

# 3-D Photopolymer Disks for Correlation and Data Storage, and Cross-talk in Volume Holographic Memories

Thesis by  
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## 0.2 ABSTRACT

This thesis considers the optical storage and processing of data using volume holography. First, cross-talk noise due to geometrical considerations is calculated for volume holographic memories. Cross-talk is due to energy diffracted from non-Bragg matched gratings. The SNR (signal-to-noise-ratio) of holographic memories due to cross-talk noise is calculated for Fourier transform holograms stored by angle, wavelength, phase-coded, and rotational multiplexing methods. Considerations include page size, geometry, angular bandwidth of the optical system, wavelength of the light used, material size, spatial light modulator (SLM) contrast, and the phase of the image plane. The SNR for angle multiplexed image plane holograms is also calculated and compared to the results for Fourier transformed angle multiplexed holograms. A comparison of the various multiplexing methods based on cross-talk is presented, and then the effect of geometry and material dynamic range is included to determine when cross-talk will be the dominant noise source. The use of photopolymers as a holographic element is then presented. The recording characteristics of the DuPont photopolymer are described and a method of multiplexing multiple holograms in the photopolymer is given. A new method for multiplexing holograms (called peristrophic multiplexing) is described. This method significantly increases the storage capacity of thin films. After this, a 3-D disk-based correlator and storage device using the photopolymer is described and demonstrated. In this device, holograms are multiplexed at a given spot and then disk rotation/head motion are used to access multiple spots on the disk. Theoretical correlation speed, read-out rates, and the storage capacity of the 3-D disk as limited by geometry and laser power are given.

Advisor: Professor Demetri Psaltis.

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# Chapter 1

## Introduction

### 1.1 Holography

In 1948, Dennis Gabor proposed [1] a wavefront reconstruction method of imaging, and holography was born. When a coherent reference beam is interfered with an object beam, the intensity of the interference pattern contains both the amplitude and phase information of the beams. Since materials such as film record the intensity of the light, this method allows for the total information of the object beam to be recorded in the film. Gabor called the interference pattern a hologram, meaning “total recording.”

Through out the fifties, Gabor [2, 3] and a few others [4, 5, 6] expanded the principles laid out in his original paper. During this time, holograms were recorded on-axis. This meant that both real and virtual images were reconstructed at the same time and in the same direction, and therefore the quality of the reconstruction was seriously limited. It was not until Leith and Upatnieks [7] developed off-axis holography that the quality of holographic reconstructions significantly improved. In off-axis holography, the reference beam is incident on the film at an angle while the object is incident to the medium at a different angle. This allows the virtual and real images, as well as the undiffracted light, to be separated. With this breakthrough in recording technique came new applications besides the original application of imaging.

For example, 3-D image reconstruction [8], information storage [9] – [13], filtering [14, 15], aberration correction in imaging systems [16, 17], and synthetic-aperture radar processing [18, 19] were all proposed and to some extent demonstrated during the 1960's and early 1970's.

During the 1960's, the material used in most experimental work was photographic film. A lot of work was done on evaluating the holographic quality [20, 21, 22] and the linearity of these films [23, 24]. The film was typically a thin, two dimensional (2-D) layer of recording material. Thin holograms lack sensitivity to changes in angle of the reconstruction beam. This property allows for the construction of shift invariant filters for correlation, but also restricts the capacity to approximately one hologram per spatial location on the film.

## 1.2 Volume Holography

The limited storage density of 2-D holograms can be overcome by using volume holography. In volume holography, the recording medium is a three dimensional (3-D) solid. Due to the fact that the interference pattern is recorded throughout the material, thick holograms are sensitive to changes in angle or wavelength of the reconstructing beam. This effect was first noticed in X-ray diffraction from crystals by Bragg in 1942 [25]. In 1969, Kogelnik [26] analyzed the diffraction of light from a hologram written in a volume by interfering two plane waves. Kogelnik's result showed how Bragg selectivity applied to thick, optical, volume holograms. Later analysis [27, 28, 29, 30] discussed Bragg selectivity and diffraction from materials whose thickness is between the thick and thin holographic regimes. Using the Bragg selectivity inherent to volume holography, multiple holograms can be stored in the same volume by changing either the angle or wavelength of the recording beams. Since Bragg selectivity allows for many holograms to be superimposed inside the

same volume, it significantly increases the storage density of the medium.

Historically, the most important type of material for volume holography is photorefractive crystals. In photorefractives, light excites electrons out of traps in the material and into the conduction band. Once in the conduction band, the electrons either diffuse out of the illuminated regions or are forced out of these regions by built in or applied electric fields. Once in the dark regions, they are retrapped. Therefore, in the illuminated regions there are uncompensated positive charges while in the dark regions there are uncompensated negative (trapped electrons) charges. This separation of charge has an associated electric field with it, which in turn generates a change in index of refraction of the material by the electro-optic effect. Thus the sinusoidal interference pattern due to two coherent plane waves is recorded by a sinusoidal perturbation in the index of refraction of the material called a phase grating. Phase gratings can have high diffraction efficiency ( up to 100 percent). The theoretical details of this recording process in photorefractives were worked out by several authors [31]–[36] including Kukhtarev [37].

Holographic storage in photorefractives was demonstrated in the late 1960’s and early 1970’s by several authors [38]–[41]. Particularly impressive was the work of Amodei, Philips, and Staebler in iron doped  $\text{LiNbO}_3$ . They demonstrated storage of several hundred holograms in  $\text{LiNbO}_3\text{:Fe}$  [42]. Recently, several thousand holograms have been stored in iron doped  $\text{LiNbO}_3$  [43, 44] and recalled with good signal-to-noise ratio.

Although photorefractives are inherently dynamic materials, there are a few methods of “fixing” the holograms – making the holograms permanent. Thermal fixing (heating the crystals to several hundred degrees Celsius) has been demonstrated in several photorefractives [45]–[49]. In thermal fixing, the electronic gratings established by the optical fields are compensated by ionic gratings. Since the ions are

only mobile at very high temperatures, these ionic gratings are effectively permanent at room temperature. In Strontium Barium Niobate (SBN), electrical fixing [51]–[56] is accomplished by electric domain reversal in the ferroelectric crystal. In this method, the electronic gratings set up by the optical field are partially compensated by electric domain reversal after a depoling electric field is applied to the crystal. Another method of preventing hologram decay is the continual refreshing of the holograms (copying), and this method has been experimentally demonstrated [57, 58, 59]. Reading out the holograms with a wavelength that is non-destructive has also been shown [60, 61, 62, 63, 64]. Thus, in photorefractives, volume holographic storage has been extensively studied.

Other, less common materials for volume holography are organic polymers. Recently, polymers have been used in several holographic and non-holographic storage techniques. Examples of non-holographic storage include Rentzepis’s 3-D storage [65, 66, 67], electron trapping materials [68, 69], and for spectral hole burning applications [70, 71, 72, 73]. In addition, there have been several recent ideas for dynamic holographic photopolymer materials. These include liquid crystal doped polymers [74]–[79], and the development of photorefractive polymers [80]–[99]. In photorefractive polymers, the recording mechanism is similar to the mechanism in inorganic crystals except that the mobile charges are holes, and strong electric fields have to be applied to enable charge migration and to enhance the electro-optic effect. Fixed gratings have yet to be demonstrated in these photorefractive polymers. With this new emphasis on polymers in recent years, there has been renewed interest in the original type of photopolymer. In this type of photopolymer light induces polymerization of monomers inside a polymer host.

In the early 1970’s, while photorefractives were initially being developed, photopolymers were also pursued for the purpose of holographic storage [100]–[110].

These photopolymers consist of monomers and a light sensitive dye inside a polymer matrix. The light excites the dye and initiates a free radical reaction. This locally polymerizes the monomers in the illuminated regions and sets up a gradient in monomer concentration. The monomers then diffuse into these regions and are polymerized, which increases the density of the film in the illuminated regions. This density change directly relates to a change in the index of refraction. Thus, like photorefractives, these polymers record phase gratings. Unlike photorefractives, these films inherently record permanent gratings since the polymerization reaction is strongly irreversible. After the holograms are stored, the film is uniformly exposed to ensure complete polymerization of the film. In 1976, Bartolini et al. [111] angle multiplexed 550 holograms in one centimeter thick photopolymer film. Using similar material, DuPont had a holographic photopolymer that was commercially available in the 1970's. Recently, DuPont, Polaroid, and others have photopolymer films [112]–[117] that have good sensitivity, resolution, and relatively large index changes. These films have been used for filters [118, 119, 120], holographic storage [121, 122], and the verification of drivers licenses and credit cards (although, the holograms on these cards are mechanically, not holographically reproduced). In addition, the theory of polymerization and monomer diffusion in photopolymer films has been investigated [123, 124].

### 1.3 Volume Holographic Storage – Then and Now

The advantages of holographic storage are high storage density and the parallel access capability. These features were recognized in the 1960's and 1970's [9, 39, 111] and serious efforts were made to take advantage of the technology. Unfortunately, this work did not lead to a commercial product. Lately, holographic storage has again received attention as a possible technology for high density data storage. A valid



question is, what has changed in the last 20 years to make holographic storage a commercially viable technology ?

Some of the more important factors that have changed in the last 20 years are listed below.

- **Need:** The emergence of applications that require storage on the order of terabytes is a very important factor. As computers transform from stand alone units to large networks with graphical/oral interfaces, and the emergence of multimedia (the fusion of text, video, and audio information), the storage requirements and data readout rates for storage devices have dramatically increased. In addition to these memory applications, new applications that require high storage density and parallel access have arisen such as neural networks, image processing, and large data bases. Clearly, the need for data storage on the terabyte level has changed significantly since the 1970's.
- **Evolution of CCD and SLM technology:** In the 1970's, presenting digital data and quickly reading out digital data from an optical system was impeded by the lack of input and output devices. The choice for input devices was limited to 1-D modulators such as acoustic-optical modulators or electro-optical modulators. With the advent of projection televisions, spatial light modulators (SLMs) based on liquid crystal technology have become commercially available. These SLMs provide an easy way to present digital data in a 2-D form to an optical system [125]–[137]. SLMs can have large transfer rates and some types support gray level encoding of the data. In addition, the dramatic changes in VLSI (very large scale integrated) technology, have made available large, fast CCD (charge coupled device) arrays [138, 139, 140, 141] for detecting and reading out the data in parallel. Thus, using both SLMs and CCD arrays, large amounts of digital data can be stored and accessed in parallel.

- **Light Sources:** Commercially viable optical memories need reasonably compact, reliable, and efficient light sources. Since most holographic materials are sensitive in the visible, these light sources have to operate in the visible to near infrared region of the spectrum. Using these criteria, there was not an acceptable light source available in the 1970's. Recently, tremendous advancements have been made in solid state lasers [142, 143, 144, 145] and even short wavelength (blue) semiconductor lasers [146, 147, 148, 149]. Currently, solid state lasers are commercially available at 532 nanometers with the high power necessary for holographic storage.
- **Holographic Storage Techniques:** Very significant advancement has occurred in holographic storage techniques. These include understanding the scaling of diffraction efficiency with the number of holograms stored , recording schedules for storing large numbers of holograms with equal diffraction efficiency, understanding cross-talk, increased stability of recording setup, optimization and characterization of storage materials, and most importantly computer control of the storage process. The advancement in computer control was an enabling technology that made the repeated recording of a large number of holograms technically feasible.
- **Optical Components:** The manufacture and design of optical systems has made dramatic improvements in the last 20 years. The development of diffractive optics to make diffusers and lenses has significantly increased the flexibility of holographic storage system design. Computers themselves have changed how optical systems can be designed and manufactured to decrease both the cost and the volume of the systems.
- **Storage Materials:** The availability of storage materials has been signifi-

cantly improved, and now several materials for volume holographic storage are commercially available in quantity. Also, procedures for the optimization and characterization of photorefractive crystals have been developed. Materials are still critical for improved recording rates and for the future extension of holographic memories from the current write once and read many times (WORM) or read only memories (ROM) memory architectures to dynamic read/write memories.

Clearly, there have been many important advances since the 1970's. Essentially, the holographic storage effort of the 1970's did not have the supporting technology needed to have a chance of making a commercially viable memory. Whether or not holographic storage has a future now is still an open question. Currently, there is no technical impediment to making holographic memory systems, but cost and market niche are certainly important factors to consider when faced with competition from mature storage technologies such as magnetic, semiconductor, and conventional optical memories.

## 1.4 Outline of Thesis

The major sections of this thesis involve cross-talk calculations, photopolymer storage, and a 3-D holographic disk for data storage and correlation. Cross-talk is presented first, since it is inherent in any volume holographic memory or processor. Next DuPont's photopolymer is investigated as an example of a particular storage medium. Then the holographic photopolymer disk is presented as a new holographic storage and processing architecture.

Chapter 2 details the cross-talk calculations. Cross-talk due to purely geometrical factors is calculated for Fourier plane holograms stored by angle, wavelength, phase codes, and rotation multiplexing. This is extended to image plane angle multiplexed

holograms, and the effect of finite image contrast is also analyzed.

Chapter 3 compares the various multiplexing methods on the basis of cross-talk. The best geometrical arrangements to minimize the cross-talk noise are discussed. Then, the circumstances where cross-talk is the dominant noise source are determined.

Holographic recording characteristics of DuPont's HRF-150 photopolymer are investigated in Chapter 4. The modulation transfer function is determined and explained. The effect of intensity, fringe visibility, grating period, and grating tilt on diffraction efficiency is determined. The lateral spread of the polymerizing reaction is also investigated. A method for recording multiple holograms in DuPont photopolymer films is then explained. A new method for multiplexing holograms (peristrophic multiplexing) is then explained and demonstrated. This new technique significantly increases the storage capacity of thin films.

The details of a 3-D holographic disk-based correlator and storage device is introduced in Chapter 5. The experimental correlator generates full 2-D correlations using disk motion, and was experimentally validated using photopolymer film. The possible use of this architecture as a data storage system is examined. The geometrically limited storage capacity of such a system is determined as well as the power limited read-out rate.

## Chapter 2

# Cross-talk in Volume Holographic Memories

### 2.1 Introduction

Volume holography achieves high storage density by superimposing many holograms inside the same volume, but as the old adage says, “you don’t get something for (sic) nothing.” Part of the price for this increase in storage capacity is the amount of noise that the other non-Bragg matched holograms add to the desired reconstruction. Figure 2.1 shows a volume with two gratings inside of it. The reference beam, incident from the left, has the correct wavelength and angle to be Bragg matched to one of the gratings. This Bragg matched grating diffracts a strong signal beam, but the other grating, even though it is not Bragg matched, still diffracts some light. Therefore, in a holographic memory, the non-Bragg matched holograms in the volume are also reconstructed with lower efficiency and with distortions due to Bragg mismatch. This diffraction of light from non-Bragg matched gratings is called cross-talk noise. Cross-talk noise places an upper limit on the signal-to noise ratio (SNR) of the system, and therefore the storage capacity of the memory.

The original efforts to calculate cross-talk [150, 151, 152, 153, 154] were hindered by the lack of understanding about how to space the holograms in angle. Bragg selectivity applies to a change of angle in only one direction. The superposition of

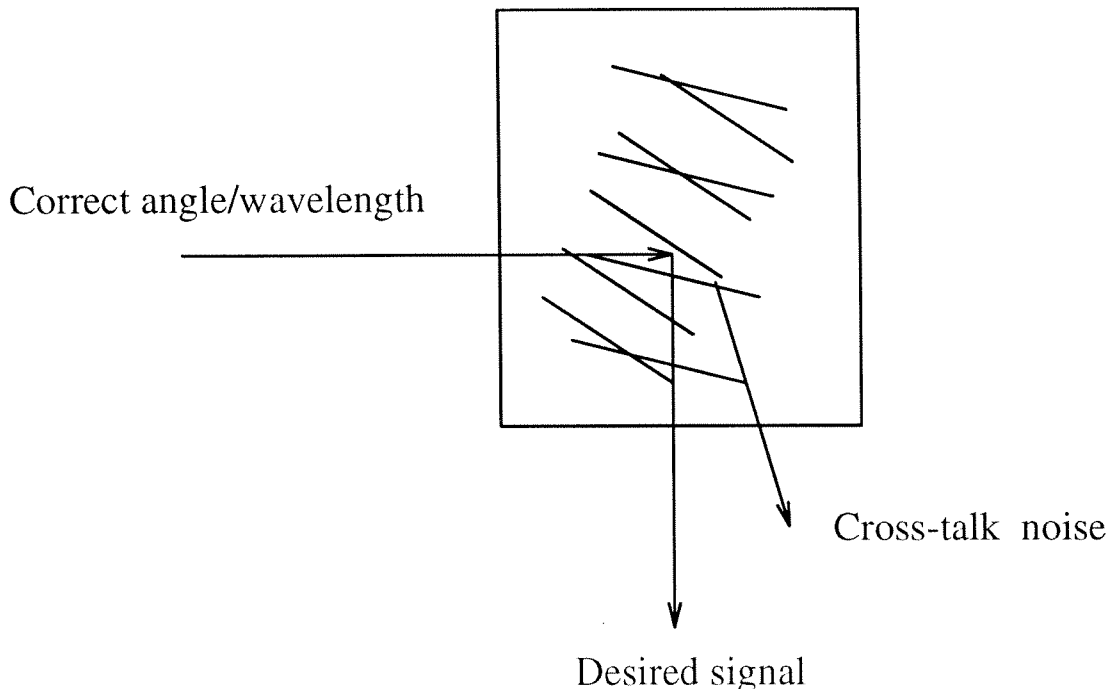


Figure 2.1: Cross-talk from non-Bragg matched gratings.

holograms can only (ignoring fractal sampling grids for the moment) be done by changing the reference beam angle in the plane of intersection. In the other direction, the gratings are very insensitive to changes in angle. This effect is called grating degeneracy. Grating degeneracy means that certain sets of angles in the reference and object beams are connected by a single grating. Most of the early calculations of cross-talk were dominated by this “first order” cross-talk noise. Schemes for eliminating cross-talk due to grating degeneracy, called fractal sampling grids, were derived and demonstrated by Psaltis et al. [155] and others [156, 157].

Other work focused on the effects of two-wave mixing, and using coupled-wave analysis to determine cross-talk for a small number of gratings [158] – [167]. Recently, it was shown that for low (less than 1%) throughput holograms, the cross-talk due to angular sidelobe overlap is the dominant cross-talk effect [168, 169]. Since for holographic memories, a large number of holograms means low diffraction efficiency, this geometrical cross-talk is the relevant cross-talk source.

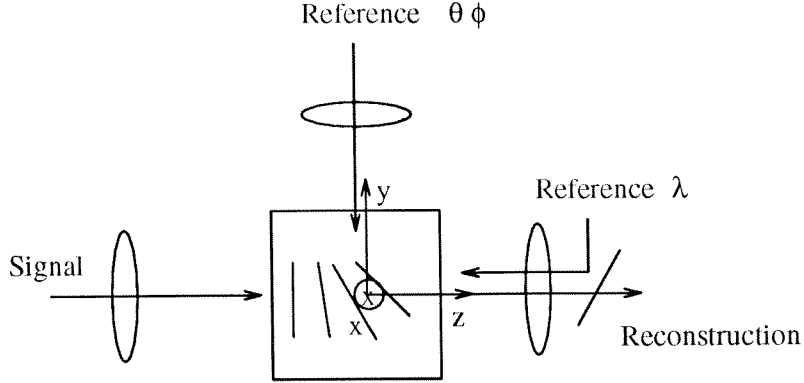


Figure 2.2: Geometry for wavelength, phase coded, and angle multiplexing.

The information storage capacity of volume holographic memories is limited by geometry and material dynamic range. The SNR is a convenient metric that we use to assess the storage capability of holographic memories. In this chapter we calculate the SNR due to cross-talk between holograms purely due to geometrical constraints. The derivation of the cross-talk for the various multiplexing methods we present here is analogous to the method used to by Gu et al. [170] for angle multiplexing and Curtis et al. [171, 172, 173, 174] for the other multiplexing methods.

### 2.1.1 Multiplexing Methods

Multiple holograms can be recorded in a material by changing either the reference beam angle [42, 43] (angle multiplexing), the recording wavelength [176, 175] (wavelength multiplexing), or by phase coding the reference beam [177, 178, 179]. These volume holographic memories have potentially high storage capacity and fast – page at a time – readout. The geometries used for recording are shown in Figure 2.2. In all geometries shown in Figure 2.2, the input data is presented on a spatial light modulator (SLM) from the left and detected on the right. For angle multiplexing, the reference is a plane wave input at a given angle (shown as  $90^\circ$  in Figure 2.2). Multiple holograms are recorded by changing the angle of the reference beam. Wavelength multiplexing is done in reflection geometry for maximum capacity. Multiplexing is

accomplished by changing the wavelength of the recording beams. In phase-code multiplexing, the reference beam consists of a set of plane waves each with an individual phase. The set of phases is called a phase-code. Multiple holograms are recorded by changing the phase-code of the reference beam.

## 2.2 Wavelength Multiplexing

### 2.2.1 Introduction

In the past, most of the work has concentrated on angle multiplexing, primarily because convenient frequency-tunable light sources have not been available. With the advent of semiconductor and solid-state tunable lasers, there has been renewed interest in wavelength multiplexing. Rakuljic et al. [175] have argued that wavelength multiplexed memories have low cross-talk noise. The section contains two main results. First, we show that for the optimum counter-propagating reflection geometry, the SNR for wavelength multiplexed holograms saturates for a large number of holograms [181]. Second, we derive a closed-form expression for the SNR for a large number of holograms (more than several hundred) in a general reflection geometry which shows that, for a given SNR, cross-talk limits the size of the stored images.

### 2.2.2 Theory for Wavelength Multiplexing

Fourier-transform holograms are wavelength multiplexed in a volume holographic medium using the setup shown in Figure 2.3. A plane wave  $R_m$  is generated at a wavelength  $\lambda_m$  and at an angle  $\theta$  with respect to the  $z$  axis by collimating light from a point source. This interferes inside the holographic medium with  $S_m$  – the Fourier-transform of the  $m$ th object image also of wavelength  $\lambda_m$ . By labeling the holograms  $m = -M, -(M - 1), \dots, 0, \dots, M$ , the presence of these  $N = 2M + 1$  holograms modulates the permittivity of the material such that the change in the permittivity



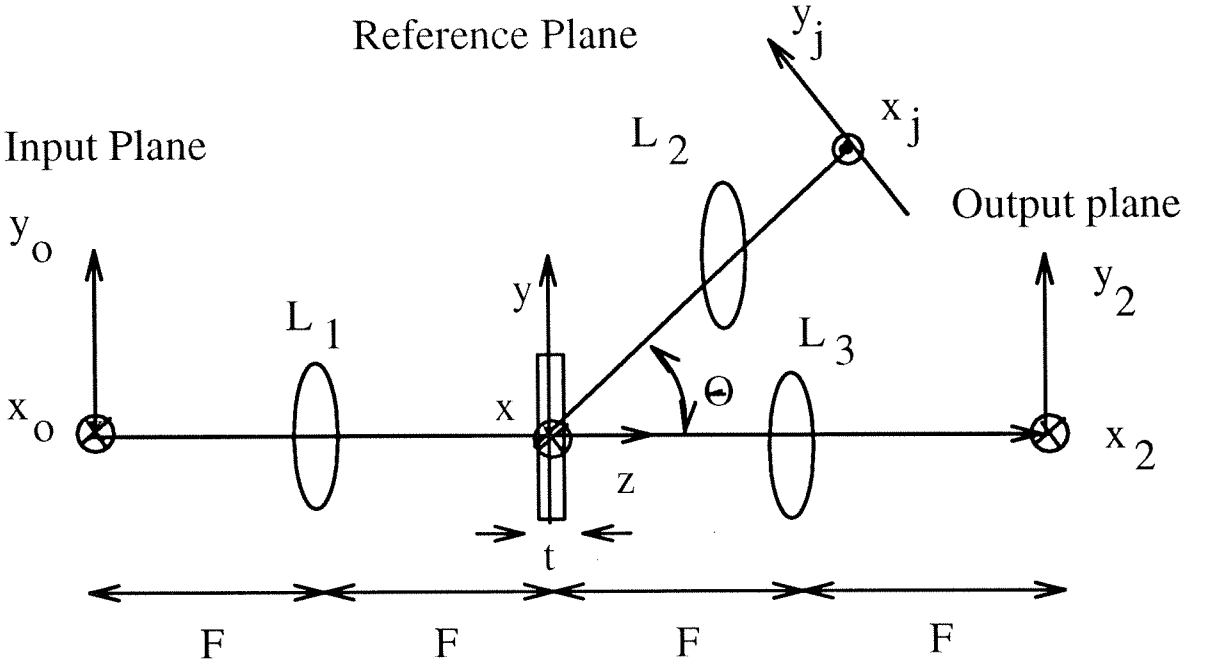


Figure 2.3: Recording and readout geometry for  $\lambda$  multiplexing.

of the medium can be written as

$$\Delta\epsilon \sim \sum_{m=-M}^M R_m^* S_m + c.c. \quad (2.1)$$

where *c.c.* is the complex conjugate term. The reference wave is given by

$$R_m = e^{ik_m \cdot r}, \quad (2.2)$$

where  $k_m$  is the wave vector coming from the reference point source with wavelength  $\lambda_m$ . In addition, using standard Fourier-optics analysis [182], the  $m$ th signal beam can be expressed as

$$S_m(r) \approx e^{i\frac{2\pi}{\lambda_m}z} \iint dx_o dy_o f_m(x_o, y_o) e^{-i\frac{2\pi}{\lambda_m F}(xx_o + yy_o)} e^{-i\frac{\pi z}{\lambda_m F^2}(x_o^2 + y_o^2)}. \quad (2.3)$$

In the expression above,  $f_m(x_o, y_o)$  is the  $m$ -th object image,  $x, y, z$  are the coordinates at the back focal plane of lens  $L_1$ , and  $F$  is the focal length of all three lenses  $L_1, L_2, L_3$ . By substituting Equations 2.2 and 2.3 into Equation 2.1,  $\Delta\epsilon$  can be written out explicitly.

