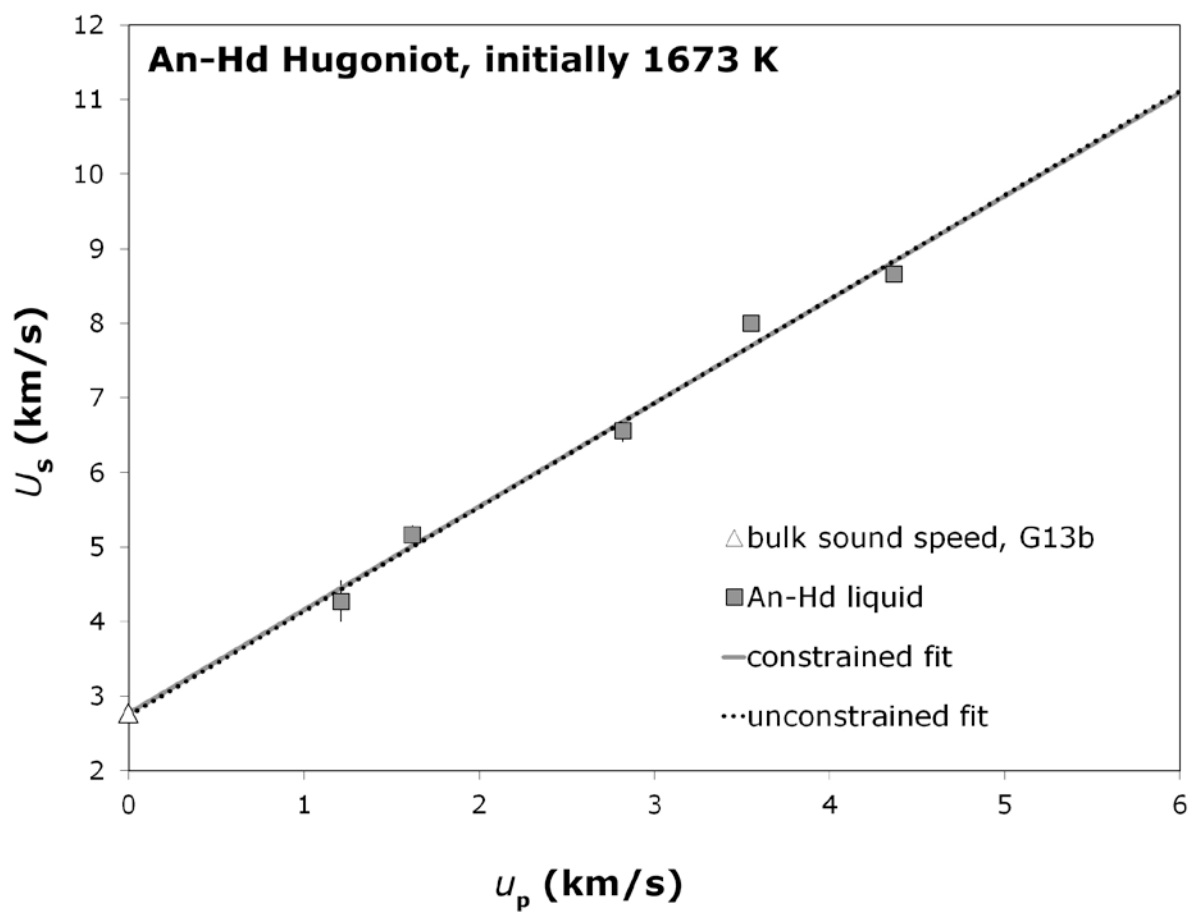


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