### 2.2.3 Calculation of Cross-talk

The recorded medium is illuminated with one of the reference plane waves. Ideally, this reconstructs the hologram associated with that particular reference. In what follows we will calculate the deviations of the reconstructed wave from the desired image. The diffracted plane wave component,  $E(\vec{k}_d)$ , with wave vector  $\vec{k}_{d,j}$  and wavelength  $\lambda_j$ , can be derived from standard scalar diffraction theory [183] (see Appendix 2.0 for details). The following expression assumes that the Born and paraxial approximations are valid.

$$E(\vec{k}_d) \approx \int d\vec{r} e^{-i\vec{K} \cdot \vec{r}} \Delta\epsilon(\vec{r}), \quad (2.4)$$

where  $\vec{K} = \vec{k}_{d,j} - \vec{k}_j$  and  $\vec{k}_j$  is the illuminating reference beam with wavelength  $\lambda_j$ . The stored perturbation  $\Delta\epsilon$  contains a grating vector component  $\vec{K}$  that will give the desired response as a strong contribution to the integral in Equation 2.4. However, the finite volume over which the integration is carried out causes frequency components other than  $\vec{K}$  to also yield non-zero contributions to the above integral. Therefore, Equation 2.4 describes both the diffracted light due to the desired  $S_j$  plus the noise due to the diffracted light from the other holograms.

Using Equations 2.1 – 2.4, the diffracted light due to the presence of the permittivity perturbation can be explicitly stated. This result can then be related to the detected pattern by using the paraxial approximation to describe how lens  $L_3$  maps  $k_{d,j}$  to a point  $x_2, y_2$  at the output plane:

$$k_{d,j} = \left( \frac{2\pi x_2}{\lambda_j F}, \frac{2\pi y_2}{\lambda_j F}, \frac{2\pi}{\lambda_j} \left( 1 - \frac{x_2^2}{2F^2} - \frac{y_2^2}{2F^2} \right) \right). \quad (2.5)$$

Carrying out the integration over the volume of the medium, the electric field at the output plane can be written as

$$E(x_2, y_2) \approx \sum_{m=-M}^M \iint dx_o dy_o f_m(x_o, y_o) \operatorname{sinc} \left( \frac{a}{2\pi} \left( \Delta K_{mjx} + \frac{2\pi}{F} \left( \frac{x_2}{\lambda_j} + \frac{x_o}{\lambda_m} \right) \right) \right)$$

$$\begin{aligned} & \times \text{sinc}\left(\frac{b}{2\pi}\left(\Delta K_{m jy} + \frac{2\pi}{F}\left(\frac{y_2}{\lambda_j} + \frac{y_o}{\lambda_m}\right)\right)\right) \\ & \times \text{sinc}\left(\frac{t}{2\pi}\left(\Delta K_{m jz} + \left(\frac{2\pi}{\lambda_j} - \frac{2\pi}{\lambda_m}\right) + \frac{2\pi}{F^2}\left(\frac{x_o^2 + y_o^2}{\lambda_m} - \frac{x_2^2 + y_2^2}{\lambda_j}\right)\right)\right). \end{aligned} \quad (2.6)$$

Making use of the assumption that the transverse dimensions of the medium ( $a$  and  $b$  in Eq. 2.6) are much larger than the spatial bandwidth of the images so that the transverse sinc functions can be approximated as delta functions, the electric field at the output plane can be written as

$$\begin{aligned} E(x_2, y_2) & \approx \sum_{m=-M}^M f_m\left(\frac{-\lambda_m}{\lambda_j}x_2 - \frac{F\lambda_m}{2\pi}\Delta K_{m jx}, \frac{-\lambda_m}{\lambda_j}y_2 - \frac{F\lambda_m}{2\pi}\Delta K_{m jy}\right) \\ & \times \text{sinc}\left(\frac{t}{2\pi}\left(\Delta K_{m jz} + \left(\frac{2\pi}{\lambda_j} - \frac{2\pi}{\lambda_m}\right) + \frac{\pi(x_2^2 + y_2^2)}{F^2}\left(\frac{\lambda_m}{\lambda_j^2} - \frac{1}{\lambda_j}\right) + \right.\right. \\ & \quad \left.\left. \frac{\lambda_m}{\lambda_j F}(\Delta K_{m jx}x_2 + \Delta K_{m jy}y_2) + \frac{\lambda_m}{4\pi}(\Delta K_{m jx}^2 + \Delta K_{m jy}^2)\right)\right), \end{aligned} \quad (2.7)$$

where  $\Delta K_{mj} = k_m - k_j$  is the difference between the  $m$ th reference wave vector  $k_m$ , and the illuminating beam's wave vector  $k_j$  and  $\Delta K_{mj\alpha}$  is the component of  $\Delta K_{mj}$  in the  $\alpha$  direction. In Equation 2.7,  $t$  is the thickness of the material in the  $z$  direction. Notice that when  $k_m = k_j$ ,  $E_j(x_2, y_2) = f_j(-x_2, -y_2)$  which is just the mirror-image of the desired stored image.

The cross-talk noise arises from the  $m \neq j$  terms in Equation 2.7. We first write explicit expressions for the Bragg-mismatch terms:

$$\begin{aligned} \Delta K_{m jx} & = 0 \\ \Delta K_{m jy} & = -\left(\frac{2\pi}{\lambda_m} - \frac{2\pi}{\lambda_j}\right) \sin \theta \\ \Delta K_{m jz} & = -\left(\frac{2\pi}{\lambda_m} - \frac{2\pi}{\lambda_j}\right) \cos \theta \end{aligned} \quad (2.8)$$

where  $\theta$  is the angle between the reference beam and the optical axis of the output plane. To estimate the noise-to-signal ratio (NSR) we divide the total average noise power by the magnitude squared of the signal. We assume that each pixel of the

stored images is an independent random variable taking the values minus one and one with equal probability. In this case, the NSR is given by

$$\text{NSR} = \sum_{m \neq j} \text{sinc}^2 \left( -t \left( \frac{1}{\lambda_m} - \frac{1}{\lambda_j} \right) \cos \theta - t \left( \frac{1}{\lambda_m} - \frac{1}{\lambda_j} \right) + \frac{t(x_2^2 + y_2^2)}{2F^2} \left( \frac{\lambda_m}{\lambda_j^2} - \frac{1}{\lambda_j} \right) - \frac{t\lambda_m y_2}{F\lambda_j} \left( \frac{1}{\lambda_m} - \frac{1}{\lambda_j} \right) \sin \theta + \frac{t\lambda_m}{2} \left( \frac{1}{\lambda_m} - \frac{1}{\lambda_j} \right)^2 \sin^2 \theta \right). \quad (2.9)$$

In order to minimize the NSR, the spacing between the holograms in frequency (or wavelength) is chosen to place the center pixel of each hologram at the zero of the sinc function of the two adjacent holograms:  $\Delta\nu = c/[t(1 + \cos \theta)] \equiv \Delta$ .

### 2.2.4 SNR Results

In order to calculate the NSR using this frequency schedule, we assume that the individual frequencies are given in terms of  $\nu_m = m\Delta + \nu_o$ . Rewriting Equation 2.9 in terms of frequency and substituting in the expression for the individual frequencies results in

$$\text{NSR} = \sum_{m \neq j} \text{sinc}^2 \left( -(m - j) - \frac{(x_2^2 + y_2^2)(j\Delta + \nu_o)}{2F^2(1 + \cos \theta)(m\Delta + \nu_o)}(m - j) - \left( \frac{j\Delta + \nu_o}{m\Delta + \nu_o} \right) \frac{y_2(m - j)}{F(1 + \cos \theta)} \sin \theta + \frac{c(m - j)^2}{2t(m\Delta + \nu_o)(1 + \cos \theta)^2} \sin^2 \theta \right). \quad (2.10)$$

The NSR is a function of the hologram number  $j$  (or equivalently the reconstructing  $\lambda_j$ ) and the location at the output plane ( $x_2, y_2$ ). Given  $t, \nu_o, F, \theta$ , and  $N$ , the above expression can be numerically evaluated. The NSR on the output plane for the hologram stored at the center frequency ( $j = 0$ ) was calculated by numerically evaluating Equation 2.10 with  $F = 30$  cm,  $t = 1$  cm,  $x_{2\max} = y_{2\max} = 1.5$  cm,  $\nu_o = 6 \times 10^{14}$  Hertz, and  $N = 401$  holograms. Figure 2.4 shows the output plane for  $\theta = 0.0^\circ$  while Figure 2.5 shows the output plane for  $\theta = 15.0^\circ$ .

In both cases, the maximum NSR is at the edges of the output plane. Setting  $x_2$  and  $y_2$  to their maximum values, we can plot the worst case SNR vs hologram

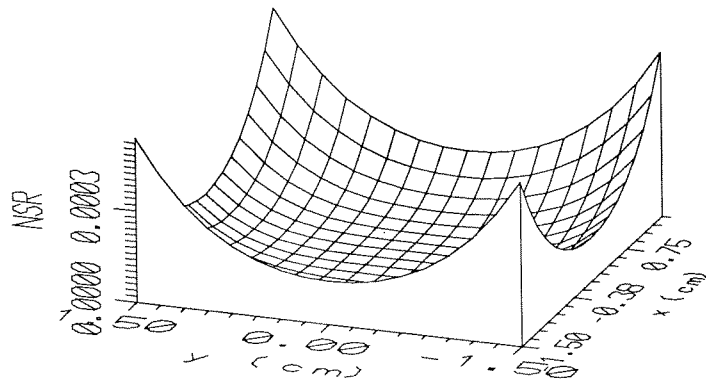


Figure 2.4: NSR at the output plane for the  $j = 0$  hologram with  $\theta = 0.0^\circ$ .

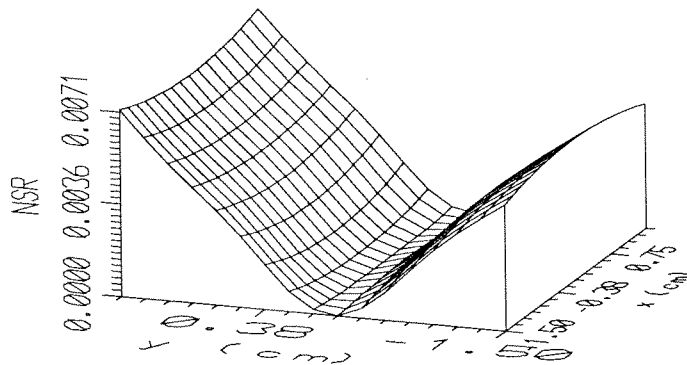


Figure 2.5: NSR at the output plane for the  $j = 0$  hologram with  $\theta = 15.0^\circ$ .

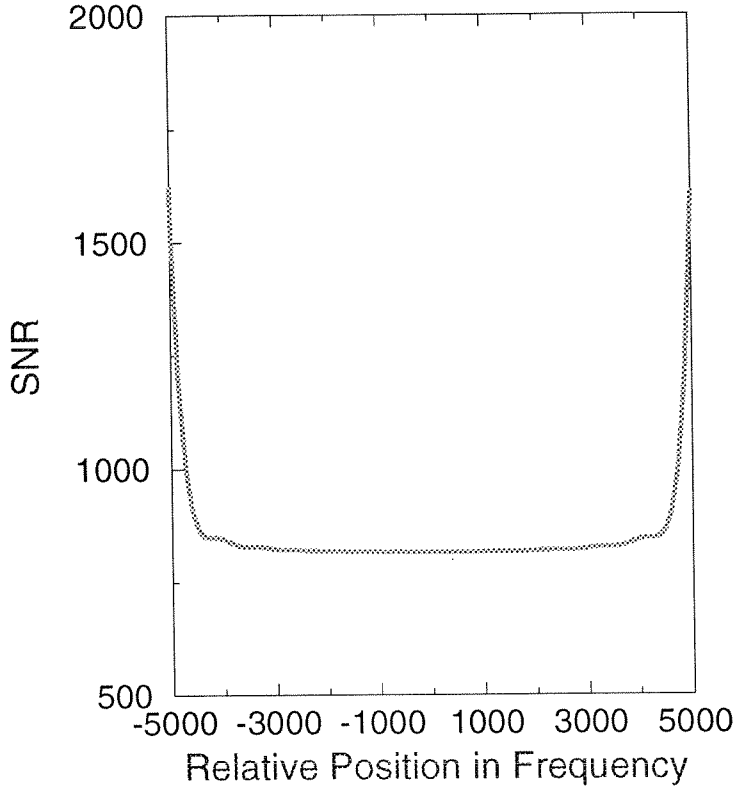


Figure 2.6: The SNR vs hologram number ( $j$ ) for  $N = 10,001$  holograms.

number ( $j$ ). Figure 2.6 shows the results for  $\theta = 0.0^\circ$  with  $N = 10,001$  holograms. Notice that the last ( $j = \pm 5,000$ ) holograms for both cases has a SNR that is twice as large as the middle ( $j = 0$ ) hologram. This is because the  $j = \pm M$  holograms only have holograms on one side. Therefore, the minimum SNR hologram is always the middle hologram in the frequency schedule. However, even this worst case SNR is more than 800 for  $\theta = 0.0^\circ$  which is excellent.

From Figure 2.6 we know that the middle ( $j = 0$ ) hologram has the minimum SNR and that the noise is symmetric about this hologram. Therefore, we can set  $j = 0$  in Equation 2.10 and use it to find a closed-form solution for the maximum NSR. Let us first define  $a$  as follows:

$$a = \frac{(x_2^2 + y_2^2)}{2F^2(1 + \cos \theta)} + \frac{y_2}{F(1 + \cos \theta)} \sin \theta. \quad (2.11)$$

We can drop the last term in the argument of the sinc function in Equation 2.10

since it only becomes comparable to the other terms when  $m$  is large. When  $m$  is large the sinc function itself adds a negligible contribution to the summation in Equation 2.10. Assuming that  $\frac{\nu_o}{m\Delta + \nu_o} \approx 1$  and using the identities  $\sin(x + y) = \sin x \cos y + \cos x \sin y$  and  $\sin^2 x = \frac{1}{2} - \frac{1}{2} \cos 2x$ , Equation 2.10 can be written as

$$\text{NSR}_{\max} = \frac{1}{\pi^2(1+a)^2} \times \left( \sum_{m=1}^M \frac{1}{m^2} - \sum_{m=1}^M \frac{\cos(2\pi am)}{m^2} \right). \quad (2.12)$$

These sums can be approximately evaluated for  $M$  very large by using the sums:

$$\begin{aligned} \sum_{k=1}^{\infty} \frac{\cos kx}{k^2} &= \frac{\pi^2}{6} - \frac{\pi x}{2} + \frac{x^2}{4} \quad \text{for } 0 \leq x \leq 2\pi, \\ \sum_{k=1}^{\infty} \frac{1}{k^2} &= \zeta(2) = \frac{\pi^2}{6}, \end{aligned} \quad (2.13)$$

from references [184] and [185] respectively. Using these expressions, the  $\text{NSR}_{\max}$  goes to

$$\text{NSR}_{\max} = \frac{a}{(1+a)^2} - \frac{a^2}{(1+a)^2}. \quad (2.14)$$

This can be further simplified, if  $a \ll 1$  (true for typical lens systems), into

$$\text{NSR}_{\max} \approx a \approx \frac{(x_2^2 + y_2^2)}{2F^2(1 + \cos \theta)} + \frac{y_2}{F(1 + \cos \theta)} \sin \theta, \quad (2.15)$$

as the number of holograms stored becomes very large. Notice that the best SNR is achieved in the  $\theta = 0^\circ$  geometry. In this case the SNR becomes

$$\text{SNR}_{\min} \approx \frac{4F^2}{(x_2^2 + y_2^2)}. \quad (2.16)$$

Figure 2.7 is a plot of the SNR at the worst point in the output plane ( $x_2 = y_2 = 1.5\text{cm}$ ) vs the total number of holograms stored using the same parameters as given for Figures 2.4 and 2.5. Equation 2.15 above predicts that at  $\theta = 0^\circ$  the SNR goes to 800 and at  $\theta = 15.0^\circ$  the SNR goes to approximately 127. The plots show that at  $N = 4,000$  holograms the SNR is 837 and 131 respectively, showing that the asymptotic expression we derived is in good agreement with the actual values for typical parameters.

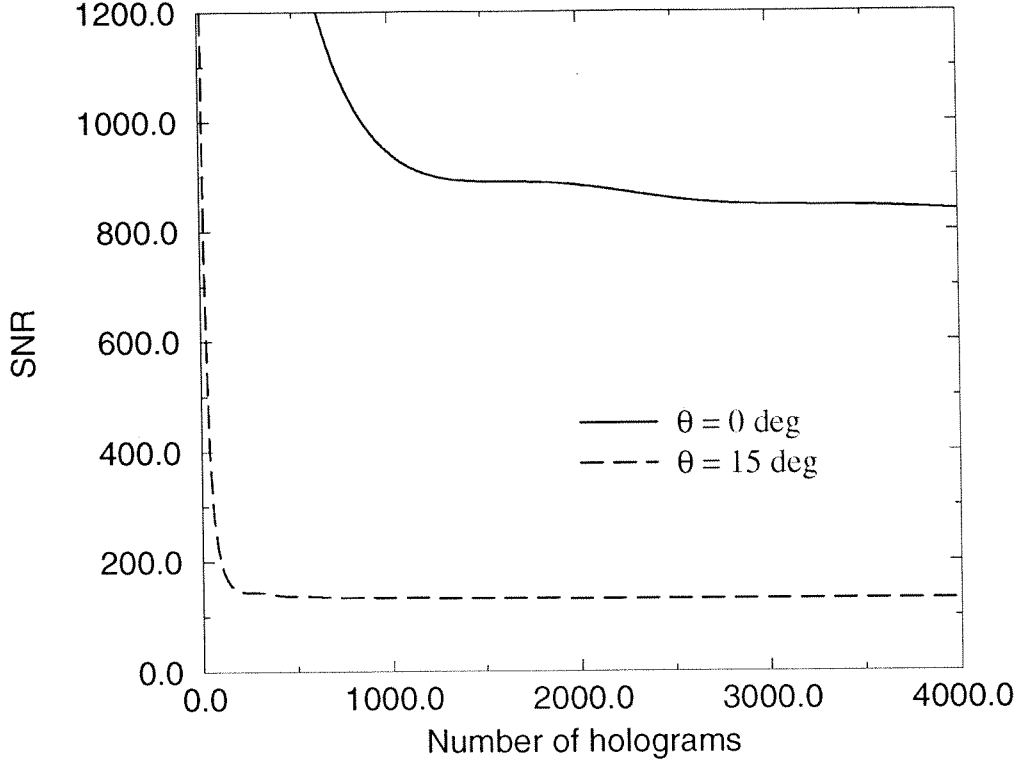


Figure 2.7: The SNR vs the total number of holograms.

## 2.3 Angle Multiplexing

### 2.3.1 Introduction

Cross-talk for angle multiplexed Fourier transform holograms is calculated for a recording schedule for which the center of each stored image coincides with the nulls of the selectivity function for the adjacent reference beams. The main result of this section is the closed-form expression that shows that the SNR goes down as one over the total number of holograms stored.

### 2.3.2 Theory for Angle Multiplexing

Fourier-transform holograms are angle multiplexed in a volume holographic medium using the setup shown in Figure 2.8. A plane wave  $R_m$  is generated with wavelength  $\lambda$  and at an angle  $\theta$  with respect to the  $z$  axis by collimating light from a point source. To change the angle the point source is moved in the  $y_j$  direction. The angle is only



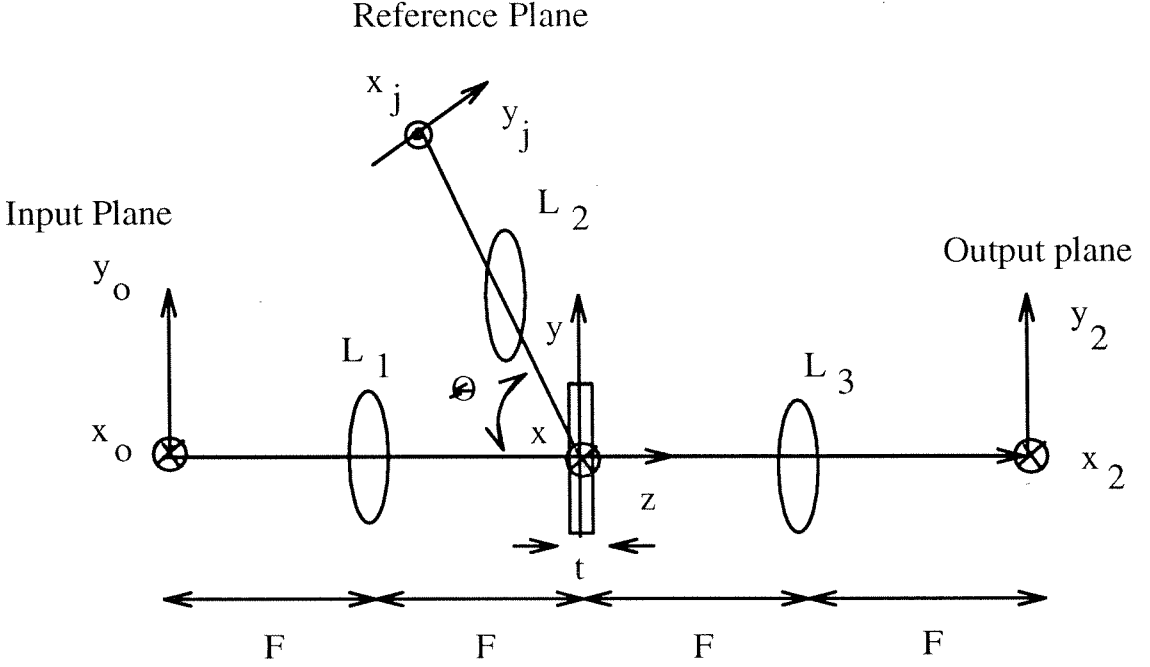


Figure 2.8: Recording and readout geometry for angle multiplexing.

changed in one dimension to eliminate the degenerate grating problem. This plane wave interferes inside the holographic medium with  $S_m$  – the Fourier-transform of the  $m$ th object image. By labeling the holograms  $m = -M, -(M - 1), \dots, 0, \dots, M$ , the presence of these  $N = 2M + 1$  holograms modulates the permittivity of the material such that the change in the permittivity of the medium can be written as

$$\Delta\epsilon \approx \sum_{m=-M}^M R_m^* S_m + c.c. \quad (2.17)$$

where *c.c.* is the complex conjugate term. The reference wave is given by

$$R_m = e^{ik_m \cdot r}, \quad (2.18)$$

where  $k_m$  refers to the wave vector coming from the reference point source at  $y_m$ . In addition, using standard Fourier-optics analysis [182], the  $m$ th signal beam can be expressed as

$$S_m(r) \approx e^{i\frac{2\pi}{\lambda}z} \iint dx_o dy_o f_m(x_o, y_o) e^{-i\frac{2\pi}{\lambda F}(xx_o + yy_o)} e^{-i\frac{\pi z}{\lambda F^2}(x_o^2 + y_o^2)}. \quad (2.19)$$

In the expression above,  $f_m(x_o, y_o)$  is the  $m$ -th object image,  $x, y, z$  are the coordinates at the back focal plane of lens  $L_1$ , and  $F$  is the focal length of all three lenses  $L_1, L_2, L_3$ . By substituting Equations 2.18 and 2.19 into Equation 2.17,  $\Delta\epsilon$  can be written out explicitly.

### 2.3.3 Calculation of Cross-talk

In readout, the recorded medium is illuminated with one of the reference plane waves. Ideally, this would only reconstruct the hologram associated with that particular reference beam. In what follows, we will calculate the deviations of the reconstructed wave from the desired image. The amplitude of the diffracted plane wave components,  $E(\vec{k}_d)$ , with wave vectors  $\vec{k}_d$ , can be derived from standard scalar diffraction theory [183]. The following expression assumes that the Born and paraxial approximations are valid.

$$E(\vec{k}_d) \approx \int d\vec{r} e^{-i\vec{K} \cdot \vec{r}} \Delta\epsilon(\vec{r}), \quad (2.20)$$

where  $\vec{K} = \vec{k}_d - \vec{k}_j$ ,  $\vec{k}_j$  is the wave vector of the illuminating reference beam.

Lens  $L_3$  maps each diffracted plane wave component with wave vector  $\vec{k}_d$  to a point  $x_2, y_2$  at the output plane given by:

$$\vec{k}_d = \left( \frac{2\pi x_2}{\lambda F}, \frac{2\pi y_2}{\lambda F}, \frac{2\pi}{\lambda} \left( 1 - \frac{x_2^2}{2F^2} - \frac{y_2^2}{2F^2} \right) \right), \quad (2.21)$$

assuming the paraxial approximation. Making use of the assumption that the transverse dimensions of the medium are much larger than the spatial bandwidth of the images, substituting relations 2.18, 2.17, 2.19, and 2.21 into Equation 2.20, and carrying out the integration over the volume of the medium, the electric field at the output plane can be written as

$$E(x_2, y_2) \approx \sum_{m=-M}^{M-1} f_m \left( -x_2 - \frac{F\lambda}{2\pi} \Delta K_{mjx}, -y_2 - \frac{F\lambda}{2\pi} \Delta K_{m jy} \right) \times \text{sinc} \left( \frac{t}{2\pi} \left( \Delta K_{mjz} + \frac{1}{F} (\Delta K_{mjx} x_2 + \Delta K_{m jy} y_2) + \frac{\lambda}{4\pi} (\Delta K_{mjx}^2 + \Delta K_{m jy}^2) \right) \right), \quad (2.22)$$

where  $\Delta \vec{K}_{mj} = \vec{k}_m - \vec{k}_j$  is the difference between the  $m$ -th reference wave vector  $\vec{k}_m$  and the illuminating beam's wave vector  $\vec{k}_j$ . Also  $\Delta K_{mj\alpha}$  is the component of  $\Delta \vec{K}_{mj}$  in the  $\alpha$  direction. In Equation 2.22,  $t$  is the thickness of the material in the  $z$  direction.

If  $\Delta K_{mj\alpha} = 0$  for all  $m, j, \alpha$  then  $E(x_2, y_2) = N f_m(-x_2, -y_2)$  and we get our desired reconstruction. When  $\Delta K_{mj\alpha} \neq 0$ , cross-talk noise arises from the  $m \neq j$  terms in Equation 2.22. We can write the components of  $\Delta \vec{K}_{mj}$  in terms of the reference coordinates  $y_m$  and  $y_j$ :

$$\begin{aligned}\Delta K_{mjx} &= 0, \\ \Delta K_{m jy} &= \frac{2\pi}{\lambda F} \cos \theta (y_j - y_m) + \frac{\pi}{\lambda F^2} \sin \theta (y_m^2 - y_j^2) \\ \Delta K_{mjz} &= \frac{2\pi}{\lambda F} \sin \theta (y_j - y_m) + \frac{\pi}{\lambda F^2} \cos \theta (y_j^2 - y_m^2),\end{aligned}\tag{2.23}$$

where  $\theta$  is the angle between the normal to the reference plane and the optical axis of the input plane. To estimate the noise-to-signal ratio (NSR), we divide the total average noise power by the magnitude squared of the high signal level. We assume that each pixel of the stored images is an independent random variable taking the values minus one and one with equal probability. This results in

$$\text{NSR} = \sum_m \text{sinc}^2 \left( \frac{t}{2\pi} \left( \Delta K_{mjz} + \frac{\Delta K_{m jy} y_2}{F} + \frac{\lambda \Delta K_{m jy}^2}{4\pi} \right) \right).\tag{2.24}$$

The spacing between the reference pixels is chosen be  $\Delta y = \frac{\lambda F}{t \sin \theta} \equiv \Delta$ . This makes the sinc function of Equation 2.24 essentially zero at the center pixel ( $y_2 = 0$ ) of each hologram and hence minimizes cross-talk at that location.

### 2.3.4 SNR Results

In order to calculate the NSR, we assume that  $\theta = 90^\circ$  to maximize capacity and that the reference pixels are spaced by  $\Delta$ . Rewriting Equation 2.24 and keeping terms to

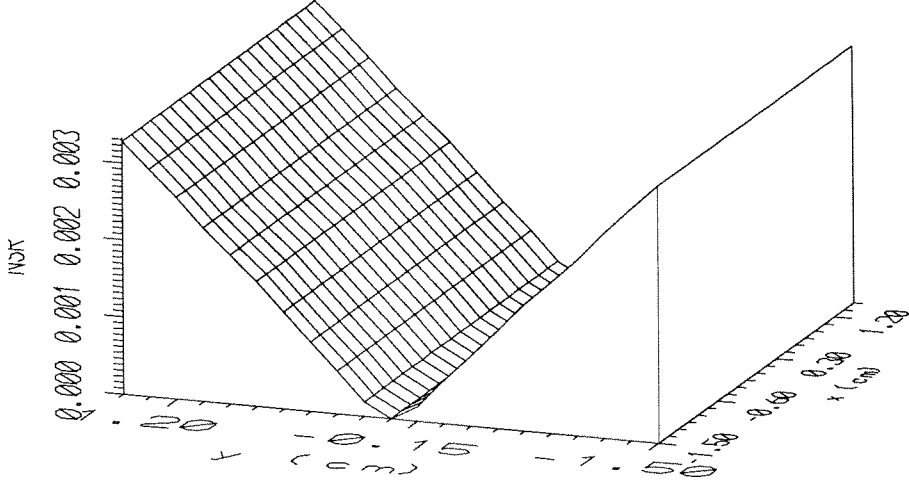


Figure 2.9: NSR at the output plane for the  $j = 0$  hologram.

third order in  $(y/F)$  results in

$$\text{NSR} = \sum_{m \neq j} \text{sinc}^2 \left( m - j + \frac{y_2 \lambda}{2Ft} (m^2 - j^2) \right). \quad (2.25)$$

The NSR is a function of the hologram number  $j$  (or equivalently the reconstructing angle) and the location at the output plane  $(x_2, y_2)$ . Given  $t$ ,  $\lambda$ ,  $F$ ,  $\theta$ , and  $N$ , the above expression can be numerically evaluated. The NSR on the output plane for the hologram stored at the center angle ( $j = 0$ ) was calculated by numerically evaluating Equation 2.25 with  $F = 30$  cm,  $t = 1$  cm,  $x_{2\max} = y_{2\max} = 1.5$  cm,  $\lambda = 500$  nm, and  $N = 4,001$  holograms, and is shown in Figure 2.9. The maximum NSR is always at the edges of the output plane. Setting  $x_2$  and  $y_2$  to their maximum values, we can plot the worst case SNR vs hologram number ( $j$ ). Figure 2.10 shows the results for  $N = 4,001$  holograms. Notice that the last ( $j = \pm 2,000$ ) holograms for both cases has a SNR that is twice as large as the worst hologram. This is because the  $j = \pm M$  holograms only have holograms on one side. The minimum SNR hologram is close to these  $j = \pm M$  holograms in the angular schedule with the middle hologram having

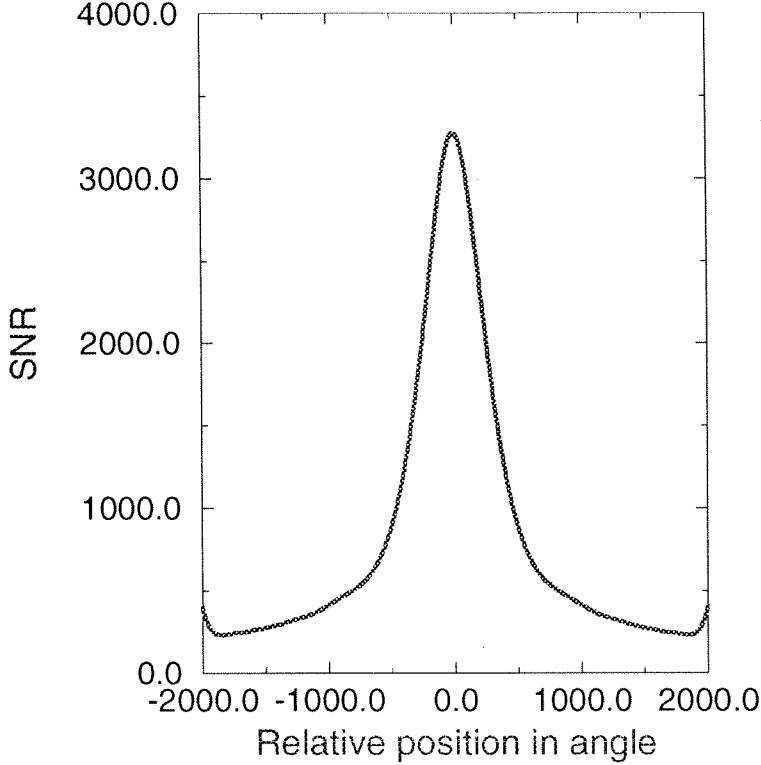


Figure 2.10: The SNR vs hologram position ( $j$ ) for  $N = 4,001$  holograms.

the best SNR. This is because the angular separation between holograms was set for the middle hologram and as the multiplexing angle deviates from this optimum angle the spacing required to put the holograms in the nulls changes. Thus the holograms multiplexed at an angle other than the center are noisier.

From Figure 2.10 we know that the end ( $j = \pm M$ ) holograms have the minimum SNR and that the noise is symmetric about the center hologram. Therefore, we can set  $j = M$  in Equation 2.25 (multiply the NSR by 2 make take into account the  $j = \pm M$  holograms only have half the neighboring holograms) and use it to find a closed-form solution for the maximum NSR. By defining  $a$  as

$$a = \frac{\lambda y_2 M}{tF}, \quad (2.26)$$

rearranging the sum, and dropping one term because it does not contribute signifi-

cantly to the sum, Equation 2.25 can be written as

$$\text{NSR} = 2 \sum_{n=1}^N \text{sinc}^2(n + an). \quad (2.27)$$

Using the identities  $\sin(x + y) = \sin x \cos y + \cos x \sin y$  and  $\sin^2 x = \frac{1}{2} - \frac{1}{2} \cos 2x$ , Equation 2.27 can be written as

$$\text{NSR}_{\max} = \frac{1}{\pi^2(1+a)^2} \times \left( \sum_{n=1}^N \frac{1}{n^2} - \sum_{n=1}^N \frac{\cos(2\pi an)}{n^2} \right). \quad (2.28)$$

These sums can be approximately evaluated for  $N$  very large by using the sums:

$$\begin{aligned} \sum_{k=1}^{\infty} \frac{\cos kx}{k^2} &= \frac{\pi^2}{6} - \frac{\pi x}{2} + \frac{x^2}{4} \quad \text{for } 0 \leq x \leq 2\pi, \\ \sum_{k=1}^{\infty} \frac{1}{k^2} &= \zeta(2) = \frac{\pi^2}{6}, \end{aligned} \quad (2.29)$$

from references [184] and [185] respectively. Using these expressions, the  $\text{NSR}_{\max}$  goes to

$$\text{NSR}_{\max} = \frac{a}{(1+a)^2} - \frac{a^2}{(1+a)^2}. \quad (2.30)$$

This can be further simplified, if  $a \ll 1$  (true for typical lens systems), into

$$\text{NSR}_{\max} \approx a \approx \frac{\lambda y_2 N}{2tF}, \quad (2.31)$$

as the number of holograms stored becomes very large. Thus the worst case SNR goes as

$$\text{SNR}_{\min} \approx \frac{2tF}{\lambda y_2 N}. \quad (2.32)$$

Figure 2.11 is a plot of the SNR at the worst point in the output plane ( $y_2 = 1.5\text{cm}$ ) vs the total number of holograms stored using the same parameters as given for Figures 2.9. Notice that the worst case SNR goes down as  $1/N$  as predicted.

## 2.4 Phase Code Multiplexing

### 2.4.1 Introduction

The cross-talk between holograms multiplexed with Walsh-Hadamard phase codes is analyzed. Each hologram is stored with a reference beam that consists of  $N$  phase

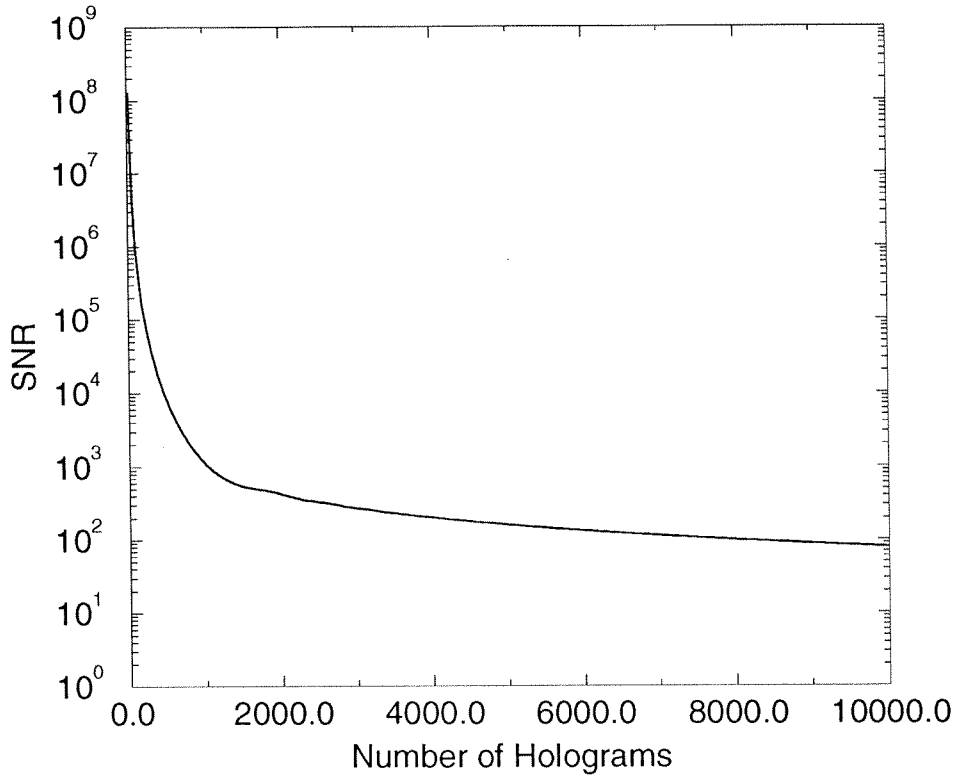


Figure 2.11: The SNR vs the total number of holograms.

coded plane waves. The SNR is calculated for a recording schedule for which the center of each stored image coincides with the nulls of the selectivity function for the adjacent plane wave components of the reference beam. The section contains two main results. First, we show that each phase-coded hologram has almost the same SNR upon reconstruction once we exclude one “bad” code. This is in contrast to  $\theta$  multiplexing where holograms have much better SNR than the worst hologram depending on their angular position with respect to the other holograms. Second, we show that, the SNR for a given number of holograms is better for  $\phi$  than for  $\theta$  multiplexing.

### 2.4.2 Theory for Phase Code Multiplexing

Fourier-transform holograms are multiplexed in a volume holographic medium using the setup shown in Figure 2.12. The  $m$ -th hologram is formed by interfering the

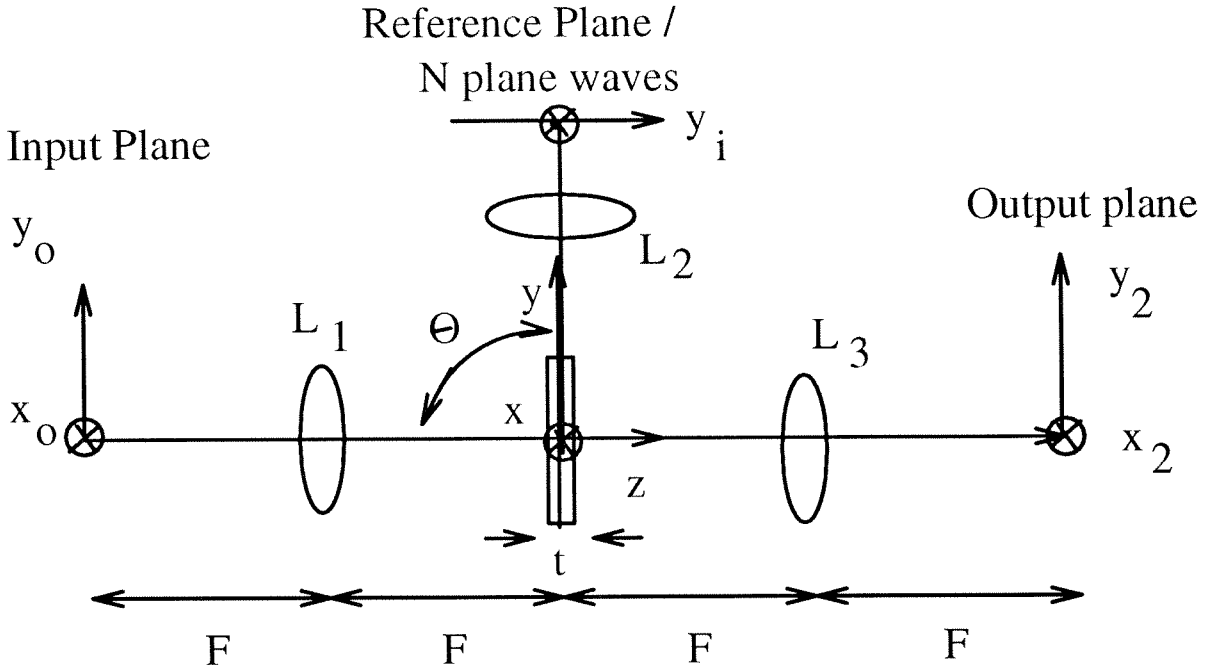


Figure 2.12: Recording and readout geometry for  $\phi$  multiplexing.

signal with a reference beam  $R_m$  that consists of  $N$  plane wave components, with the phase of the  $i$ -th component being modulated by  $P_i^m$ :

$$R_m = \sum_{i=-N/2}^{N/2-1} P_i^m e^{ik_i r}. \quad (2.33)$$

In this section we assume that the  $P_i^m$ 's are either 1 or  $-1$  and that they are the Walsh-Hadamard codes which are orthogonal and easy to construct. The method used to construct them is given in reference [180]. The codes are constructed using the kernel

$$m = \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix}. \quad (2.34)$$

Taking this matrix larger matrices can be constructed by recursively substituting the previous matrix into a new matrix  $m_{i+1}$  shown below:

$$m_{i+1} = \begin{pmatrix} m_i & m_i \\ m_i & -m_i \end{pmatrix}. \quad (2.35)$$

This procedure generates orthogonal codes of length  $2^{(i+1)}$ . Examples of the first 64 codes are shown in Figure 2.13. In the figure above, white stands for a 1 and black



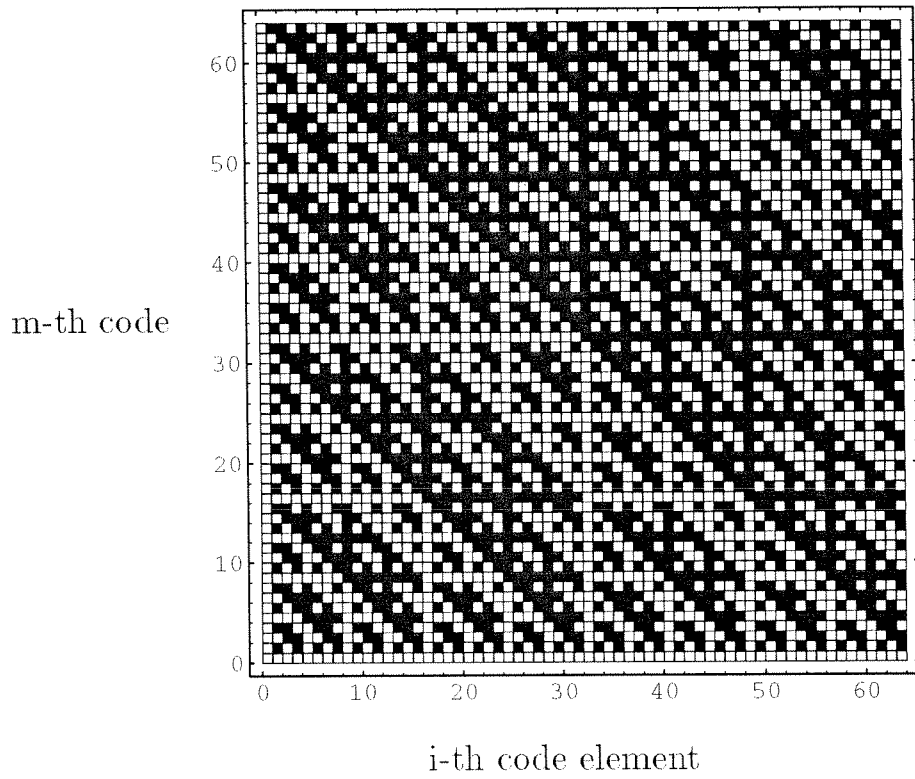


Figure 2.13: First 64 Walsh-Hadamard codes for  $\phi$  multiplexing.

represents a - 1. Notice that the second code has the highest frequency – it alternates 1 and -1. Since Walsh-Hadamard codes are orthogonal and easy to construct they are a logical choice for phase codes for the evaluation of cross-talk in phase coded holographic memory.

This reference beam interferes inside the holographic medium with  $S_m$  – the Fourier-transform of the  $m$ th object image. The presence of these  $N = 2M$  holograms, labeled  $m = -M, -(M-1), \dots, 0, \dots, M-1$ , modulates the permittivity of the material such that the change in the permittivity of the medium can be written as

$$\Delta\epsilon \approx \sum_{m=-M}^{M-1} R_m^* S_m + c.c. \quad (2.36)$$

where  $c.c.$  represents the complex conjugate term. In addition, using standard Fourier-optics analysis [182], the  $m$ -th signal beam can be expressed as

$$S_m(x, y, z) \approx e^{i\frac{2\pi}{\lambda}z} \iint dx_o dy_o f_m(x_o, y_o) e^{-i\frac{2\pi}{\lambda F}(xx_o + yy_o)} e^{-i\frac{\pi z}{\lambda F^2}(x_o^2 + y_o^2)}. \quad (2.37)$$

In the expression above,  $f_m(x_o, y_o)$  is the  $m$ -th object image,  $x, y, z$  are the coordinates at the back focal plane of lens  $L_1$ , and  $F$  is the focal length of all three lenses  $L_1, L_2, L_3$ . By substituting Equations 2.33 and 2.37 into Equation 2.36,  $\Delta\epsilon$  can be written out explicitly.

### 2.4.3 Calculation of Cross-talk

In readout, the recorded medium is illuminated with one of the set of phase coded ( $P_j^n$ ) reference plane waves. Ideally, this would only reconstruct the hologram associated with that particular reference set. In what follows, we will calculate the deviations of the reconstructed wave from the desired image. The amplitude of the diffracted plane wave components,  $E(\vec{k}_d)$ , with wave vectors  $\vec{k}_d$ , can be derived from standard scalar diffraction theory [183]. The following expression assumes that the Born and

paraxial approximations are valid.

$$E(\vec{k}_d) \approx \sum_{j=-M}^{M-1} \int d\vec{r} P_j^n e^{-i\vec{K} \cdot \vec{r}} \Delta\epsilon(\vec{r}), \quad (2.38)$$

where  $\vec{K} = \vec{k}_d - \vec{k}_j$ ,  $\vec{k}_j$  is the wave vector of one of the plane wave components that make up an illuminating reference beam,  $j$  is the read out pixel number corresponding to  $\vec{k}_j$ , and  $n$  is the read out codeword number.

Lens  $L_3$  maps each diffracted plane wave component with wave vector  $\vec{k}_d$  to a point  $x_2, y_2$  at the output plane given by:

$$\vec{k}_d = \left( \frac{2\pi x_2}{\lambda F}, \frac{2\pi y_2}{\lambda F}, \frac{2\pi}{\lambda} \left( 1 - \frac{x_2^2}{2F^2} - \frac{y_2^2}{2F^2} \right) \right), \quad (2.39)$$

assuming the paraxial approximation. Making use of the assumption that the transverse dimensions of the medium are much larger than the spatial bandwidth of the images, substituting relations 2.33, 2.36, 2.37, and 2.39 into Equation 2.38, and carrying out the integration over the volume of the medium, the electric field at the output plane can be written as

$$E(x_2, y_2) \approx \sum_{m=-M}^{M-1} \sum_{j=-M}^{M-1} \sum_{i=-M}^{M-1} P_j^n P_i^{*m} f_m \left( -x_2 - \frac{F\lambda}{2\pi} \Delta K_{ijx}, -y_2 - \frac{F\lambda}{2\pi} \Delta K_{ijy} \right) \\ \times \text{sinc} \left( \frac{t}{2\pi} \left( \Delta K_{ijz} + \frac{1}{F} (\Delta K_{ijx} x_2 + \Delta K_{ijy} y_2) + \frac{\lambda}{4\pi} (\Delta K_{ijx}^2 + \Delta K_{ijy}^2) \right) \right), \quad (2.40)$$

where  $\Delta \vec{K}_{ij} = \vec{k}_i - \vec{k}_j$  is the difference between the  $i$ -th reference wave vector  $\vec{k}_i$  and the illuminating beam's wave vector  $\vec{k}_j$ . Also  $\Delta K_{ij\alpha}$  is the component of  $\Delta \vec{K}_{ij}$  in the  $\alpha$  direction. In Equation 2.40,  $t$  is the thickness of the material in the  $z$  direction.