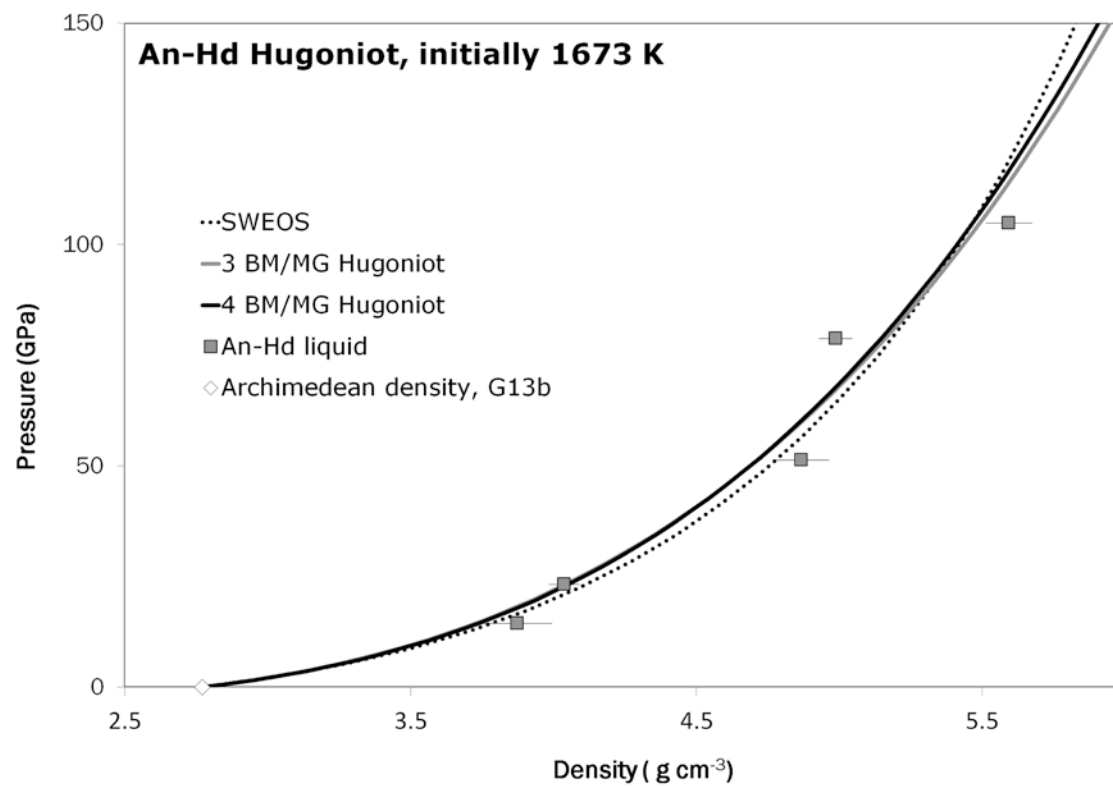