If  $\Delta K_{ij\alpha} = 0$  for all  $i, j, \alpha$  then  $E(x_2, y_2) = N f_m(-x_2, -y_2)$  and we get our desired reconstruction. When  $\Delta K_{ij\alpha} \neq 0$ , cross-talk noise arises from the  $m \neq n$  and the  $m = n$  but  $i \neq j$  terms in Equation 2.40. We can write the components of  $\Delta \vec{K}_{ij}$  in terms of the reference coordinates  $y_i$  and  $y_j$ :

$$\Delta K_{ijx} = 0,$$

$$\begin{aligned}
\Delta K_{ijy} &= \frac{2\pi}{\lambda F} \cos \theta (y_j - y_i) + \frac{\pi}{\lambda F^2} \sin \theta (y_i^2 - y_j^2) \\
\Delta K_{ijz} &= \frac{2\pi}{\lambda F} \sin \theta (y_j - y_i) + \frac{\pi}{\lambda F^2} \cos \theta (y_j^2 - y_i^2),
\end{aligned} \tag{2.41}$$

where  $\theta$  is the angle between the normal to the reference plane and the optical axis of the input plane. To estimate the noise-to-signal ratio (NSR), we divide the total average noise power by the magnitude squared of the high signal level. We assume that each pixel of the stored images is an independent random variable taking the values minus one and one with equal probability. Using the properties of Hadamard codes, the average NSR is given by

$$\text{NSR} = \frac{1}{N^2} \sum_m \left| \sum_j \sum_{i \neq j} P_j^n P_i^m \text{sinc} \left( \frac{t}{2\pi} \left( \Delta K_{ijz} + \frac{\Delta K_{ijy} y_2}{F} + \frac{\lambda \Delta K_{ijy}^2}{4\pi} \right) \right) \right|^2. \tag{2.42}$$

The spacing between the pixels of the phase code is chosen be  $\Delta y = \frac{\lambda F}{t \sin \theta} \equiv \Delta$ . This makes the sinc function of Equation 2.42 essentially zero at the center pixel ( $y_2 = 0$ ) of each hologram and hence minimizes cross-talk at that location.

#### 2.4.4 SNR Results

In order to calculate the NSR, we assume that  $\theta = 90^\circ$  to maximize capacity and that the reference pixels are spaced by  $\Delta$ . Rewriting Equation 2.42 results in

$$\text{NSR} = \frac{1}{N^2} \sum_m \left| \sum_j \sum_{i \neq j} P_j^n P_i^m \text{sinc} \left( j - i + \frac{y_2 \lambda}{2Ft} (i^2 - j^2) + \frac{\lambda^3}{8t^3} (i^2 - j^2)^2 \right) \right|^2. \tag{2.43}$$

Notice that if the sinc functions are replaced by the Kronecker  $\delta_{ij}$  then there is no cross-talk. The NSR is a function of the location at the output plane in one dimension ( $y_2$ ). Given  $t$ ,  $\lambda$ ,  $F$ ,  $\theta$ , and  $N$ , the above expression can be numerically evaluated. The NSR at the output plane for the n=8 hologram was calculated by numerically evaluating Equation 2.10 with  $F = 30$  cm,  $t = 1$  cm,  $x_{2\max} = y_{2\max} =$

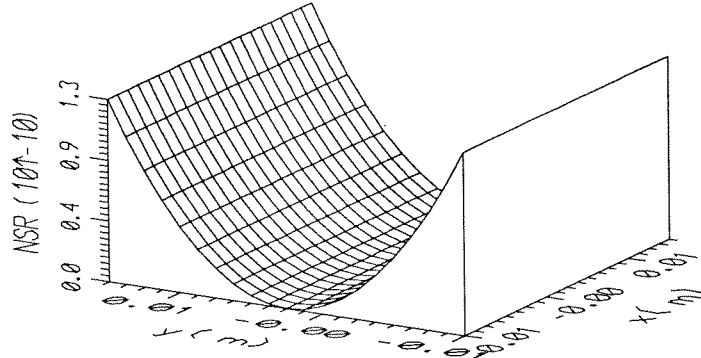


Figure 2.14: NSR vs position  $(x,y)$  on the output plane for the  $n = 8$  hologram with  $N = 16$ .

1.5 cm,  $\lambda = 500$  nm, and  $N = 16$  holograms. Figure 2.14 shows the result plotted as a function of position on the output plane. Notice that the maximum NSR occurs at the  $y_2$  boundary of the output plane.

Setting  $y_2$  to its maximum value, we can plot the worst case SNR vs  $n$ , the phase code number. Figure 2.15 shows the result for  $N = 64$  holograms. Code word number 2 is the highest frequency codeword (1,-1,1,-1 ...) and its structure complements the sinc function. Therefore when phase code number 2 is multiplied with the sinc function in Equation 2.43 we obtain a consistently positive product which adds up to a large NSR. Figure 2.16 plots the product  $P_i^m \text{sinc}\left(j - i + \frac{y_2 \lambda}{2Ft}(i^2 - j^2) + \frac{\lambda^3}{8t^3}(i^2 - j^2)^2\right)$  as a function of  $m$  and  $i' = i - M$  for  $n = 6$  and  $j = 0$ . While this particular codeword has poor SNR, the others average to a much higher SNR. The other codes in Figure 2.15 that have poor SNR are noisy due to cross-talk from code 2. By taking out codeword number 2, it is possible to record holograms that have the same SNR

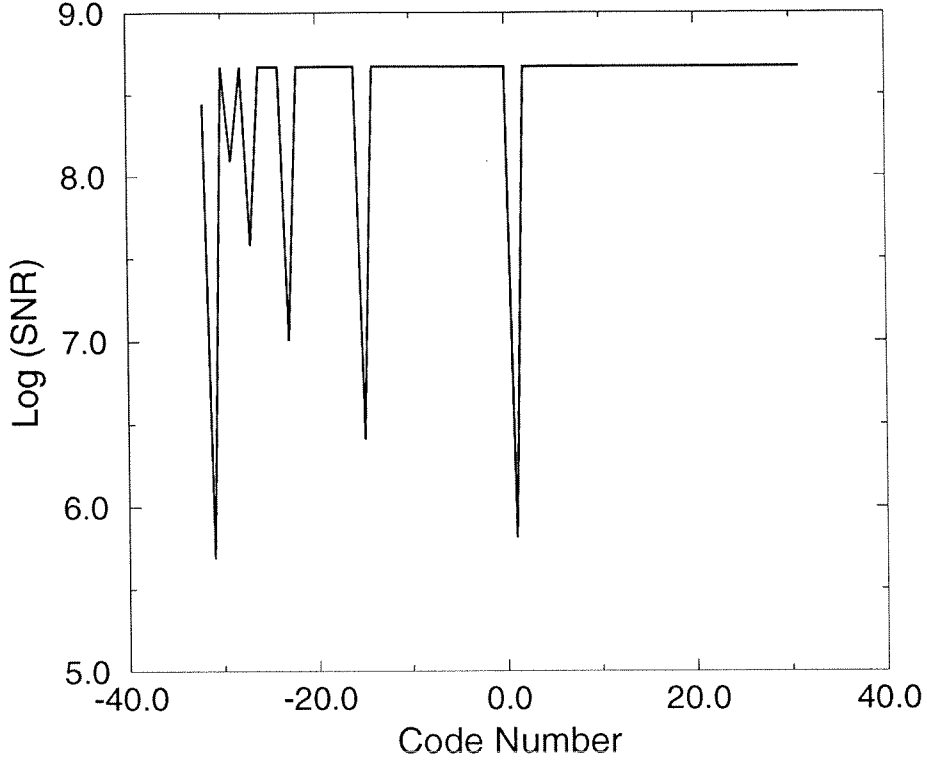


Figure 2.15: The  $\log(\text{SNR})$  vs hologram code number for  $N = 64$  holograms.

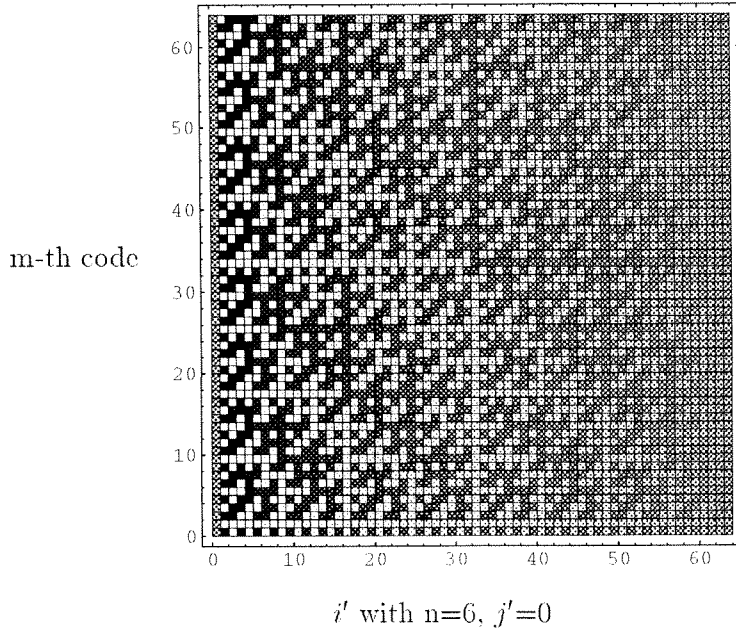


Figure 2.16: Plot of  $P_i^m \text{sinc}\left(j - i + \frac{y_2 \lambda}{2F_t}(i^2 - j^2) + \frac{\lambda^3}{8t^3}(i^2 - j^2)^2\right)$  as a function of  $m$  and  $i' = i - M$  for  $n = 6$  and  $j = 0$ .

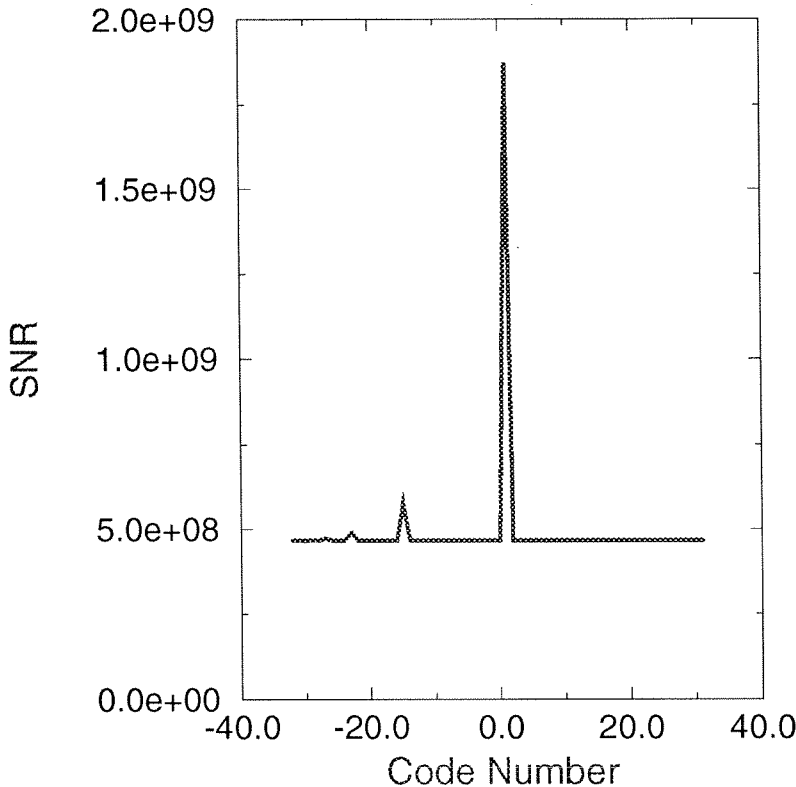


Figure 2.17: The SNR vs hologram code number without code 2.

(to a couple of decimal places) except for a few holograms that have better SNR than the others, as shown in Figure 2.17. This result is different from  $\theta$  multiplexing [170] where most holograms have better SNR than the worst hologram depending on their angular position with respect to the other holograms.

Figure 2.18 shows the worst case SNR for  $\phi$  and  $\theta$  multiplexing with the same parameters as listed above. When using codeword number 2, the worst case SNR decreases very quickly for  $\phi$  multiplexing. Taking this codeword out and recording  $2^n - 1$  holograms with  $2^n$  plane wave references, results in worst case SNR that is about 3 orders of magnitude better than  $\theta$  multiplexing. If we take the SNR for  $\theta$  multiplexing at the maximum  $y_2$ , averaged over all holograms, we obtain an average SNR that is still more than two orders of magnitude less than for  $\phi$  multiplexing for large  $N$ . Therefore, well chosen phase codes suppress, even on average, the cross-talk. For  $\lambda$  multiplexing, the worst case SNR for 500 holograms, with all other parameters

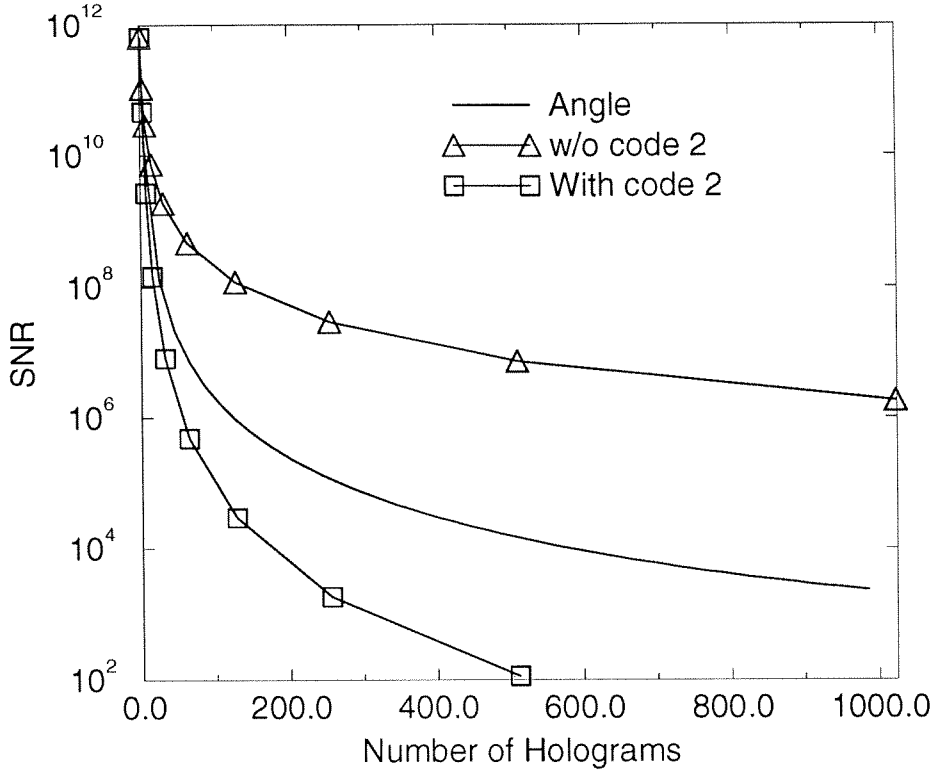


Figure 2.18: The SNR vs the total number of holograms for  $\phi$  (with and w/o code 2) multiplexing, and  $\theta$  multiplexing.

the same as in Figure 2.18, is  $1.4 \times 10^3$  [171], compared to  $10^7$  for  $\phi$  multiplexing and  $10^4$  for  $\theta$  multiplexing. We have not calculated the cross-talk for  $\phi$  multiplexing beyond 1024 holograms. Notice that the cross-talk SNR is very high in all cases, indicating that other system imperfections will most likely determine the fidelity of the reconstruction in practice.

## 2.5 Rotation Multiplexing

### 2.5.1 Introduction

Cross-talk for holograms multiplexed by rotating the storage media is calculated and compared to the angle multiplexed case. Brady and Psaltis [186] argue that rotation multiplexing can potentially have higher storage capacity than angle multiplexing since more of the grating space is accessible. Assuming that the media is a rectan-



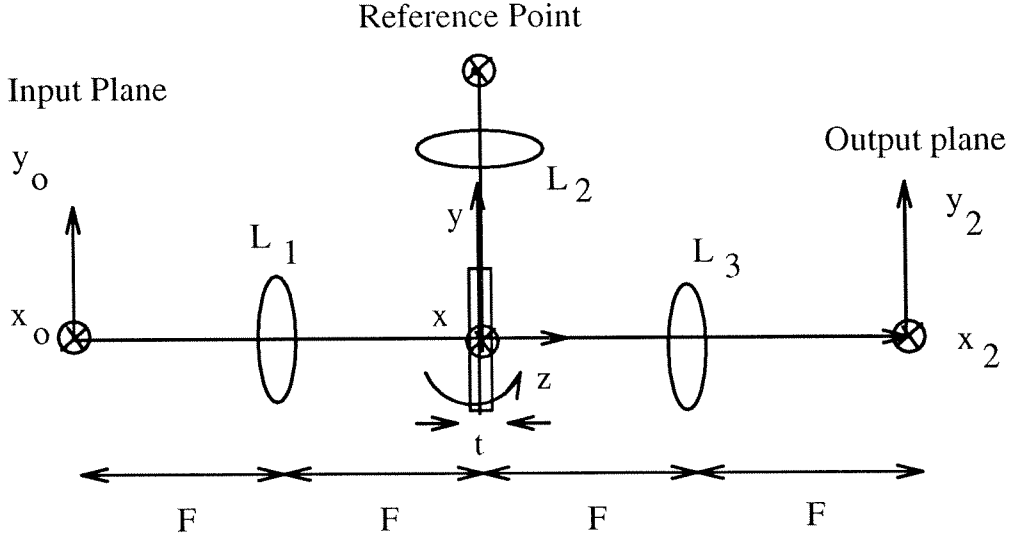


Figure 2.19: Recording and readout geometry for rotation multiplexing.

gular bar, the worst case SNR for rotation multiplexing is shown to have the same dependence on the number of holograms stored as angle multiplexing. For the case of a cylinder, the SNR is shown to saturate as the number of holograms

### 2.5.2 Storing Rotation Multiplexed Holograms

Fourier-transform holograms are rotation multiplexed in a volume holographic medium using the setup shown in Figure 2.19. A plane wave  $R_m$  is generated by collimating light from a reference point source. This interferes inside the holographic medium with  $S_m$  – the Fourier-transform of the  $m$ th object image and stores a single hologram. Multiple holograms are stored by rotating the medium around the  $x$  axis between each new hologram. The holograms are then stored in rotated  $(x'_m, y'_m)$  reference systems as shown in Figure 2.20. Notice that positive rotation is shown as counter-clockwise.

We will define the holograms that are currently being reconstructed as the origin of rotation for  $\theta_m$ , and then label all other holograms with respect to this hologram at rotation  $\theta_i$ . By labeling the reconstructed holograms as  $i = -M, -(M-1), \dots, 0, \dots, M$ ,

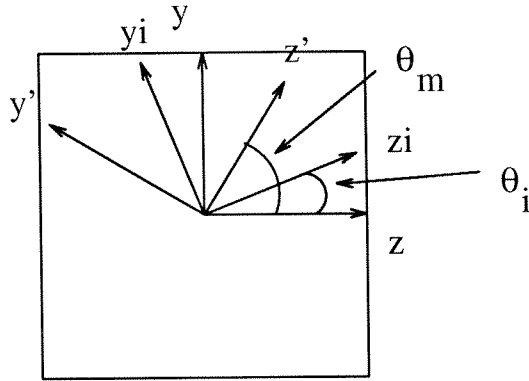


Figure 2.20: Rotated coordinate systems defining  $\theta_m$  and  $\theta_i$ .

the presence of these  $N = 2M + 1$  holograms modulates the permittivity of the material as

$$\Delta\epsilon \approx \sum_m R_m^* S_m + c.c. \quad (2.44)$$

where *c.c.* is the complex conjugate term. Since we rotate the entire setup instead of the storage media, expressions for the reference and signal beams can easily be expressed. The reference wave coincides with the  $y_i$ -axis and is given by

$$R_m = e^{-i\frac{2\pi}{\lambda}y_m}. \quad (2.45)$$

In addition, using standard Fourier-optics analysis [182], the  $m$ th signal beam can be expressed as

$$S_m(r) \approx e^{i\frac{2\pi}{\lambda}z'_m} \iint dx_o dy_o f_m(x_o, y_o) e^{-i\frac{2\pi}{\lambda F}(xx_o + y'_m y_o)} e^{-i\frac{\pi z'_m}{\lambda F^2}(x_o^2 + y_o^2)}. \quad (2.46)$$

In the expression above,  $f_m(x_o, y_o)$  is the  $m$ -th object image, and  $F$  is the focal length of all three lenses  $L_1, L_2, L_3$ . By substituting Equations 2.45 and 2.46 into Equation 2.44,  $\Delta\epsilon$  can be written out explicitly.

### 2.5.3 SNR Calculation for Rotation Multiplexing in Rectangular Media

A rectangular recording medium is illuminated with a reference plane wave in the  $y_i, z_i$  coordinate system. Ideally, this reconstructs the  $i$ -th hologram associated with

that particular rotation. In what follows, we will calculate the deviation of the reconstructed wave from the desired image. The amplitudes  $E(\vec{k}_d)$  of the diffracted plane wave components with wave vectors  $\vec{k}_d$  can be derived from standard scalar diffraction theory [183]. The following expression assumes that the Born and paraxial approximations are valid (small rotations).

$$E(\vec{k}_d) \approx \int d\vec{r} e^{-i\vec{K} \cdot \vec{r}} \Delta\epsilon(\vec{r}), \quad (2.47)$$

where  $\vec{K} \cdot \vec{r} = \vec{k}_d \cdot \vec{r}_m + \frac{2\pi}{\lambda} y_i$ , and  $-\frac{2\pi}{\lambda} y_i$  is the wave vector of the illuminating reference beam. Using Equations 2.44 – 2.47, we can solve explicitly for the diffracted light due to the presence of the permittivity perturbation. This result can then be related to the detected pattern by using the paraxial approximation to describe how lens  $L_3$  maps  $k_d$  to a point  $x_2, y_2$  at the output plane:

$$\vec{k}_d = \left( \frac{2\pi x_2}{\lambda F}, \frac{2\pi y_2}{\lambda F}, \frac{2\pi}{\lambda} \left( 1 - \frac{x_2^2}{2F^2} - \frac{y_2^2}{2F^2} \right) \right). \quad (2.48)$$

Making use of the assumption that the height of the medium in the  $x$  direction is much larger than the spatial bandwidth of the images (forces the resulting sinc function to a delta function in  $x$ ). Transformed back into  $x, y, z$  coordinates and integrating over the volume of the media, the electric field at the output plane can be written as

$$E(x_2, y_2) \approx \sum_m \int dy_o f_m(-x_2, y_o) \times \frac{\sin(At \cos \theta_i + Bt \sin \theta_i)}{At \cos \theta_i + Bt \sin \theta_i} \frac{\sin(Bt \cos \theta_i - At \sin \theta_i)}{Bt \cos \theta_i - At \sin \theta_i}. \quad (2.49)$$

In Equation 2.49,  $t$  is one half of the thickness of the material in the  $y$  and  $z$  directions, while  $B$  is given by

$$B = -\frac{2\pi y_2}{\lambda F} \sin \theta_m - \frac{\pi y_2^2}{\lambda F^2} \cos \theta_m + \frac{2\pi}{\lambda} \sin \theta - \frac{2\pi y_o}{\lambda F} \sin \theta_m + \frac{\pi y_o^2}{\lambda F^2} \cos \theta_m, \quad (2.50)$$

and  $A$  by

$$A = \frac{\pi y_o^2}{\lambda F^2} \sin \theta_m + \frac{2\pi y_o}{\lambda F} \cos \theta_m + \frac{2\pi y_2}{\lambda F} \cos \theta_m - \frac{\pi y_2^2}{\lambda F^2} \sin \theta_m + \frac{2\pi}{\lambda} (1 - \cos \theta_m). \quad (2.51)$$

Equation 2.49 can be approximated by letting the first sinc function go to a  $\delta$  function with integration of  $y_o$ . Using the condition that the argument of the  $\delta$  function must be zero, and that the angles are small, results in  $y_o$  being given by

$$y_o = -y_2 - \frac{F\theta_m^2}{2} - F\theta_m\theta_i. \quad (2.52)$$

Substituting this expression for  $y_o$  into the last sinc and discarding terms that are more than order 3 in  $\theta$  gives an expression for the scattered electric field. We can estimate the noise-to-signal ratio (NSR) by dividing the total average noise power by the signal's magnitude squared. Assuming that each pixel of the stored images is an independent random variable taking the values minus one and one with equal probability, the NSR is given by

$$\text{NSR} = \sum_{m \neq 0} \text{sinc}^2 \left( \frac{L}{\lambda} \theta_m + \frac{Ly_2}{2\lambda F} \theta_m^2 + \frac{Ly_2}{\lambda F} \theta_m \theta_i \right), \quad (2.53)$$

with  $L = 2t$  as the thickness of the bar in the  $y$  and  $z$  directions.

In order to minimize the NSR, the rotational angle between the holograms is chosen to place the center pixel ( $y_2 = 0$ ) of each hologram at the zero of the sinc function of the adjacent holograms. This leads to a rotational angle of  $\Delta\theta_m = \frac{\lambda}{L}$ . In order to calculate the NSR using this rotational schedule, we assume that the individual rotations are given in terms of  $\theta_m = m\frac{\lambda}{L}$ . The NSR is a function of the hologram position in the recording sequence  $i$  (implicit in the limits of the sum), the location at the output plane ( $y_2$ ), and the total number of holograms stored ( $N$ ). Given  $L$ ,  $F$ ,  $\lambda$ , and  $N$ , the above expression can be numerically evaluated.