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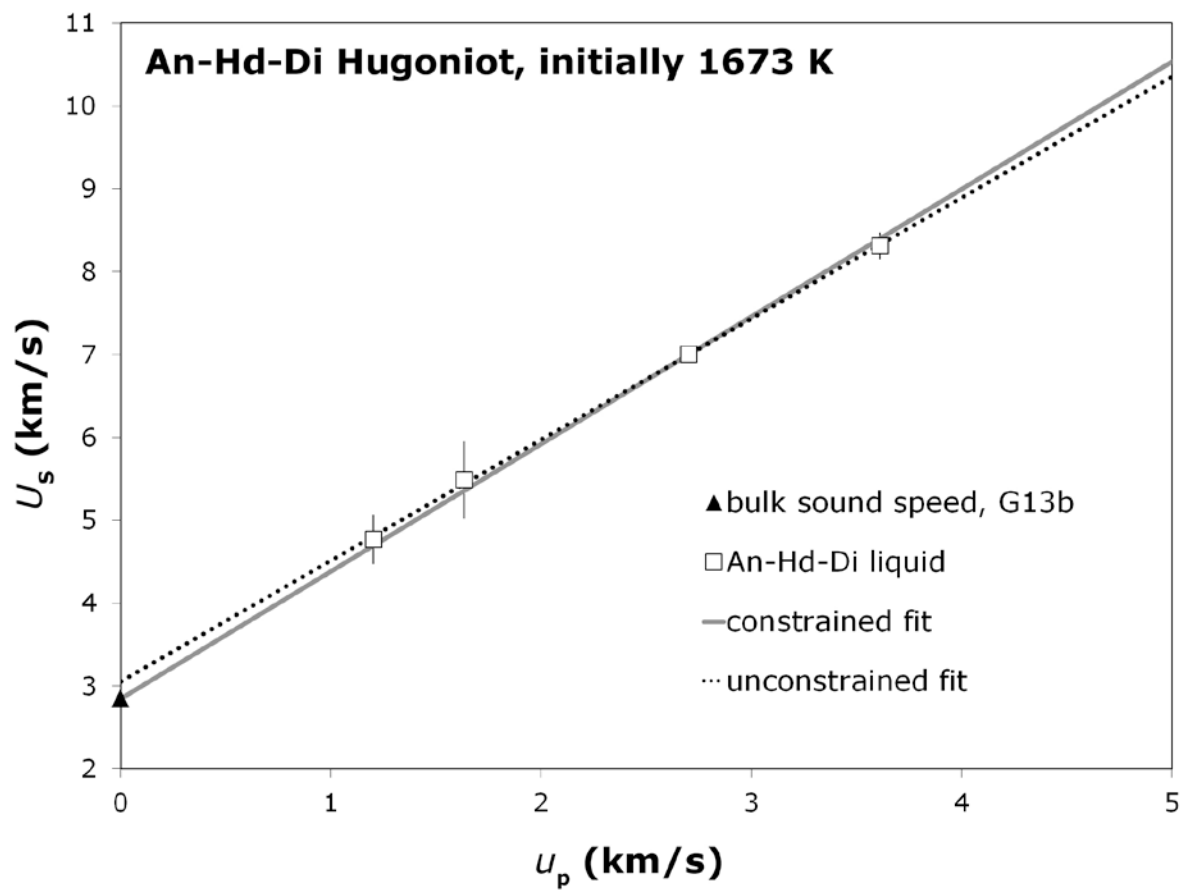


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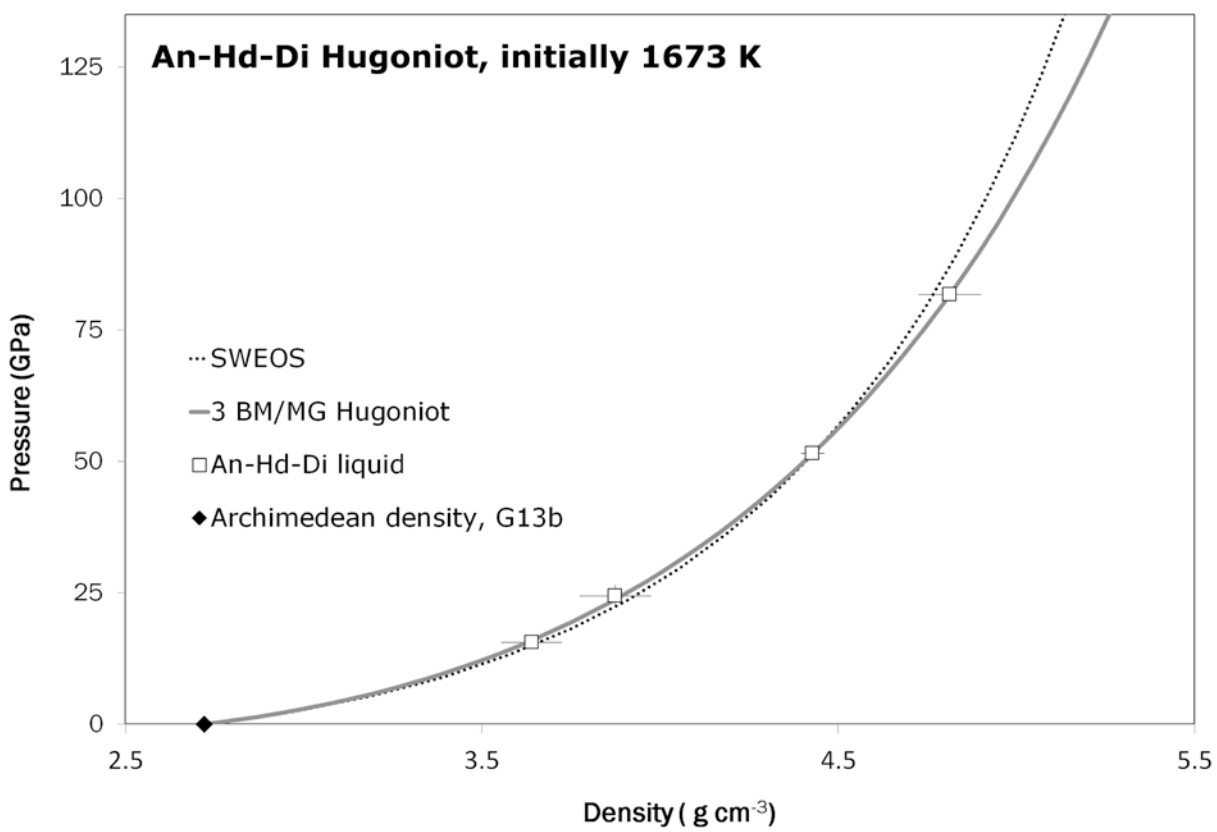