The NSR on the output plane for the  $i = 100$  hologram was calculated by numerically evaluating Equation 2.53 with  $F = 30$  cm,  $L = 1$  cm,  $x_{2\max} = y_{2\max} =$

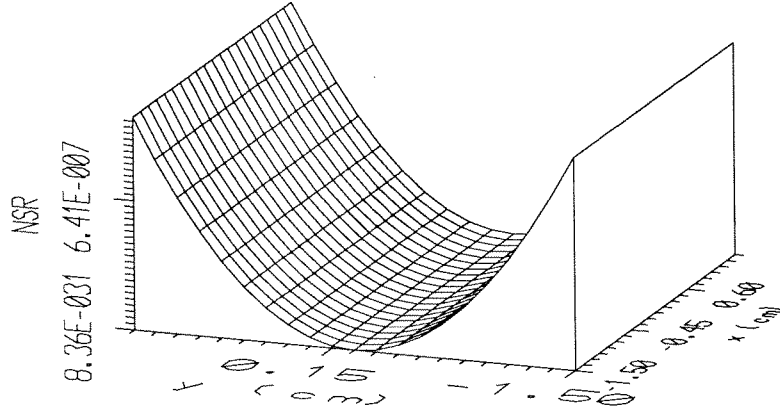


Figure 2.21: NSR vs position on the output plane for rectangular media.

1.5 cm,  $\lambda = 500$  nm,  $N = 201$  and is shown in Figure 2.21. The  $y_2 = -1.5$  side of the reconstruction has slightly more noise than the other side of the image. This is because the gratings on this side of the image are closer to the other holograms' gratings in k-space. The  $i = -100$  hologram has an output plane that is similar to Figure 2.21 but the largest NSR is at positive  $y_2$ . Figure 2.22 shows the SNR at the worst location on the output plane for 4001 holograms vs  $i$  (position in rotation). This is similar to the result given for angle multiplexing where the center holograms – zero rotation – are better than the holograms at large rotation (large angle). Figure 2.23 shows the minimum SNR (worst case in both position on output plane and in hologram number  $i$ ) vs the total number of holograms stored. This follows exactly the curve for angle multiplexing given in reference [170]. Therefore, for both rotation and angle multiplexing in rectangular media, SNR decreases as  $1/N$ .

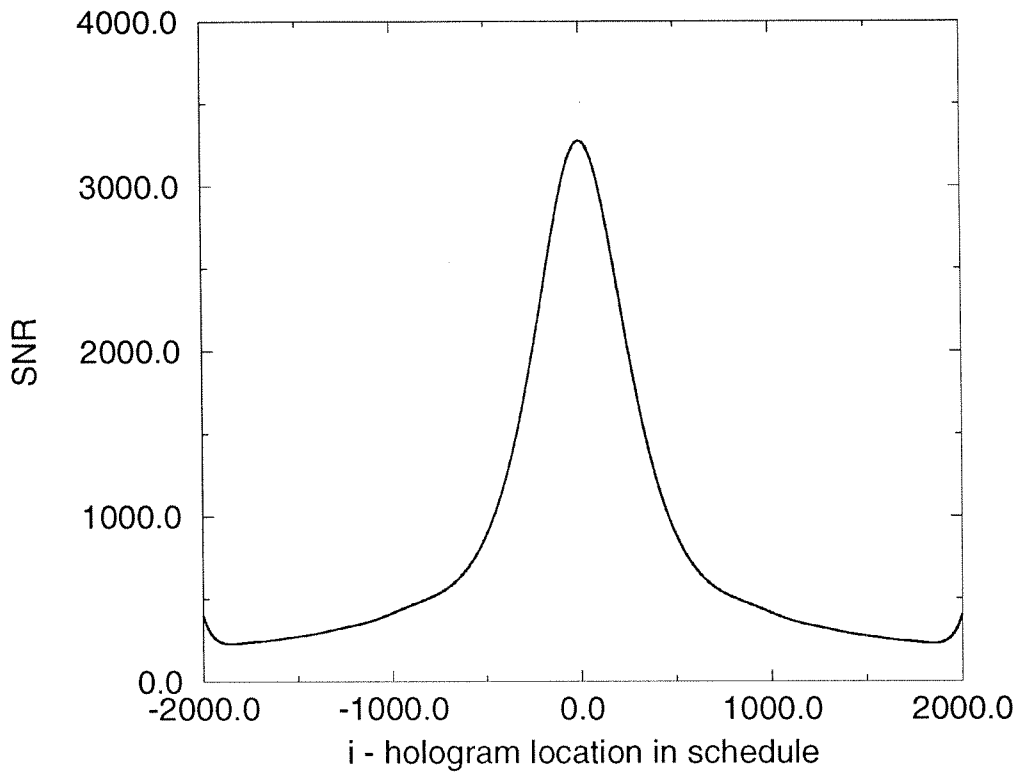


Figure 2.22: SNR vs hologram position in recording schedule.

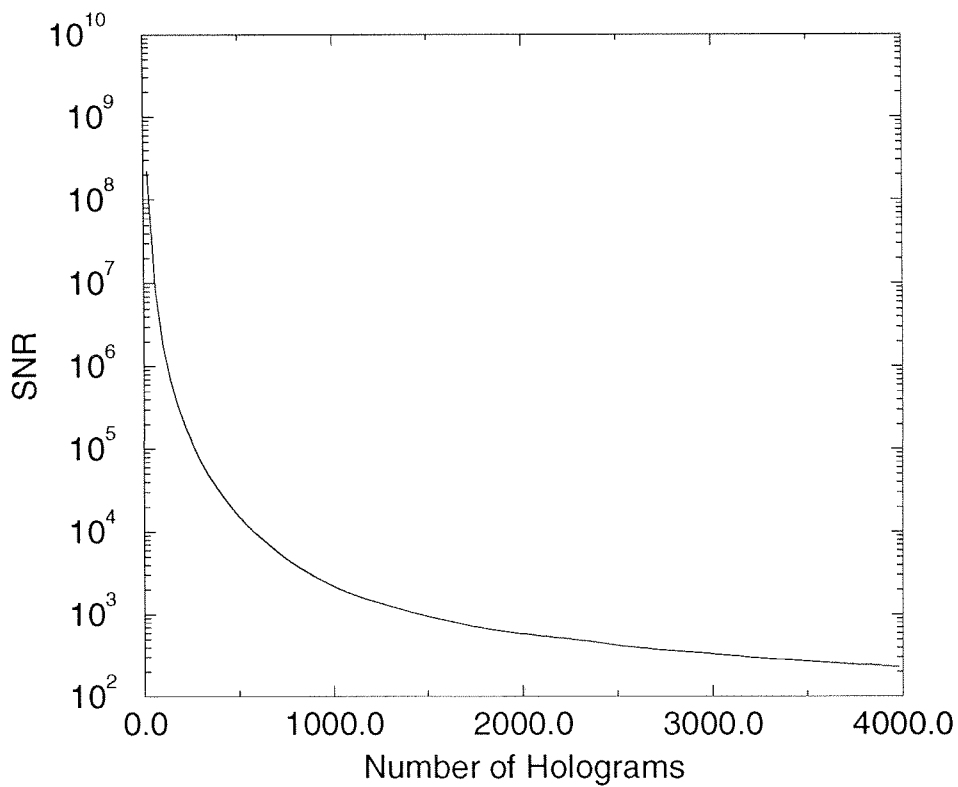


Figure 2.23: SNR vs the number of holograms stored for rectangular media.

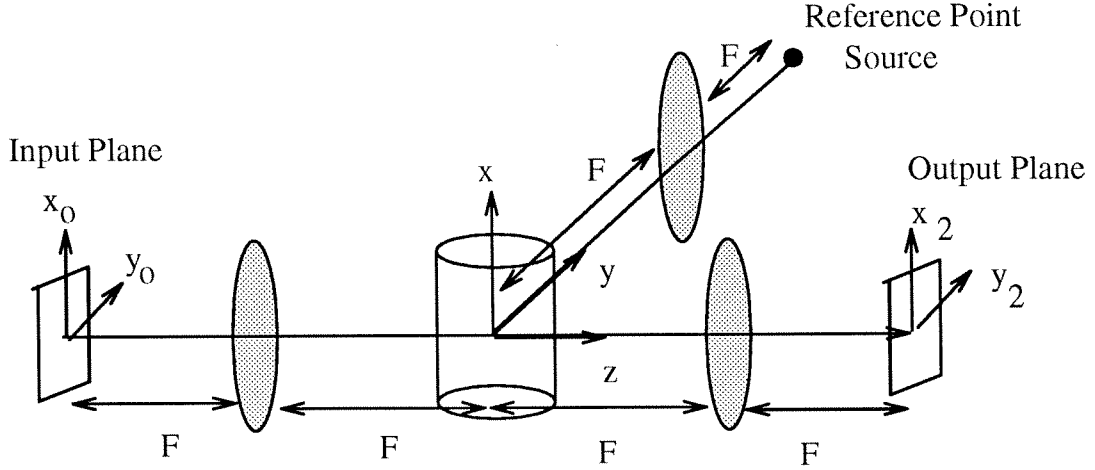


Figure 2.24: Recording and readout geometry for rotation multiplexing in a cylinder.

### 2.5.4 SNR Calculation for Rotation Multiplexing in a Cylinder

Fourier-transform holograms are recorded in a cylinder using the same setup as in Figure 2.24. The cylinder is situated such that it traces out a circle in the  $y - z$  plane with radius  $r_o$ , and holograms are multiplexed by rotating the cylinder around the  $x$  axis. Ignoring refraction at the surface (the cylinder is surrounded by an index matching material, for instance), Equations 2.44 – 2.48 are identical except that the integral in relation 2.47 is now over the volume of the cylinder. Making use of the assumption that the height of the medium in the  $x$  direction is much larger than the spatial bandwidth of the images ( again letting the resulting sinc function go to a delta function in  $x$ ), but before integrating over the volume of the media, transformed into cylindrical coordinates, the electric field at the output plane can be written as

$$E(x_2, y_2) \approx \sum_m \int dy_o f_m(-x_2, y_o) \int_0^{r_o} dr \, 2r \int_0^\pi d\phi \, e^{-ir\sqrt{A^2+B^2}\cos\phi} \quad (2.54)$$

where  $A$  and  $B$  are given in relations 2.50 and 2.51. The resulting integrals can be solved using

$$\int_0^\pi e^{i\beta\cos x} \cos(nx) dx = i^n \pi J_n(\beta) \quad (2.55)$$

[187] and

$$\int_0^1 x^{v+1} J_v(ax) dx = a^{-1} J_{v+1}(a) \quad (2.56)$$

[188]. Using Gegenbauer's formula [189] to simplify gives the electric field as

$$E(x_2, y_2) \approx \sum_m \int dy_o f_m(-x_2, y_o) \sum_{k=0}^{\infty} \frac{J_1(r_o A)}{r_o A} \frac{J_1(r_o B)}{r_o B} (1+k) C_k^1(0) \quad (2.57)$$

where the  $C_k^v()$ 's are the Gegenbauer polynomials. Using  $J_1(x)/x$  as a  $\delta$  function, as we did earlier with the sinc functions, leads to the requirement that  $r_o A = 0$ . Using this requirement to solve for  $y_o$  yields

$$y_o = -y_2 - \frac{F\theta_m^2}{2}, \quad (2.58)$$

which is similar to the expression for  $y_o$  found before. Substituting this result into our expression for B, and using the properties of  $C_k^1(0)J_k(r_o A)/(r_o A)$  as  $r_o A$  approaches zero, results in a simplified expression for the electric field. The noise-to-signal ratio is again calculated by dividing the total average noise power by the magnitude squared of the signal. With the same assumptions that each pixel of the stored images is an independent random variable taking the values minus one and one with equal probability, the NSR is given by

$$\text{NSR} = 4 \sum_{m \neq 0} \frac{\left[ J_1\left(\frac{\pi D}{\lambda} \theta_m + \frac{\pi D y_2}{2\lambda F} \theta_m^2\right) \right]^2}{\left( \frac{\pi D}{\lambda} \theta_m + \frac{\pi D y_2}{2\lambda F} \theta_m^2 \right)^2}, \quad (2.59)$$

where  $D$  is the diameter of the cylinder.

### SNR Results for Cylinder

In order to minimize the NSR, the spacing between the holograms in rotation is chosen to place the center pixel ( $y_2 = 0$ ) of each adjacent hologram at the zeroes of the  $J_1(x)/x$  function. Locating the zeroes was done by numerically evaluating the



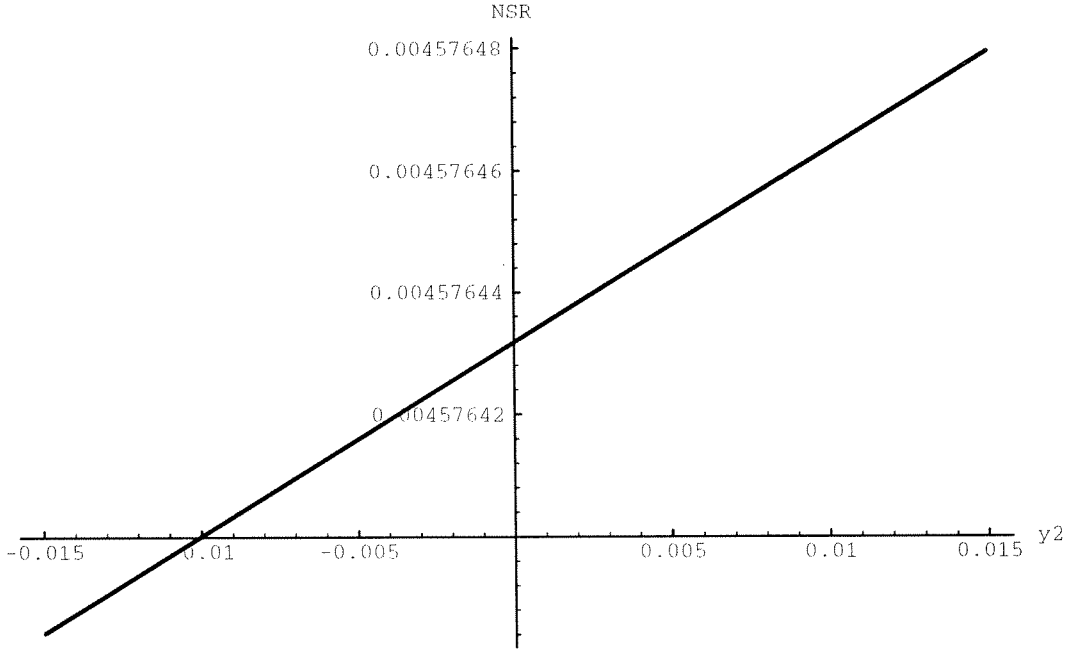


Figure 2.25: NSR vs position on the output plane for cylindrical media for the hologram stored at the furthest clockwise position.

$J_1(x)/x$  function. Compared to the rectangular media spacing ( $\frac{\lambda}{L}$ ) the first zero is approximately 1.22 times larger, the second 2.23 times larger, and the third zero has 3.23 times larger rotation between holograms. This means that the holograms cannot be multiplexed as close together when using cylindrical media.

Using the same parameters as for the rectangular media ( $D = L = 1\text{cm}$ ), Equation 2.59 can be numerically evaluated. Figure 2.25 shows the NSR versus  $y_2$  (position on the output plane) for the hologram that is stored at the furthest clockwise position with a total of 21 holograms stored. Again the positive  $y_2$  side is noisier since its grating vectors are closer to the other holograms. The NSR is plotted against position on the output plane for the center hologram in Figure 2.26. In this hologram, the noise is symmetric since there are an equal number of holograms on each side. The worst noise for all holograms occurs at one of the edges of the output plane – the one which is closest to the largest number of holograms. Taking the noise at the worse

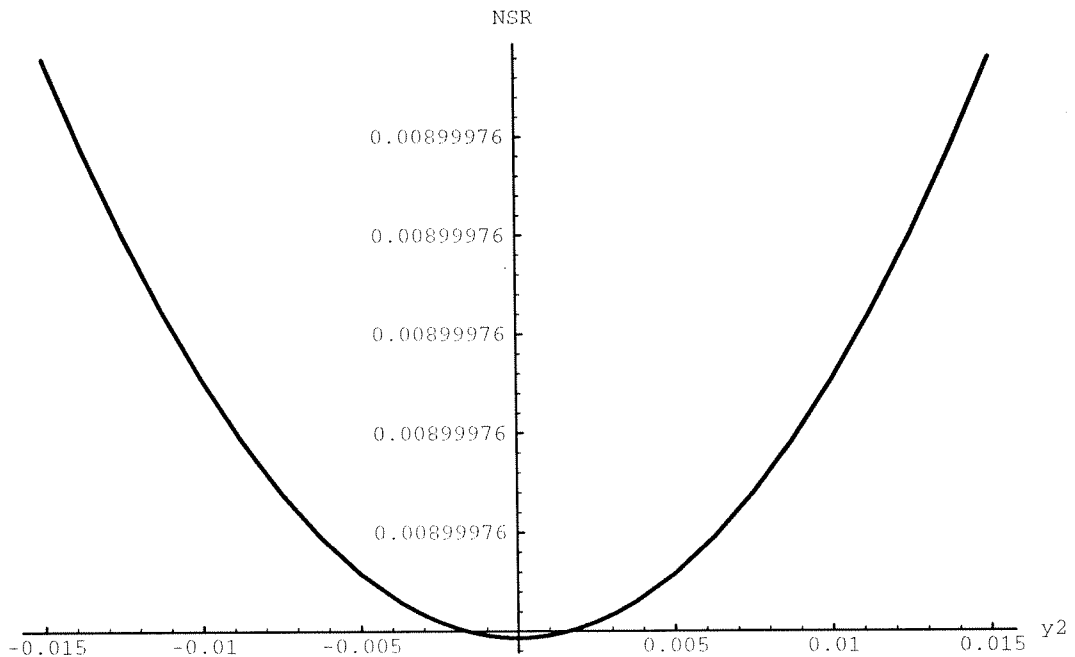


Figure 2.26: NSR vs position on the output plane for cylindrical media for the center hologram.

point in the output plane, we show, in Figure 2.27, SNR as a function of hologram position for 100 holograms. The worse SNR is always the center hologram since it borders the largest number of other holograms. Figure 2.28 shows the worst case SNR versus the total number of holograms stored for the case of multiplexing at the first, second, or third zero of the  $J_1(x)/x$  function. Notice that unlike rotation multiplexing or angle multiplexing in a rectangular media, the SNR saturates quickly to a fixed level in a cylindrical media. This is partly due to the fact that as the cylinder is rotated the interaction distance does not change; therefore, the width between nulls of the selectivity function ( $J_1(x)/x$ ) remains identical for all holograms. In addition, rotation tends to further separate gratings from different holograms unlike the case for angle multiplexing. In addition, the noise saturates quickly since  $J_1(x)/x$  falls off much faster than the sinc function.

Figure 2.28 shows the SNR for scheduling at the first, second, and third zeroes.

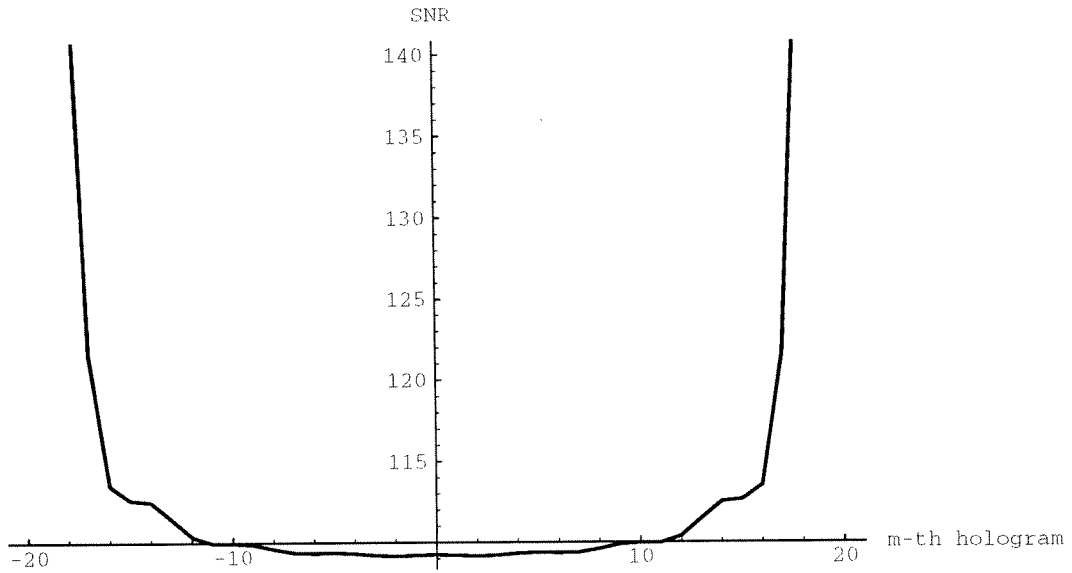


Figure 2.27: SNR vs hologram position in recording schedule for cylindrical media.

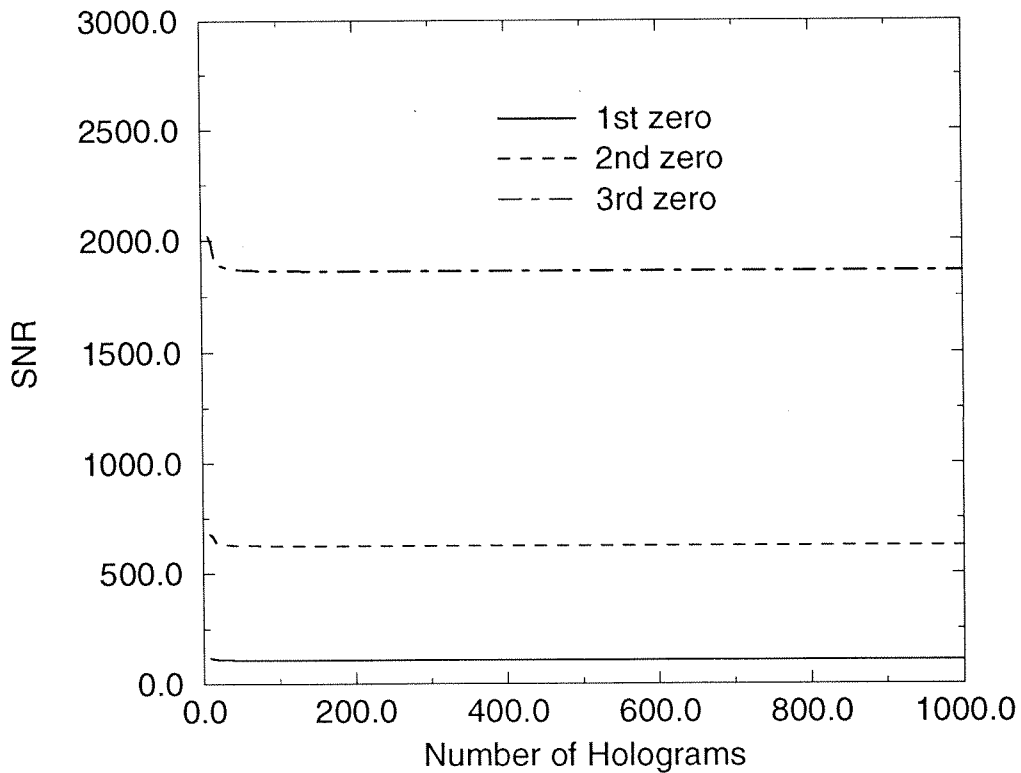


Figure 2.28: SNR vs the number of holograms stored for cylindrical media.

Notice that the SNR can be improved quickly in this way. For comparison, the SNR for the second zero results in saturation at about 640 for the cylindrical media compared to 800 for first zero wavelength multiplexing using the otherwise identical parameters in rectangular media. The SNR for 1,000 holograms for either angle or rotation multiplexing in a rectangular medium is about 2,000 which is comparable to placing the adjacent holograms in the third zero for the cylindrical media. Thus with cylindrical media the holograms cannot be stored as close together as with rectangular media, but since it is possible to access more grating space it is still possible to get comparable SNR. Since the increase in the number of possible angles is around 10 times the number available using angle multiplexing, the factors of 2 – 3 lost because of the wider selectivity function is not significant in terms of storage capacity. Thus different geometries such as cylindrical media may prove to have interesting possibilities for storage.

## 2.6 Image Plane Angle Multiplexed Holograms

### 2.6.1 Introduction

The cross-talk for angle multiplexed image plane holograms is derived. The SNR is calculated for a recording schedule that places the center frequency of each image at the null in angle of the adjacent hologram. This section contains two main results. First, we show that for image plane holograms the SNR is independent of the pixel location on the output plane. We also show that the worse case SNR for image plane is better than the worse case SNR for Fourier transformed holograms recorded with typical parameters.