Figure 6

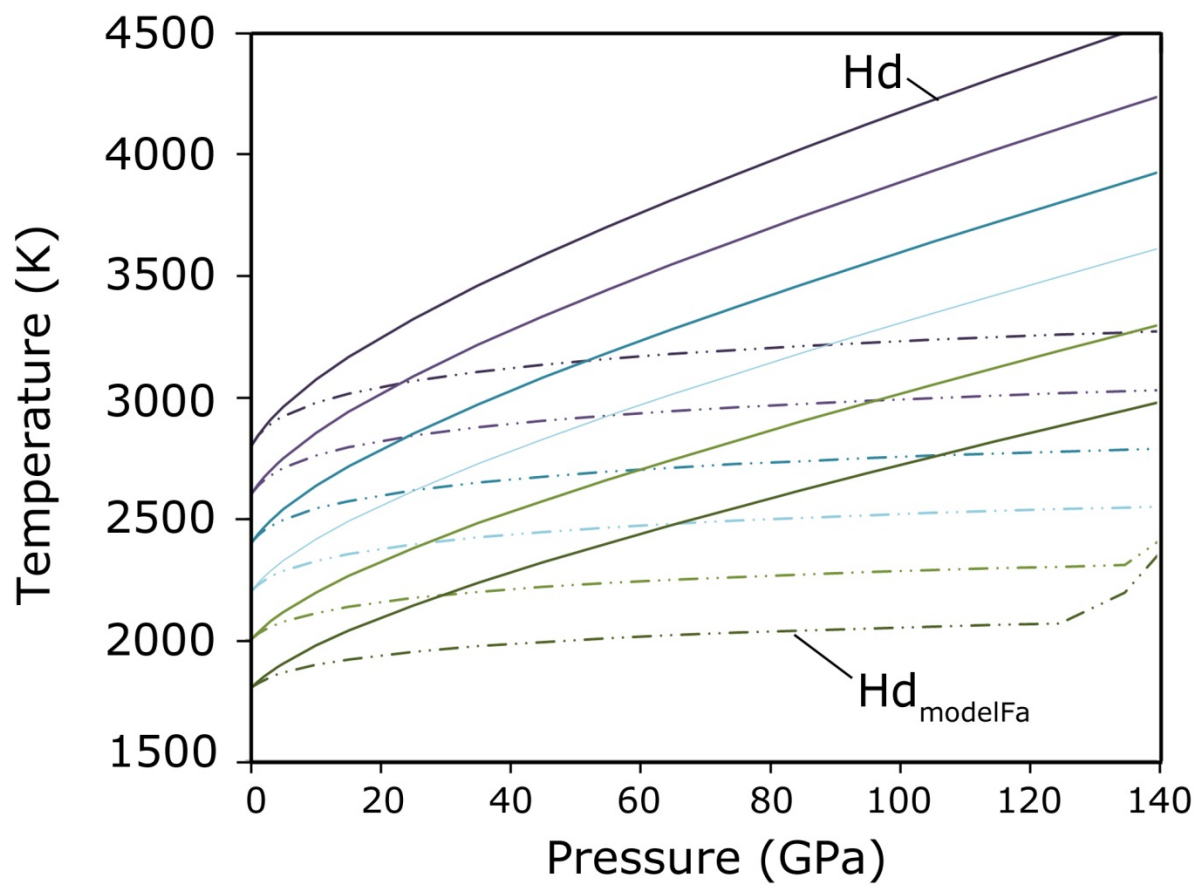


Figure 7