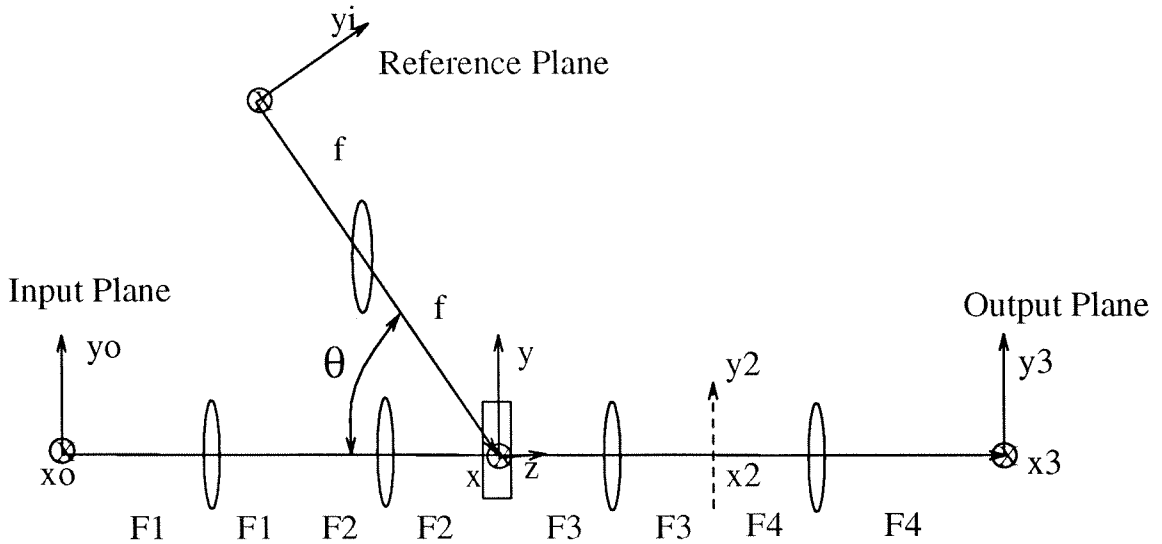


Figure 2.29: Recording and readout geometry for angle multiplexing image plane holograms.

### 2.6.2 Theory

Image plane holograms are multiplexed in a volume holographic medium using the setup shown in Figure 2.29. A plane wave  $R_m$  is generated at the material by collimating light from a point source at the reference plane. This interferes inside the holographic medium with  $S_m$  – the magnified image of the  $m$ -th object image. By labeling the holograms  $m = -M, -(M - 1), \dots, 0, \dots, M$ , the presence of these  $N = 2M + 1$  holograms modulates the permittivity of the material such that the change in the permittivity of the medium can be written as

$$\Delta\epsilon \approx \sum_{m=-M}^M R_m^* S_m + c.c. \quad (2.60)$$

where  $c.c.$  is the complex conjugate term. The reference wave is given by

$$R_m = e^{ik_m \cdot r}, \quad (2.61)$$

where  $k_m$  refers to the wave vector coming from the reference point source at  $y_m$ . In addition, using standard Fourier-optics analysis [182], the  $m$ -th signal beam can be

expressed as

$$S_m(r) \approx \iint dx_o dy_o \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} du dv f_m(x_o, y_o) e^{i\frac{2\pi}{\lambda_m} z} e^{-i\pi\lambda z(u^2+v^2)} e^{i2\pi u(x+M_1 x_o)} e^{i2\pi v(y+M_1 y_o)}. \quad (2.62)$$

In the expression above,  $f_m(x_o, y_o)$  is the  $m$ -th object image,  $x, y, z$  are the coordinates at the recording medium,  $M_1 = F2/F1$ , and  $u$  and  $v$  are spatial frequencies in the  $x$  and  $y$  directions respectively. By substituting Equations 2.61 and 2.62 into Equation 2.60,  $\Delta\epsilon$  can be written out explicitly.

### 2.6.3 Cross-talk Calculations

In readout, the recorded medium is illuminated with a reference plane wave from position  $y_i$  on the reference plane. Ideally, this would only reconstruct the hologram associated with that particular reference. In what follows, we will calculate the deviations of the reconstructed wave from the desired image. The amplitude of the diffracted plane wave components,  $E(\vec{k}_d)$ , with wave vectors  $\vec{k}_d$ , can be derived from standard scalar diffraction theory [183]. The following expression assumes that the Born and paraxial approximations are valid.

$$E(\vec{k}_d) \approx \int d\vec{r} e^{-i\vec{K}\cdot\vec{r}} \Delta\epsilon(\vec{r}), \quad (2.63)$$

where  $\vec{K} = \vec{k}_d - \vec{k}_i$ ,  $\vec{k}_i$  is the wave vector of the reference plane wave.

Lens with focal length  $F3$  maps each diffracted plane wave component with wave vector  $\vec{k}_d$  to a point  $x_2, y_2$  at the  $x_2, y_2$  plane given by:

$$\vec{k}_d = \left( \frac{2\pi x_2}{\lambda F_3}, \frac{2\pi y_2}{\lambda F_3}, \frac{2\pi}{\lambda} \left( 1 - \frac{x_2^2}{2F_3^2} - \frac{y_2^2}{2F_3^2} \right) \right), \quad (2.64)$$

assuming the paraxial approximation. Making use of the assumption that the transverse dimensions of the medium are much larger than the spatial bandwidth of the images, substituting relations 2.61, 2.60, 2.62, and 2.64 into Equation 2.63, and carrying out the integration over the volume of the medium and the spatial frequencies, the electric field at the  $x_2, y_2$  plane can be written as

$$E(x_2, y_2) \approx \sum_{m=-M}^{M-1} F_m \left( -\frac{M_1 x_2}{\lambda F_3} - \frac{M_1}{2\pi} \Delta K_{mix}, \left( -\frac{M_1 y_2}{\lambda F_3} - \frac{M_1}{2\pi} \Delta K_{miy} \right) \right) \times \text{sinc} \left( \frac{t}{2\pi} \left( \Delta K_{miz} + \frac{1}{F} (\Delta K_{mix} x_2 + \Delta K_{miy} y_2) + \frac{\lambda}{4\pi} (\Delta K_{mix}^2 + \Delta K_{miy}^2) \right) \right), \quad (2.65)$$

where  $\Delta \vec{K}_{mi} = \vec{k}_m - \vec{k}_i$  is the difference between the  $m$ -th reference wave vector  $\vec{k}_m$  and the illuminating beam's wave vector  $\vec{k}_i$ . Also  $\Delta K_{mi\alpha}$  is the component of  $\Delta \vec{K}_{mi}$  in the  $\alpha$  direction and  $F_m()$  is the Fourier transform of  $f_m()$ . The grating vector  $\Delta \vec{K}_{ij\alpha}$ 's are given by  $\Delta K_{ijx} = 0$ ,  $\Delta K_{ijy} = \frac{2\pi}{\lambda f} (y_i - y_m) \cos \theta + \frac{\pi^2}{\lambda f^2} (y_m^2 - y_i^2) \sin \theta$ , and  $\Delta K_{ijz} = \frac{2\pi}{\lambda f} (y_i - y_m) \sin \theta + \frac{\pi^2}{\lambda f^2} (y_i^2 - y_m^2) \cos \theta$ . Also, in Equation 2.65,  $t$  is the thickness of the material in the  $z$  direction. This is exactly the expression given for Fourier plane holograms.

Setting  $\theta = 90^\circ$  in order to maximize the storage capacity, and taking the Fourier transform of equation 2.65 results in an expression for  $E_m(x_3, y_3)$ . To estimate the noise-to-signal ratio (NSR) we divide the total average noise power by the magnitude squared of the signal. We assume that each pixel of the stored images is an independent random variable taking the values minus one and one with equal probability. In this case, the NSR is given by

$$\text{NSR} = \sum_{m \neq i} \text{sinc}^2 \left( \frac{t}{\lambda f} (y_i - y_m) + \frac{\pi t}{4\lambda f^4 M_1} (y_m^2 - y_i^2)^2 \right). \quad (2.66)$$

### 2.6.4 SNR Results

To minimize the noise the reference pixels ( $y_i$ 's) need to be spaced by  $\lambda f/t$  which is exactly the result for Fourier plane holograms. Notice that the NSR is independent of pixel position on the output plane. This means that all pixels in a hologram have the same SNR. Using the spacing described above, the SNR as a function of hologram number ( $i$ ) was calculated with  $N = 513$ ,  $F = 30$  cm,  $t = 1$  cm,  $M_1 = 1$ , and  $\lambda = 500\text{nm}$  and is shown in Figure 2.30. Notice that the holograms with the worst

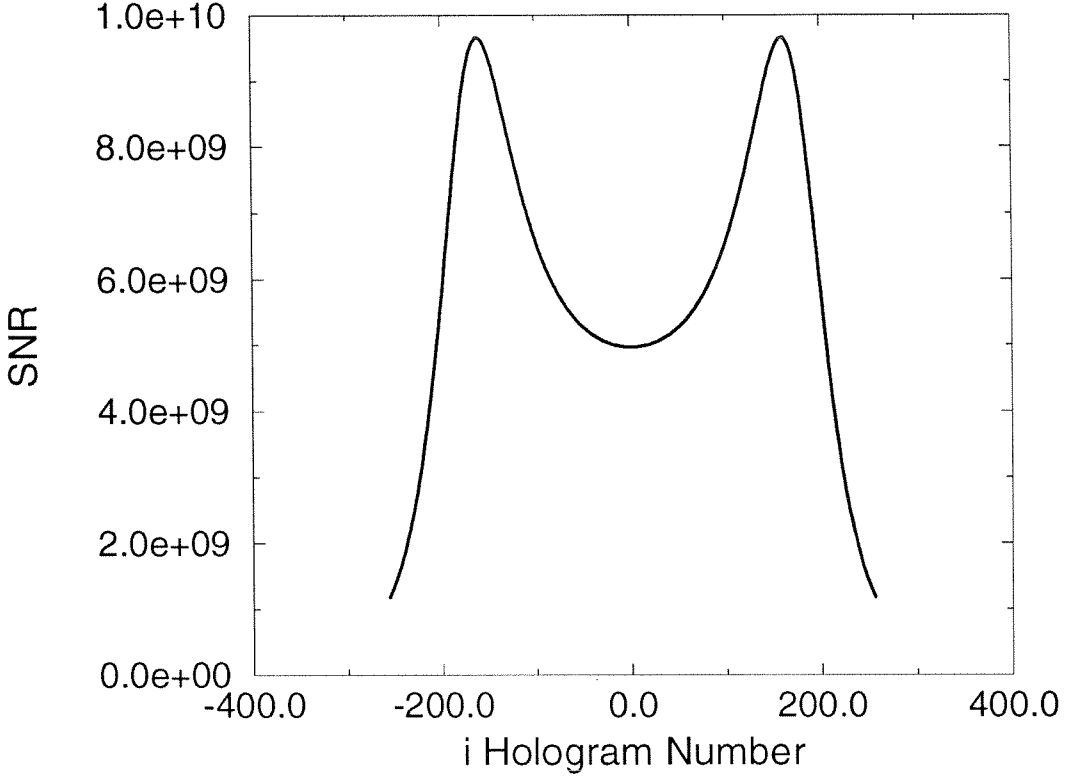


Figure 2.30: SNR vs hologram number  $-i$ .

SNR are the holograms that are furthest from  $90^\circ$  ( $i = M$ , or  $i = -M$ ). Also the holograms stored a little off  $90^\circ$  have better SNR than the hologram stored at  $90^\circ$ .

Figure 2.31 shows worst case SNR for angle multiplexed holograms that are stored in both image plane and Fourier plane formats as a function of the total number of holograms stored. Image plane holograms average out the variations in the noise which results in a higher SNR. Also this higher SNR is the same for every pixel in a given hologram while for Fourier transformed holograms the edges of the image in the  $y$  direction have much lower SNR than the middle of the image. Thus, by averaging out variations over the input plane, holograms stored in the image plane have higher SNR than holograms stored in the Fourier plane format.



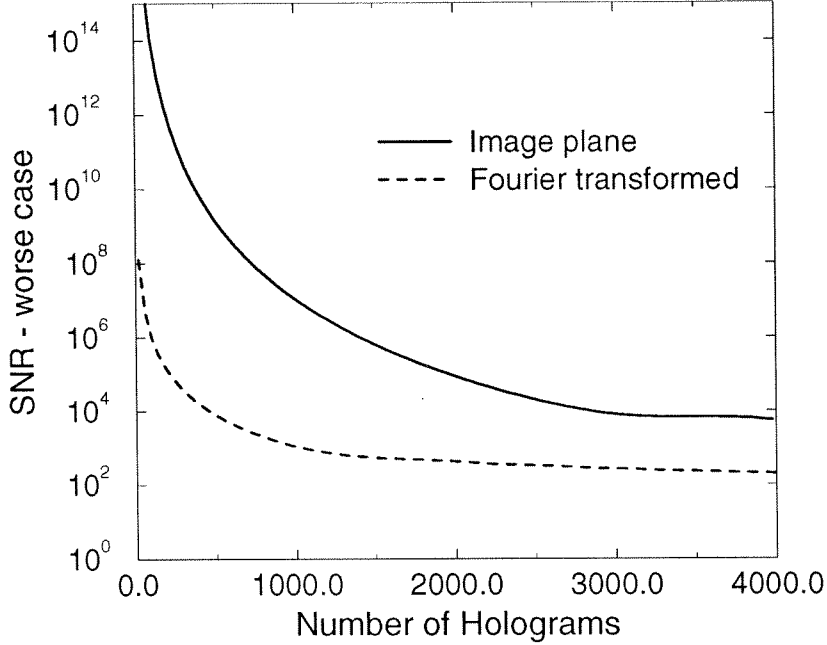


Figure 2.31: Worst case SNR for Fourier transformed and image plane holograms vs number of holograms stored.

## 2.7 Appendix 2.0: Derivation of Born Approximated Diffraction Equation

All the cross-talk calculations used an equation for diffraction from a volume assuming the both the Born and the paraxial approximations. This equation was derived for both isotropic and anisotropic material in Sidney Li's thesis [215]. The derivation for diffraction in isotropic medium directly follows that derivation and is given here for completeness.

The formula referred to is

$$E(\mathbf{r}) \approx e^{j\mathbf{k}_i \cdot \mathbf{r}} + \frac{k_0^2}{4\pi} \cdot e^{\frac{jkr}{r}} \int_V d\mathbf{r}' e^{-j\mathbf{K}' \cdot \mathbf{r}'} \Delta\epsilon(\mathbf{r}') , \quad (2.67)$$

where  $\Delta\epsilon$  is the change in the dielectric constant (not the permittivity),  $\mathbf{k}_i$  is the wave vector of the incident plane wave,  $\mathbf{k}'_d$  is the diffracted plane wave, and  $\mathbf{K}' = \mathbf{k}'_d - \mathbf{k}_i$ .

We begin the derivation with the Born's approximation formula

$$\mathbf{D}(\mathbf{r}) = \mathbf{D}^{(0)}(\mathbf{r}) + \frac{1}{4\pi} \int_V d\mathbf{r}' \frac{e^{jk|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} \nabla \times \nabla \times [\Delta\epsilon \mathbf{D}^{(0)}(\mathbf{r}')] , \quad (2.68)$$

where  $\mathbf{D}^{(0)}$  is the incident wave. If we assume the region of integration  $V$  is finite, take the origin somewhere inside  $V$ , and take the observation point far away from  $V$ , then we can make the approximation

$$\begin{aligned} |\mathbf{r} - \mathbf{r}'|^2 &= r^2 + r'^2 - 2\mathbf{r} \cdot \mathbf{r}', \\ |\mathbf{r} - \mathbf{r}'| &\approx r - 2 \left( \frac{\mathbf{r}}{r} \right) \cdot \mathbf{r}', \end{aligned} \quad (2.69)$$

since  $r \gg r'$ . Thus we have

$$\frac{e^{jk|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} \approx \frac{e^{jkr}}{r} \cdot e^{-j\mathbf{k} \cdot \mathbf{r}'}, \quad (2.70)$$

where we define

$$\mathbf{k} = k \left( \frac{\mathbf{r}}{r} \right), \quad (2.71)$$

and where  $r = |\mathbf{r}|$ .

We can therefore approximate  $\mathbf{D}$  by

$$\mathbf{D}(\mathbf{r}) = \mathbf{D}^{(0)}(\mathbf{r}) + \frac{1}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j\mathbf{k} \cdot \mathbf{r}'} \nabla \times \nabla \times [\Delta\epsilon \mathbf{D}^{(0)}(\mathbf{r}')] . \quad (2.72)$$

Next we assume that  $\Delta\epsilon$  is of the form

$$\Delta\epsilon(\mathbf{r}') = \Delta\epsilon_a e^{-j\mathbf{K} \cdot \mathbf{r}'}, \quad (2.73)$$

and  $\mathbf{D}^{(0)}$  is of the form

$$\mathbf{D}^{(0)}(\mathbf{r}') = D_0 e^{j\mathbf{k}_i \cdot \mathbf{r}'} \mathbf{e}_i, \quad (2.74)$$

where  $\mathbf{e}_i$  is the unit vector that gives the polarization of the incident wave  $\mathbf{D}^{(0)}$ .

We then have

$$\Delta\epsilon \mathbf{D}^{(0)}(\mathbf{r}') = D_0 \Delta\epsilon_a e^{j\mathbf{k}_a \cdot \mathbf{r}'} \mathbf{e}_i, \quad (2.75)$$

where  $\mathbf{k}_d = \mathbf{k}_i - \mathbf{K}$  and

$$\begin{aligned}\nabla \times \nabla \times [\Delta\epsilon \mathbf{D}^{(0)}(\mathbf{r}')] &= -D_0[\mathbf{k}_d \times (\mathbf{k}_d \times \mathbf{e}_i)] \left\{ \Delta\epsilon_a e^{j\mathbf{k}_d \cdot \mathbf{r}'} \right\}, \\ &= -D_0[\mathbf{k}_d \times (\mathbf{k}_d \times \mathbf{e}_i)] \Delta\epsilon(\mathbf{r}') e^{j\mathbf{k}_i \cdot \mathbf{r}'}\end{aligned}\quad (2.76)$$

assuming that  $\Delta\epsilon_a$  is a scalar. This then allows us to write  $\mathbf{D}$  as

$$\begin{aligned}\mathbf{D}(\mathbf{r}) &= \mathbf{D}^{(0)}(\mathbf{r}) - \frac{\mathbf{k}_d \times (\mathbf{k}_d \times \mathbf{e}_i)}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j(\mathbf{k}-\mathbf{k}_i) \cdot \mathbf{r}'} \Delta\epsilon(\mathbf{r}') \\ &= \mathbf{D}^{(0)}(\mathbf{r}) - \mathbf{u} \sin(\mathbf{e}_i, \mathbf{k}_d) \cdot \frac{k_0^2}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j(\mathbf{k}-\mathbf{k}_i) \cdot \mathbf{r}'} \Delta\epsilon(\mathbf{r}'),\end{aligned}\quad (2.77)$$

where  $\mathbf{u}$  is the unit vector in the direction of  $\mathbf{k}_d \times (\mathbf{k}_d \times \mathbf{e}_i)$ , and  $\sin(\mathbf{e}_i, \mathbf{k}_d)$  is the sine of the angle between  $\mathbf{e}_i$  and  $\mathbf{k}_d$  which is in general not equal to one. Note that  $\mathbf{u}$  is perpendicular to  $\mathbf{k}_d$ . Here we have used the fact that

$$|\mathbf{k}_d \times (\mathbf{k}_d \times \mathbf{e}_i)| = k_0^2 \sin(\mathbf{e}_i, \mathbf{k}_d). \quad (2.78)$$

If we take  $\mathbf{k}$  to be  $\mathbf{k}'_d$ , then the result in Eq. 2.77 agrees with the result in Eq. 2.67. For a general  $\Delta\epsilon$ , we assume that it can be written as a sum of sinusoidal gratings of the form

$$\Delta\epsilon(\mathbf{r}') = \sum_{\alpha} \Delta\epsilon_{\alpha} e^{-j\mathbf{K}_{\alpha} \cdot \mathbf{r}'}. \quad (2.79)$$

In this case we have

$$\begin{aligned}\mathbf{D}(\mathbf{r}) &= \mathbf{D}^{(0)}(\mathbf{r}) - \frac{1}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j\mathbf{k} \cdot \mathbf{r}'} \sum_{\alpha} \mathbf{k}_{\alpha} \times (\mathbf{k}_{\alpha} \times \mathbf{e}_i) \Delta\epsilon_{\alpha} e^{-j\mathbf{k}_{\alpha} \cdot \mathbf{r}'} \\ &= \mathbf{D}^{(0)}(\mathbf{r}) - \frac{k_0^2}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j(\mathbf{k}-\mathbf{k}_i) \cdot \mathbf{r}'} \sum_{\alpha} \sin(\mathbf{e}_i, \mathbf{k}_{\alpha}) \Delta\epsilon_{\alpha} e^{-j\mathbf{K}_{\alpha} \cdot \mathbf{r}'} \mathbf{u}_{\alpha},\end{aligned}\quad (2.80)$$

where

$$\mathbf{k}_{\alpha} = \mathbf{k}_i - \mathbf{K}_{\alpha}, \quad (2.81)$$

and  $\mathbf{u}_{\alpha}$  is the unit vector in the direction of  $\mathbf{k}_{\alpha} \times (\mathbf{k}_{\alpha} \times \mathbf{e}_i)$ . Assuming that all  $\mathbf{u}_{\alpha}$  are approximately in the same direction, we take

$$\sin(\mathbf{e}_i, \mathbf{k}_{\alpha}) \mathbf{u}_{\alpha} \approx \sin(\mathbf{e}_i, \mathbf{k}_{\alpha 0}) \mathbf{u}_{\alpha} \quad (2.82)$$

to be a constant and move it outside the summation. Eq. 2.80 then becomes

$$\begin{aligned} \mathbf{D}(\mathbf{r}) &= \mathbf{D}^{(0)}(\mathbf{r}) - \left\{ \frac{k_0^2}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j(\mathbf{k}-\mathbf{k}_i) \cdot \mathbf{r}'} \sum_{\alpha} \Delta\epsilon_{\alpha} e^{-j\mathbf{K}_{\alpha} \cdot \mathbf{r}'} \right\} \sin(\mathbf{e}_i, \mathbf{k}_{\alpha 0}) \mathbf{u}_{\alpha 0} \\ &= \mathbf{D}^{(0)}(\mathbf{r}) - \left\{ \frac{k_0^2}{4\pi} \cdot \frac{e^{jkr}}{r} \int_V d\mathbf{r}' e^{-j(\mathbf{k}-\mathbf{k}_i) \cdot \mathbf{r}'} \Delta\epsilon(\mathbf{r}') \right\} \sin(\mathbf{e}_i, \mathbf{k}_{\alpha 0}) \mathbf{u}_{\alpha 0}. \end{aligned} \quad (2.83)$$

Except for some constants, this agrees with the expression in Eq 2.67.

## 2.8 Appendix 2.1: Comments on Input Image Format

### 2.8.1 Random $\pm 1$ 's

The input data ( $a_n$ ) is always assumed to be random plus or minus 1 bits each with equal probability. To calculate the SNR the signal intensity is divided by the average noise intensity. The signal has magnitude one and the average noise intensity depends on the second moment of the random input signal. In the case of random +1 or -1's, the second moment of the random variable is

$$E\{a_n a_m\} = \delta_{mn} \quad (2.84)$$

where  $\delta_{mn}$  is a Kronecker delta function. This eliminates one of the sums in the average noise intensity calculation. Since the signal has no zeroes, the holograms consist of the maximum number of gratings.

### 2.8.2 Random $\pm 1$ 's or 0

Assuming that the signal ( $a_n$ ) is either  $|1|$  or 0 with equal probability and the 1 bits are equally likely to be -1 or +1, the second moment of  $a_n$  becomes

$$E\{a_n a_m\} = \frac{\delta_{mn}}{2} \quad (2.85)$$

where  $\delta_{mn}$  is a Kronecker delta function. Thus to convert the SNR calculated in the preceding sections into SNR for a  $\pm 1$  or 0 input just multiply by 2. This increase

in two is simply due to having 1/2 the number of recorded gratings. Thus the noise intensity due to diffraction off of these gratings is down by a factor of 2.

### 2.8.3 Random 1 or 0

For the case that the signal consists of random 1 or 0 bits each with equal probability, the second moment of the random variable is

$$E\{a_n a_m\} = \frac{\delta_{mn}}{2} + \frac{1}{4}(n \neq m). \quad (2.86)$$

Since the random variable has a mean the signal is equal the signal plus the noise. The noise is much smaller than one; therefore, the signal intensity is still approximately one. The average noise is now expressed as

$$\text{Ave. Noise} = \frac{1}{2} \sum_{i \neq j} \text{sinc}(i, j) + \frac{1}{4} \sum_{i \neq j, i' \neq j'} \text{sinc}(i, j) \text{sinc}(i', j'). \quad (2.87)$$

The second term in the expression can be numerically shown to be much smaller than the first term since it sums over positive and negative values. Thus to convert the SNR calculated in the preceding sections into SNR for a binary 1 or 0 input just multiply by 2. This increase in two is simply due to having half the number of recorded gratings. Thus the noise intensity due to diffraction off of these gratings is down by a factor of two.