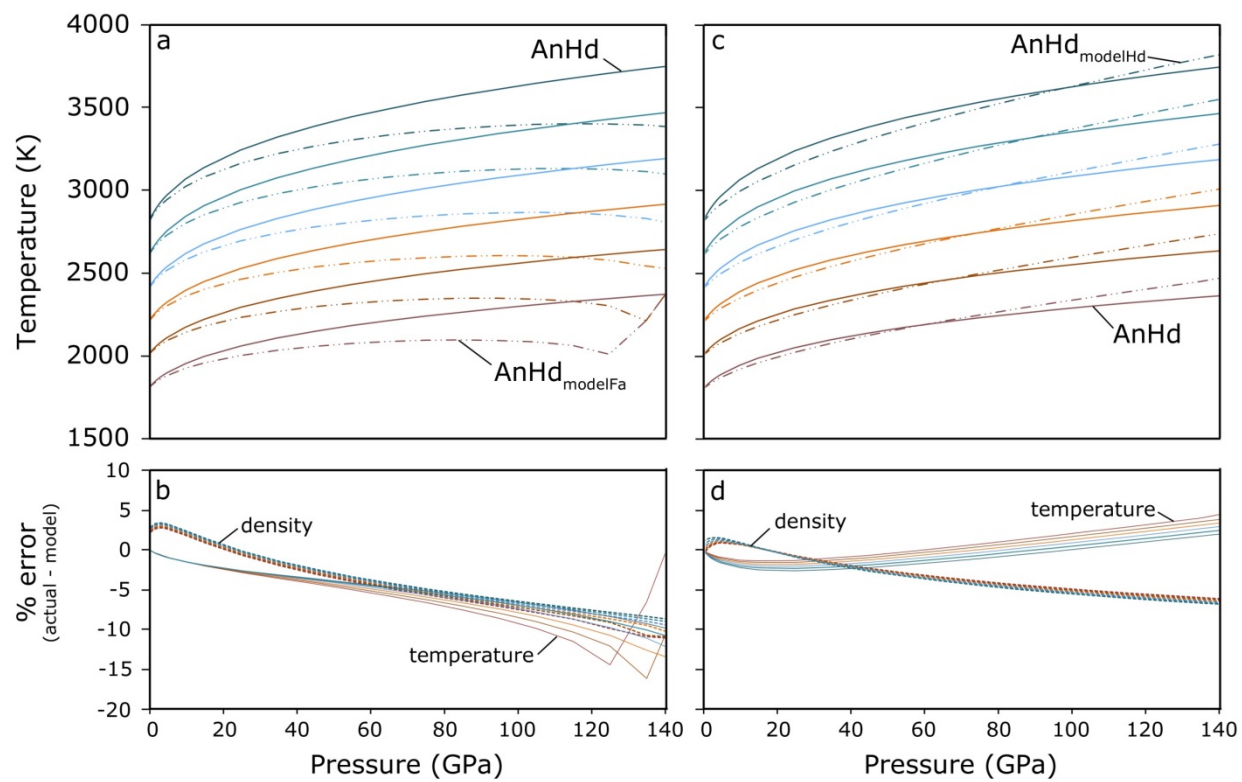


Figure 8

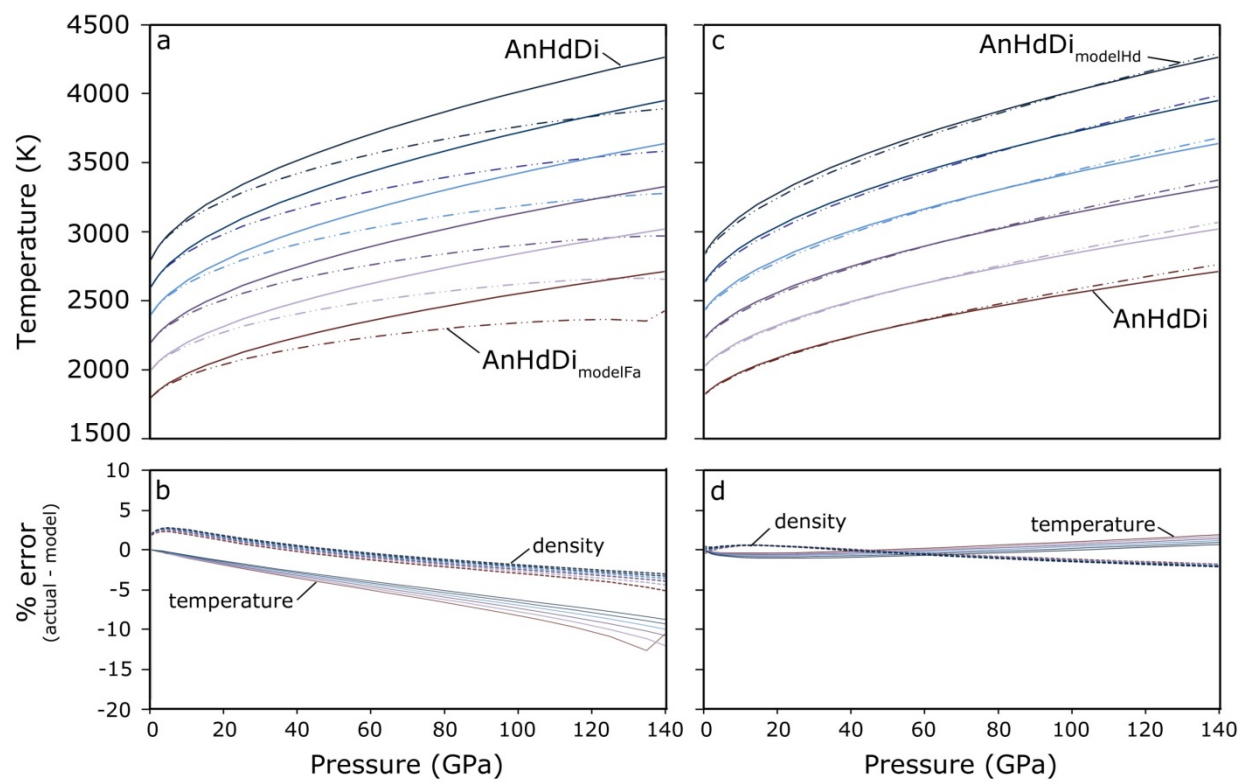


Figure 9

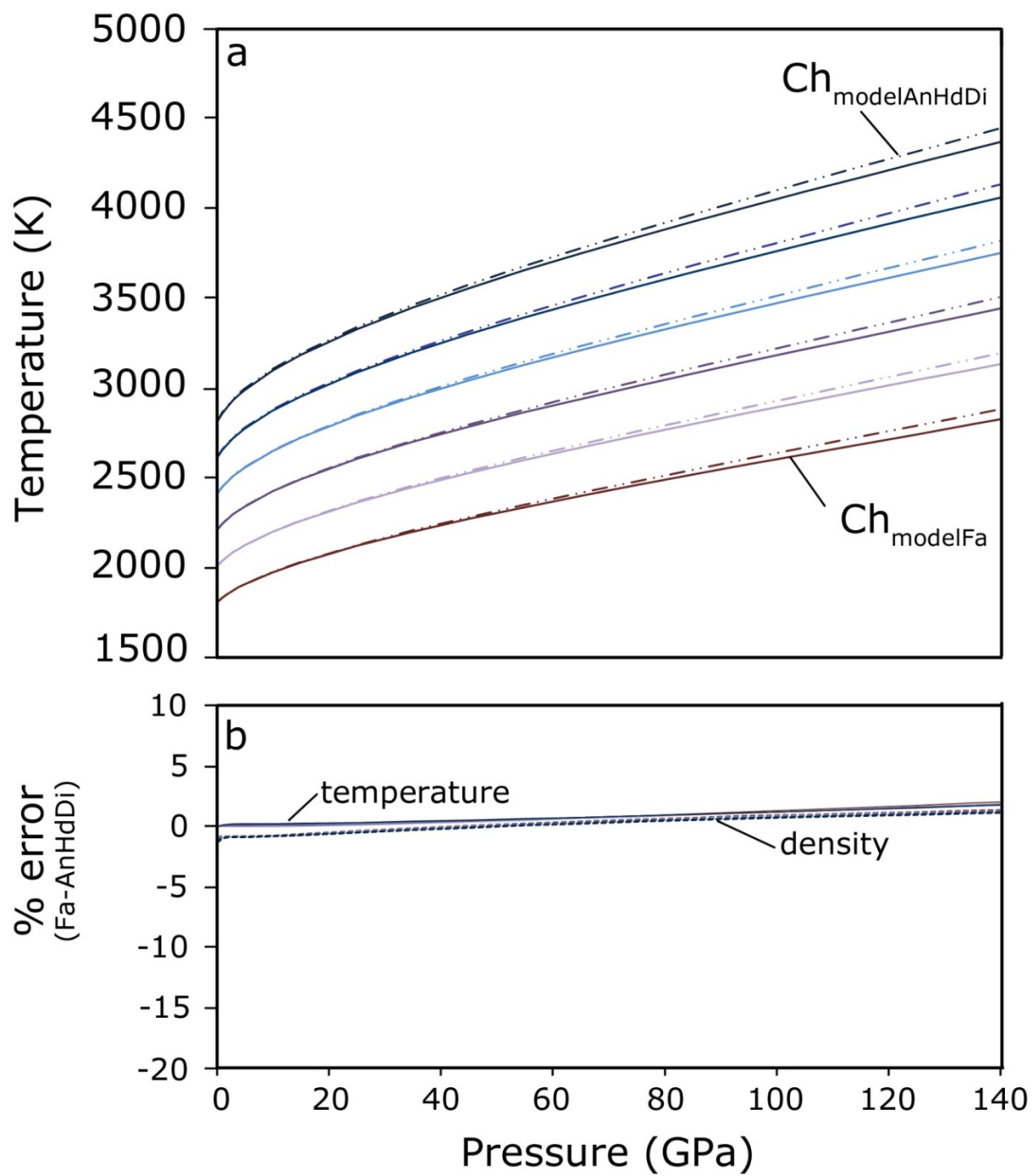


Figure 10