## 2.9 Appendix 2.2: Finite SLM Contrast Ratio

### 2.9.1 Introduction

Given that the signal is supposed to be a 0 or a  $\pm 1$  with equal probability, what happens if the SLM has a finite contrast ratio? For instance, what happens to the cross-talk if the signal is  $f_m = \epsilon_2 a_m + \epsilon_1$ , where  $a_m$  is the ideal signal and the contrast ratio is defined as  $\epsilon_2^2/\epsilon_1^2$ . The signal is given by  $\epsilon_2^2$ . Using the fact that

$$E\{a_n a_m\} = \frac{\delta_{mn}}{2}$$

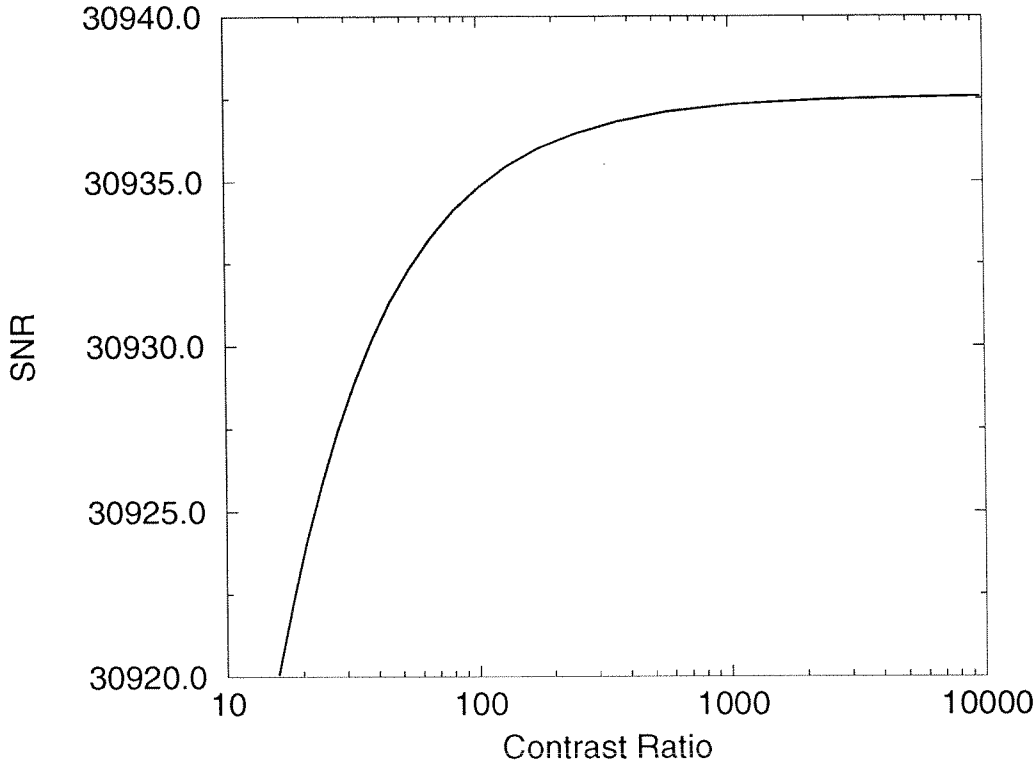


Figure 2.32: Worst case SNR for Fourier transformed holograms vs contrast ratio.

$$E\{a_m\} = 0, \quad (2.88)$$

the average noise intensity can be written as

$$\text{Ave. Noise} = \frac{1}{2} \sum_{i \neq j} \text{sinc}(i, j) + \epsilon_1^2 \sum_{i \neq j} \sum_{i' \neq j'} \text{sinc}(i, j) \text{sinc}(i', j'). \quad (2.89)$$

The SNR is the signal divided by the average noise intensity given in Eq. 2.89.

### 2.9.2 Contrast Ratio Results

Using the parameters for angle multiplexing given in section 2 of this chapter with  $N = 501$  holograms, the worst case SNR vs contrast ratio is shown in Figure 2.32. For Figure 2.32, we can conclude that as long as the contrast ratio of the SLM is reasonable, the contrast ratio has little effect on the cross-talk noise. For example, looking at Figure 2.32, holograms stored with a 16:1 contrast ratio have SNR that differs by only  $\sim 0.06\%$  from holograms stored with infinite contrast ratio.

## Chapter 3

# Comparison of Multiplexing Methods on the Basis of Cross-talk

### 3.1 Introduction

Multiple holograms can be recorded in a material by changing either the reference beam angle ( $\theta$  multiplexing), the recording wavelength ( $\lambda$  multiplexing), phase coding the reference beam ( $\phi$  multiplexing) or by rotating the media (rotation multiplexing). In the past, most of the work has concentrated on angle multiplexing, primarily because convenient frequency-tunable light sources and phase modulating SLMs have not been available. With the advent of semiconductor and solid-state tunable lasers and commercial SLMs (spatial light modulators), there has been renewed interest in wavelength and phase code multiplexing.

Recently, Rakuljic et al. [175] argued that  $\lambda$  multiplexed memories have lower cross-talk noise than  $\theta$  multiplexing. The cross-talk calculations in Chapter 2 indicate that the SNR is independent of the total number of holograms for  $\lambda$  multiplexing, whereas it drops as one over the number of holograms for  $\theta$  multiplexing, rotation multiplexing, and  $\phi$  multiplexing using Walsh-Hadamard codes. Therefore, for a large enough number of holograms,  $\lambda$  multiplexing has less cross-talk. In this section we

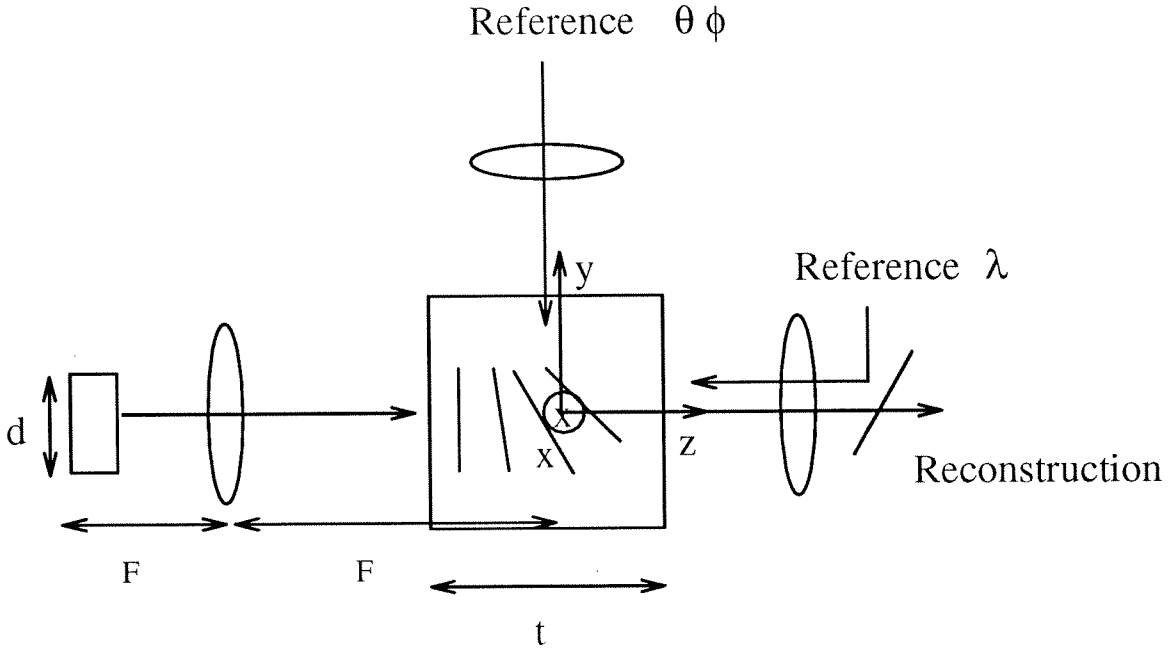


Figure 3.1: Geometry used for recording  $\theta$ ,  $\lambda$ , and  $\phi$  multiplexed Fourier transform holograms.

compare cross-talk results for the three methods and show that, in practice, if the recording geometry is optimal, the maximum number of holograms that can be stored with any of these methods is not limited by cross-talk [190, 191].

## 3.2 Summary of Cross-talk Results

### 3.2.1 Comparison of Methods

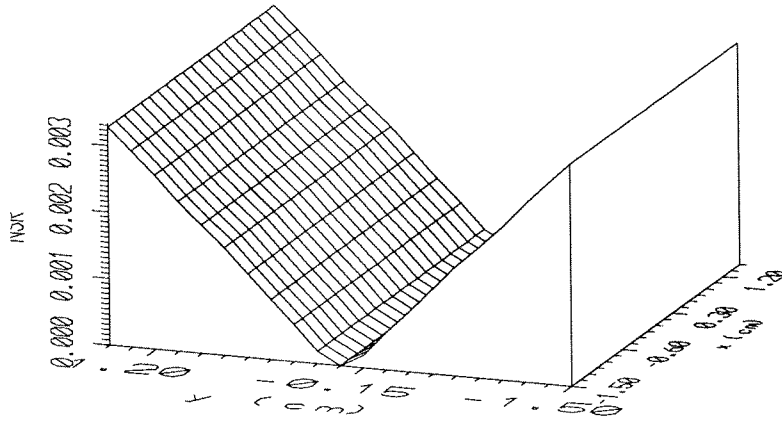
Gratings are recorded throughout the volume of a material using one of the three methods of multiplexing as shown in Figure 3.1. Since rotation multiplexing is similar to  $\theta$  multiplexing, only  $\theta$ ,  $\lambda$ , and  $\phi$  results will be summarized in this section. In order to maximize the storage capacity of each of the methods, a  $90^\circ$  geometry was assumed for  $\theta$  and  $\phi$  multiplexing and a pure reflection geometry for  $\lambda$  multiplexing. When a particular reference beam is incident on the medium, the amount of light



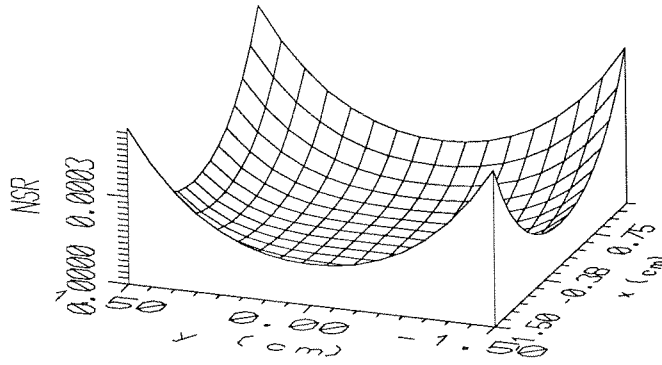
diffracted from the gratings can be calculated by assuming the Born and paraxial approximations. When the reference has the correct angle or wavelength, it is Bragg matched to one set of stored gratings, and the stored hologram is reconstructed. Cross-talk noise is due to energy diffracting off non-Bragg matched gratings. The cross-talk for  $\theta$ ,  $\lambda$ , and  $\phi$  multiplexing can be minimized by placing the center pixel of each Fourier transformed hologram at the null of the adjacent hologram's Bragg selectivity function. The calculations of cross-talk noise details are given in Chapter 2.

The noise-to-signal ratio (NSR) at the output plane for each method was numerically evaluated with (focal length)  $F = 30$  cm, (material thickness)  $t = 1$  cm, (output plane dimensions)  $x_{2max} = y_{2max} = 1.5$  cm, and  $\lambda = 500$  nm.  $\phi$  multiplexing was done using Walsh-Hadamard codes. Figure 3.2 shows the spatial distribution of the noise at the output plane. As mentioned above, the angular or wavelength separation between adjacent holograms was chosen to minimize the cross-talk of the center pixel for all three methods. Therefore, the center of the image has the lowest amount of noise while the edges of the images are noisier. For  $\phi$  and  $\theta$  multiplexing, the reference beams can only be scanned in one dimension ( $y$ ) since there is no Bragg selectivity in the second direction. For this reason, the noise is also a one dimension function at the output plane. For  $\lambda$  multiplexing, the recording geometry is radially symmetric around the z-axis which is reflected in the structure of the noise at the output plane.

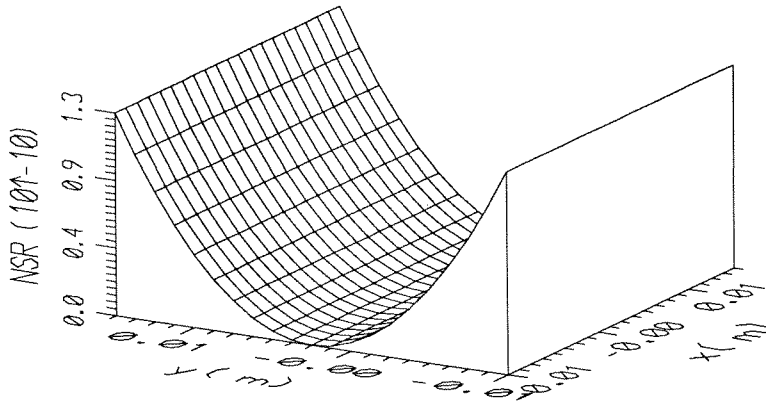
Selecting the noisiest pixel on the output plane and plotting this worst case signal-to-noise ratio (SNR) against hologram page number results in Figure 3.3. Here we show the behavior of the SNR for holograms stored at different angles, wavelengths, or phase codes. The plots in Figure 3.3 were made for 4,000, 10,000, and 63 holograms for  $\theta$ ,  $\lambda$ , and  $\phi$  multiplexing respectively. The angular separation was selected to minimize the cross-talk of the central hologram. Therefore, the SNR for  $\theta$  multiplexing is maximum at  $i = 0$  hologram. All  $\lambda$  multiplexed holograms have about the same



Angle



Wavelength



Phase Coded

Figure 3.2: NSR vs. position on the output plane for  $N = 4001, 401$ , and 15 holograms for  $\theta$ ,  $\lambda$ , and  $\phi$  multiplexing respectively.

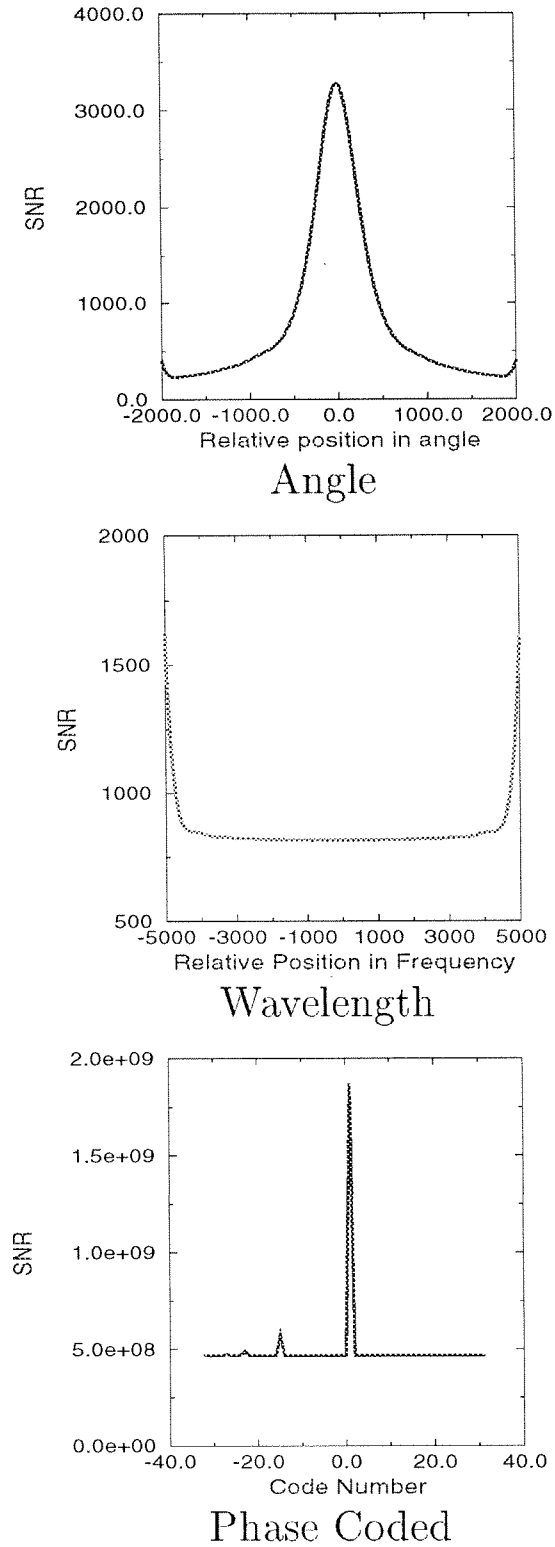


Figure 3.3: SNR vs. hologram page number in the schedule where zero represents the center angle/wavelength for  $\theta$ ,  $\lambda$  and  $\phi$  multiplexing.

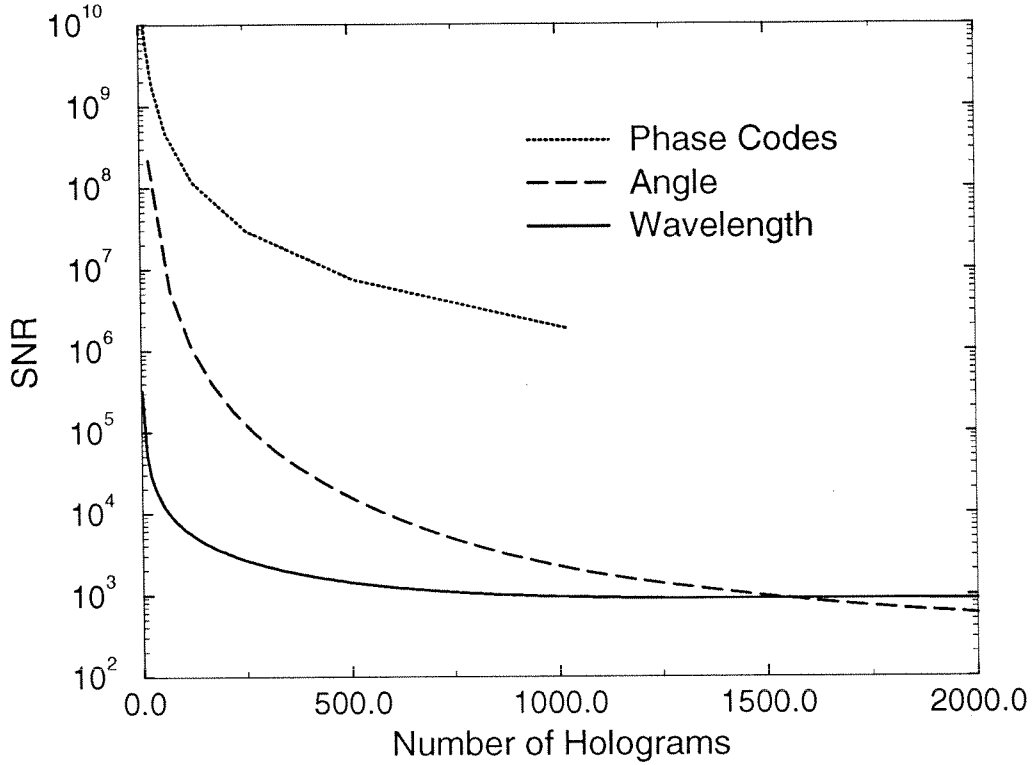


Figure 3.4: Worst case SNR vs. total number of holograms stored.

SNR except for the few holograms stored at the edge of the spectral range of the tunable source. This is due to the fact that cross-talk noise comes mostly from the neighboring holograms and that edge holograms only have neighbors on one side. For  $\phi$  multiplexing, most holograms average to the same SNR since they are recorded with all reference beam angles simultaneously. There is some variation in the SNR from code-word to code-word because they correlate differently with the non-Bragg matched stored holograms.

Figure 3.4 shows the SNR at the noisiest pixel in the output plane for the noisiest hologram stored, as a function of the total number of holograms stored. Notice that as we expected, the SNR decreases as the number of holograms for  $\phi$  and  $\theta$  multiplexing increases, while for  $\lambda$  multiplexing the SNR decreases rapidly to a saturation value. Also, if we take the SNR for  $\theta$  multiplexing at the edge of the output plane and

average over all holograms, we obtain an average SNR that is still more than two orders of magnitude less than  $\phi$  multiplexing for large  $N$ . Therefore, well chosen phase codes suppress, even on average, the cross-talk. The SNR was plotted for up to only 1,023 holograms for  $\phi$  multiplexing since the calculation is computationally intensive. Notice that the crossover point for the worst case SNR between  $\lambda$  and  $\theta$  is a little more than 1,500 holograms. Therefore, if the number of holograms stored is more than 1,500,  $\lambda$  multiplexing has lower cross-talk noise. For  $\phi$  multiplexing, the crossover point is much larger. The worst case SNR for storage of 1,000 holograms is  $\sim 1,000$  for  $\lambda$  multiplexing, compared to  $\sim 10^6$  for  $\phi$  multiplexing and  $\sim 3,000$  for  $\theta$  multiplexing. The cross-talk SNR is very high in all cases, which leads us to the conclusion that other system imperfections will determine the fidelity of the reconstruction in practice.

### 3.2.2 Image Plane vs Fourier Transform Holograms

Figure 3.5 shows worst case SNR for angle multiplexed holograms that are stored in both image plane and Fourier plane formats as a function of the total number of holograms stored. Image plane holograms average out the variations in the noise which results in a higher SNR for image plane holograms than for Fourier plane holograms. Also, this higher SNR is the same for every pixel in a given hologram while for Fourier transform holograms the edges of the image in the  $y$  direction (the direction of angle multiplexing) have much lower SNR than the middle of the image. Thus, the higher SNR and equal SNR for each pixel makes image plane holograms a desirable storage format if cross-talk noise is relevant.

### 3.2.3 Geometrically, When is Cross-talk Important

The conclusion that cross-talk is not the determining noise source was based on  $\theta$  and  $\phi$  multiplexing at  $90^\circ$ ,  $\lambda$  multiplexing in reflection geometry, and the optimum spacing

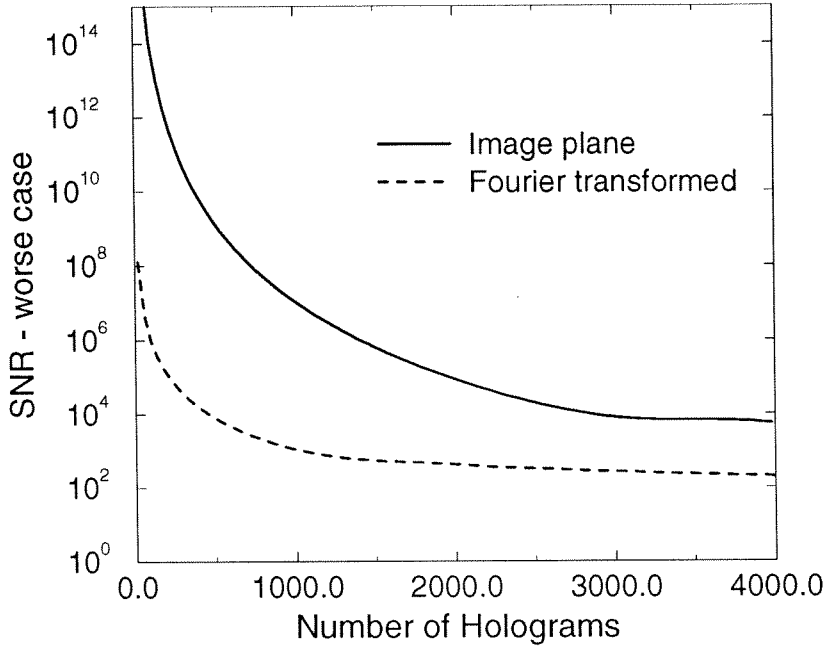


Figure 3.5: Worst case SNR for Fourier transformed and image plane holograms vs number of holograms stored.

of the holograms at the nulls of the Bragg selectivity. When  $\theta$  and  $\phi$  multiplexing is done at an intersection angle other than  $90^\circ$ , the SNR decreases rapidly. Figure 3.6 shows the worst case SNR for angle multiplexing 101 holograms as the intersection angle between the reference and signal beams is changed from  $10^\circ$  to  $90^\circ$ . Clearly, from a cross-talk point of view,  $90^\circ$  is the best geometry for  $\theta$  and  $\phi$  multiplexing. This can also be seen in the depth of the minimums of the sinc functions. For one plane wave component of the image using angle multiplexing, the nulls for  $90^\circ$  geometry are zeroes of the sinc function. For other geometries, the minima of the sinc functions are not zeroes due to other geometric terms arising (Figure 3.7). Figure 3.7 shows the diffracted power into the center pixel of the image stored (Fourier hologram) as the reference beam is scanned for different intersection angles. The SNR also decreases rapidly for geometries other than pure reflection for  $\lambda$  multiplexing and reference [171] gives a closed-form asymptotic expression for the SNR. Another

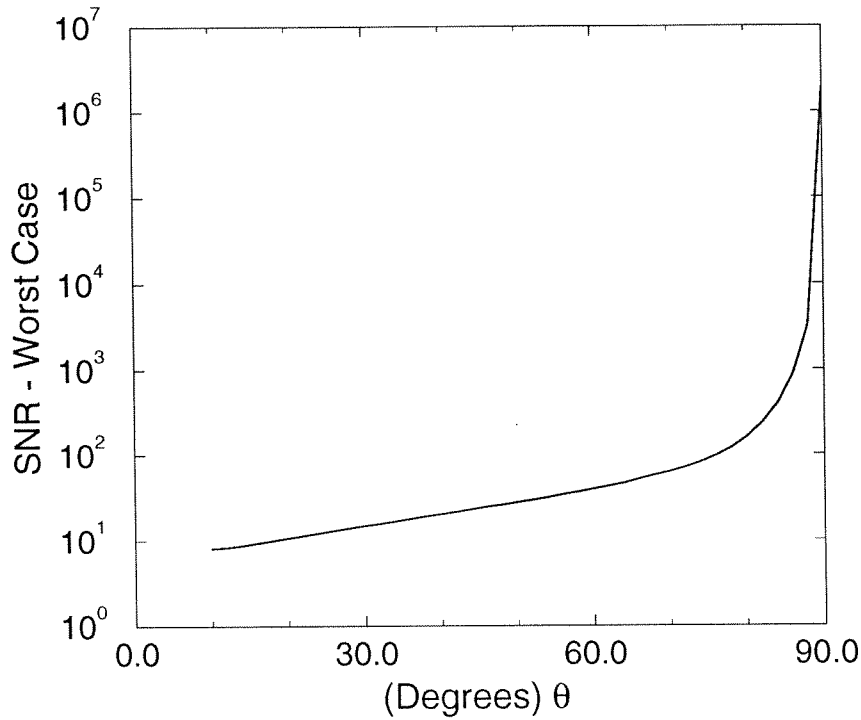


Figure 3.6: Worst case SNR for  $N = 101$  holograms as a function of angle between reference and image beams ( $\theta$ ) for **angle** multiplexing.

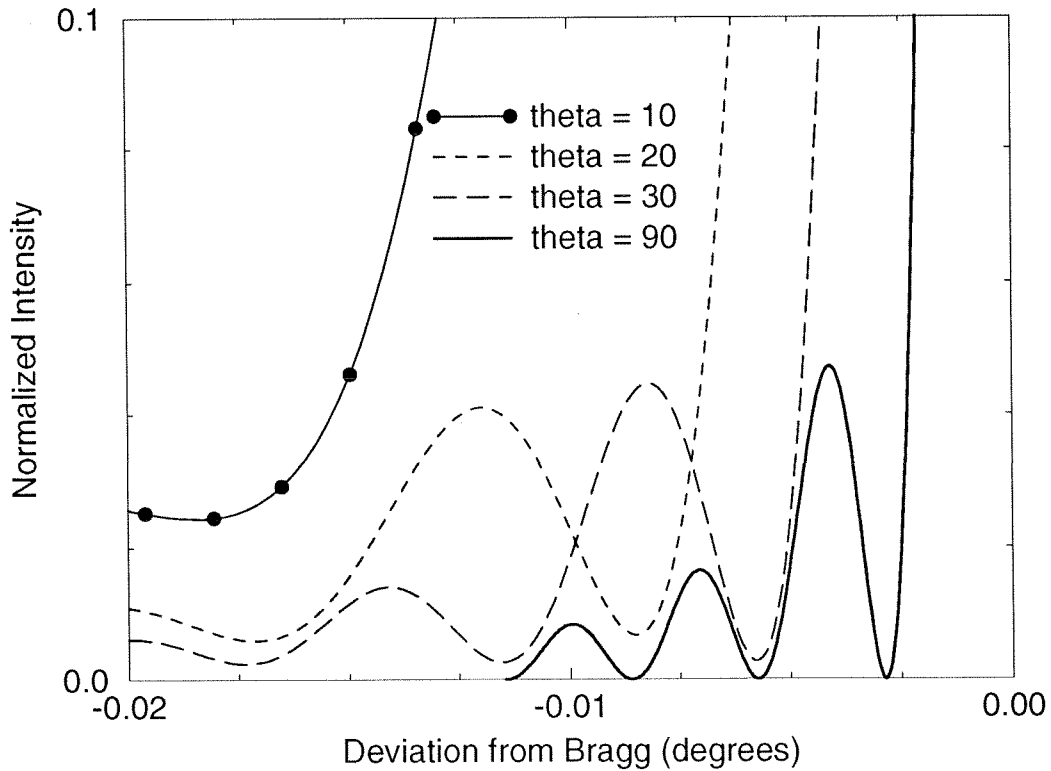


Figure 3.7: Normalized diffraction vs deviation from Bragg for  $\theta$  (intersection angle) equal to 10, 20, 30, and 90 degrees.

requirement for  $\lambda$  multiplexing is a tunable source with a sufficiently small linewidth to avoid further increases in cross-talk noise [192]. The other important criterion is that the holograms were stored at the minimum of the adjacent hologram's Bragg selectivity function. If the holograms are spaced so as they are not in the nulls, either intentionally or due to experimental error, the SNR will also decrease. Therefore, cross-talk will be irrelevant only when using the best geometry and spacing the holograms correctly in either angle or wavelength.

### 3.3 Understanding Cross-talk in Wave Vector and Grating Space

#### 3.3.1 k-space Introduction

To better understand the recording of holograms, we consider what happens in wave vector space ( $\mathbf{k}$ -space). Assuming the electric and magnetic fields are plane waves and therefore have the form of

$$\begin{aligned} \mathbf{E} e^{i(\vec{k} \cdot \vec{r} - \omega t)} \quad \text{and} \\ \mathbf{H} e^{i(\vec{k} \cdot \vec{r} - \omega t)}, \end{aligned} \quad (3.1)$$

the wave equation in an anisotropic medium can be written as

$$\vec{k} \times (\vec{k} \times \vec{E}) + \omega^2 \mu \epsilon \vec{E} = 0, \quad (3.2)$$

where  $\vec{k}$  is the wave vector,  $\vec{E}$  is the electric field,  $\omega$  is the angular frequency of the wave,  $\mu$  is the permeability tensor which is assumed to be a constant, and  $\epsilon$  is the dielectric tensor. Assuming that the coordinate system is chosen such that the dielectric tensor is a diagonal matrix, Equation 3.2 will only have nontrivial solutions if and only if

$$\det \begin{vmatrix} \omega^2 \mu \epsilon - k_y^2 - k_z^2 & k_x k_y & k_x k_z \\ k_y k_x & \omega^2 \mu \epsilon - k_x^2 - k_z^2 & k_y k_z \\ k_z k_x & k_z k_y & \omega^2 \mu \epsilon - k_x^2 - k_y^2 \end{vmatrix} = 0. \quad (3.3)$$



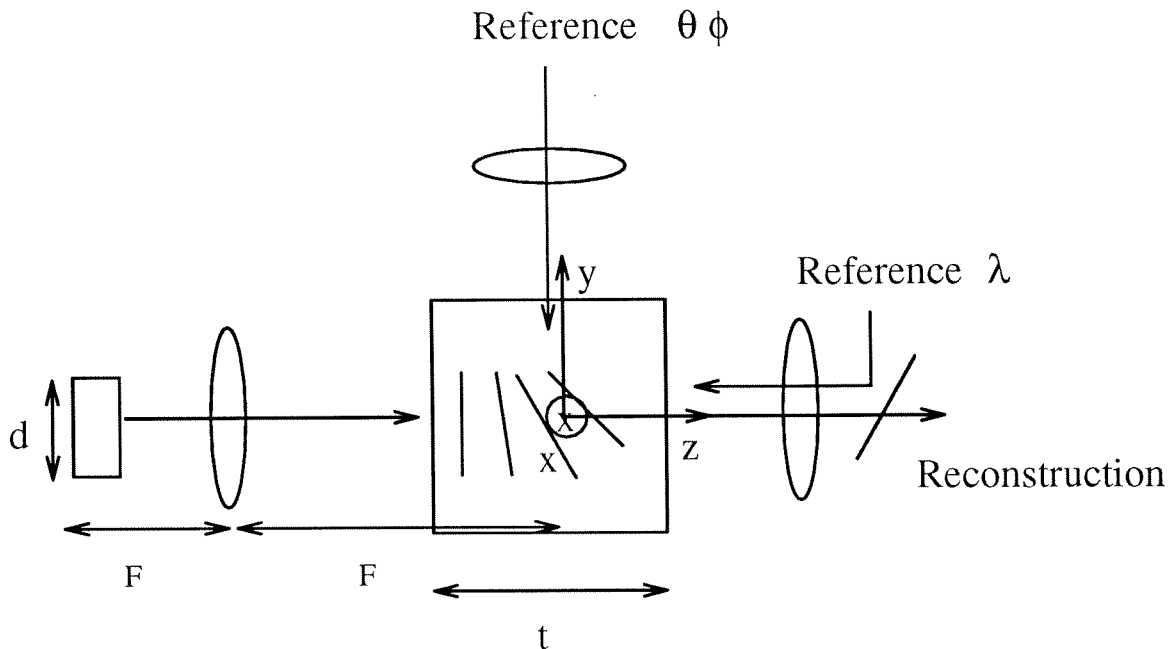


Figure 3.8: Geometry used for recording  $\theta$ , and  $\lambda$  multiplexed Fourier transform holograms.

Equation 3.3 defines a 3-dimensional  $\mathbf{k}$ -space of all possible wave vectors. In general, each polarization has a 3-dimensional ellipsoid that the tip of the vector is confined to lie on which is called the normal surface. In an isotropic medium, this constrains the vectors to lie on a sphere which can be written as

$$\frac{k_x^2}{n^2} + \frac{k_y^2}{n^2} + \frac{k_z^2}{n^2} = k_o^2 . \quad (3.4)$$

### 3.3.2 $\mathbf{k}$ -space for Angle Multiplexing

Angle multiplexing in the  $90^\circ$  geometry is shown in Figure 3.8. To construct the  $\mathbf{k}$ -space picture for the geometry in Figure 3.8, we center the normal surface on the material and represent each plane wave by a vector that starts in the middle and extends to the surface. The  $\mathbf{k}$ -space picture corresponding to the geometry in Figure 3.8 is shown in Figure 3.9. Since images are composed of a spectrum of plane waves, they are shown with several vectors. The plane wave reference beam records

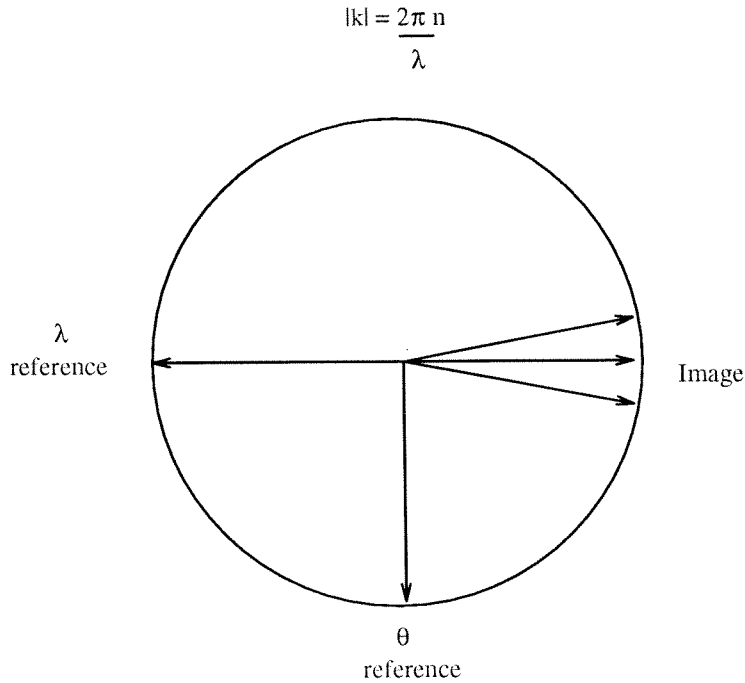


Figure 3.9: k-space diagram for  $\theta$  and  $\lambda$  multiplexing.

a grating with each plane wave component of the image.

A single grating is recorded by two plane waves interfering inside the material as shown in Figure 3.10 part (a). The grating vector is the thicker vector that connects the reference and signal plane waves and is defined as  $\vec{K} = \vec{k}_s - \vec{k}_r$ . To reconstruct the hologram, the reference beam is input, diffracts off the grating, and reconstructs the original signal, since  $\vec{K} + \vec{k}_r = \vec{k}_s$ . Since the material is of finite thickness, you can imagine that instead of a single grating the actual recording is a spectrum of gratings that have an amplitude distribution of a sinc function as shown in Figure 3.10 part (a). To angle multiplex another grating, the reference angle is changed and a new grating is recorded as in part (b). To find the cross-talk between the gratings, simply place the second grating vector at the beginning of the first vector and see where the resultant vector's sinc function crosses the normal surface. Notice that for two plane waves recorded at  $90^\circ$  it is possible to place this second resultant grating in the zero of the sinc function. Thus, it is possible to record two plane waves with zero

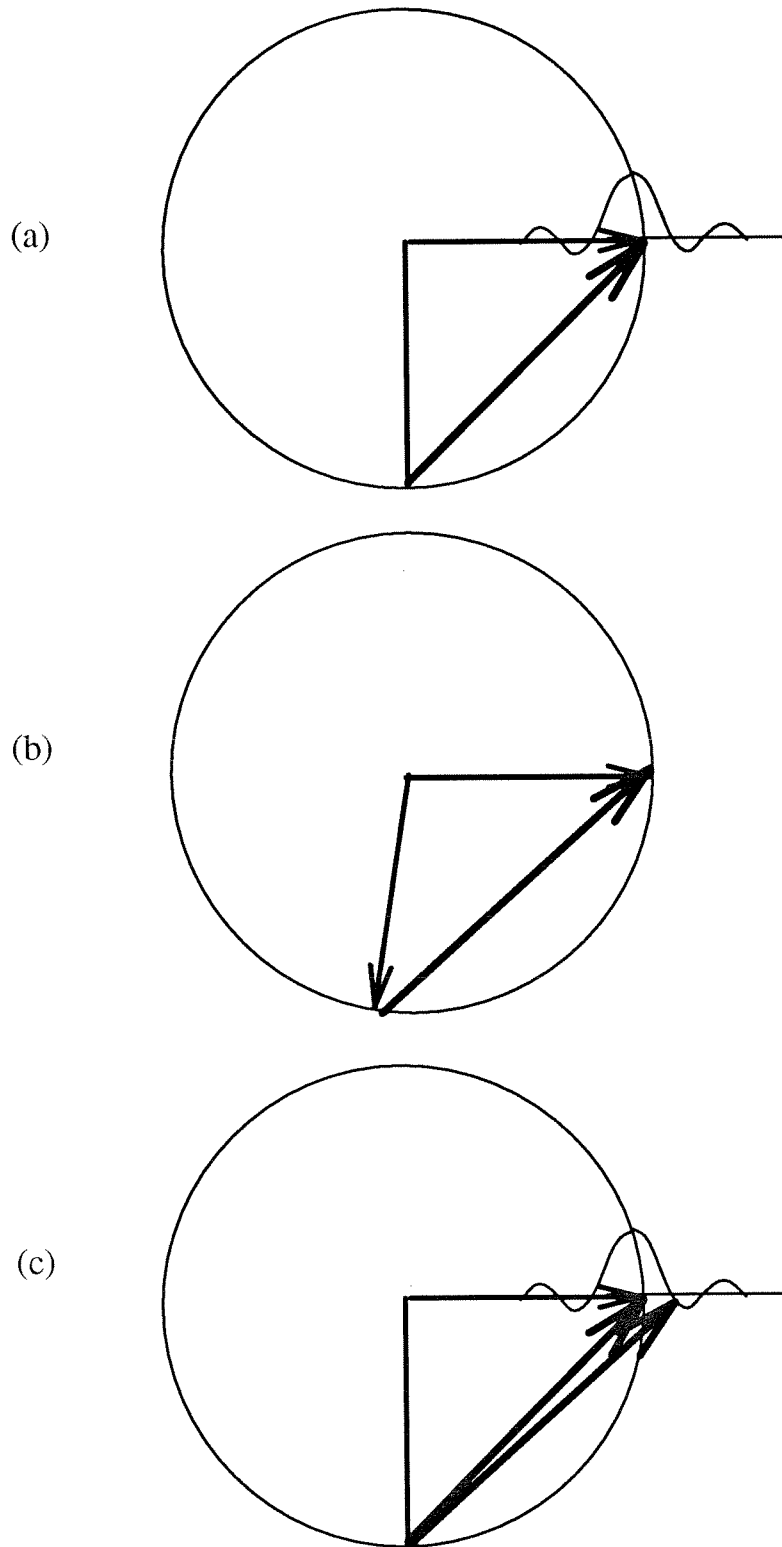


Figure 3.10:  $k$ -space diagram for reconstruction (a) single grating selectivity, (b) another grating angle multiplexing, (c) placing in nulls of first grating.

cross-talk. Unfortunately, since images are made up of many plane wave components it is **not** possible to make all of the components fall in the nulls. To better see what factors influence the cross-talk, it is advantageous to look at the grating relationships in “grating space.”

### 3.3.3 K-space for Angle Multiplexing

Grating space (or **K**-space) is simply the volume traced out by all the grating vectors after the ends are superimposed. This creates a space that is filled by the gratings and allows the distance between gratings to be easily visualized. Since only one particular value of separation in the  $z$  direction allows the sinc function to cross the normal surface at the minimum of the sinc function, any change in the distance between gratings results in cross-talk noise. Thus, **K**-space allows sources of cross-talk to be visualized.

**K**-space for angle multiplexing is made by first drawing the **k**-space pictures for the various holograms as shown in Figure 3.11. The spectrum of plane wave vectors that make up the image lie in the shaded regions of Figure 3.11. For simplicity, only the grating to the center plane wave of the image is drawn. To form **K**-space, the ends of the gratings are placed together. This results in Figure 3.12 which shows the **K**-space picture for three angle multiplexed holograms. Notice that as the angular spread of the images are increased the distance between the gratings is changed. If  $d/2$  is the maximum extent of the image plane and  $F$  is the focal length of the Fourier transforming lens, the maximum angular bandwidth of the image can be written as  $\Delta\theta_i = d/2F$ . Also notice that the outer holograms tend to rotate down on their normal surfaces relative to the center hologram. This rotation also changes the distance between the gratings and therefore introduces cross-talk. The maximum rotation is directly related to the maximum angular bandwidth of the reference beams. This spread of reference beams depends on the number of holograms and the spacing

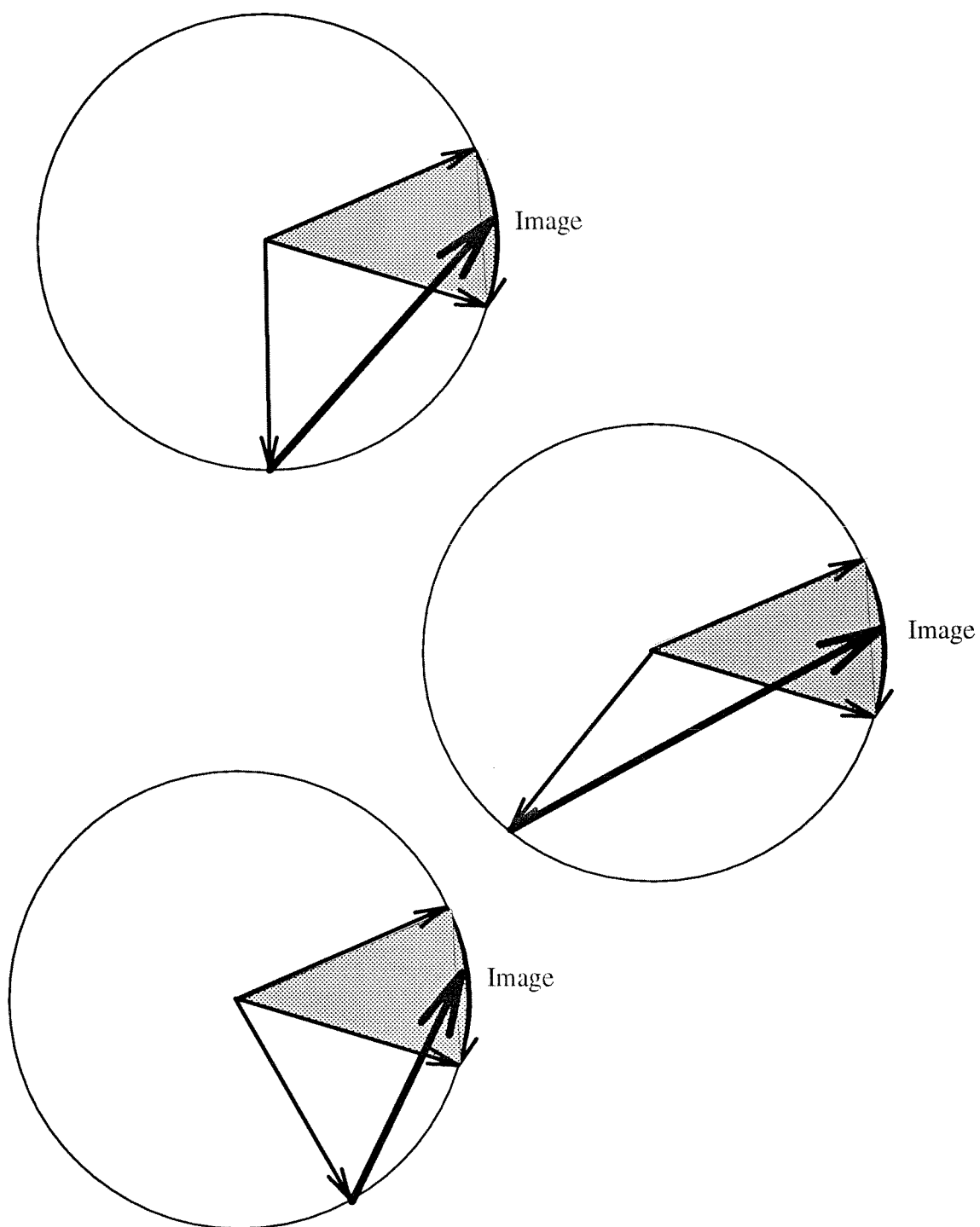
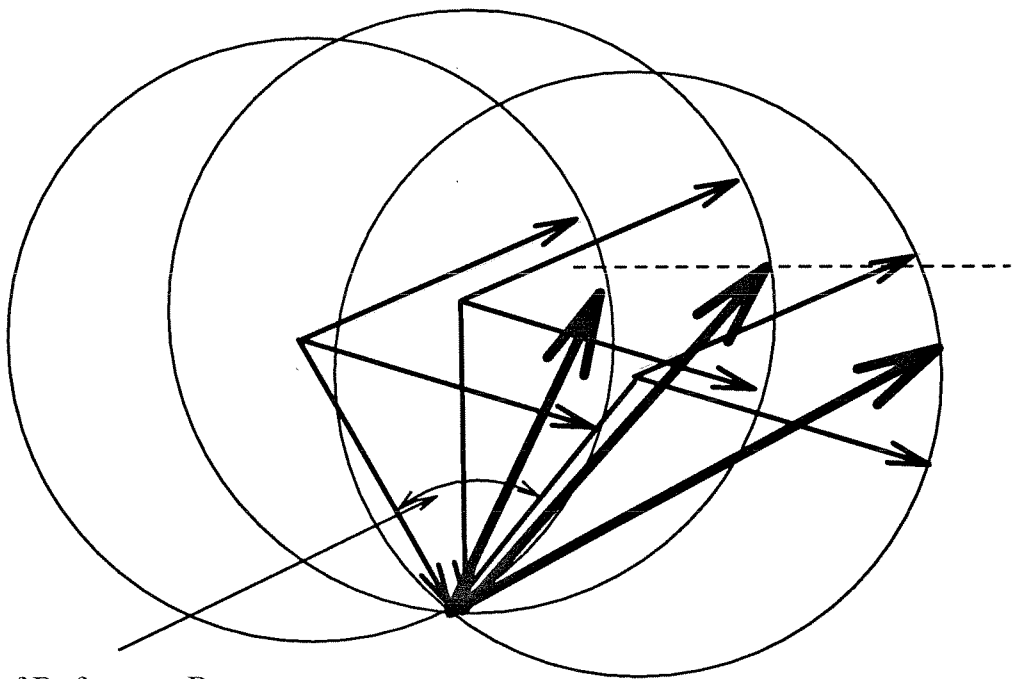


Figure 3.11: k-space diagrams for  $\theta$  multiplexing.



Sweep of Reference Beams

Figure 3.12: K-space diagrams for  $\theta$  multiplexing.

of the reference angles. In the  $90^\circ$  geometry, the bandwidth of the reference beams is equal to  $\Delta\theta_r = \lambda N_\theta / t$ , where  $N_\theta$  is the number of holograms stored, and  $t$  is the thickness of the recording medium in the  $z$  direction.

The closed form solution for SNR of  $\theta$  multiplexed holograms in the  $90^\circ$  geometry was given in Chapter 2 as

$$\text{SNR} = \frac{2tF}{\lambda d N_\theta} = \left(\frac{1}{\Delta\theta_i}\right)\left(\frac{1}{\Delta\theta_r}\right). \quad (3.5)$$

We see that the SNR for angle multiplexing goes as one over bandwidth of the reference beams and one over the bandwidth of the image beam.

Rotation and phase code have geometries similar to the geometry for angle multiplexing. Phase code multiplexing has exactly the same geometry except that the noise does not add up coherently. For rotation multiplexing, the grating space picture is shown in Figure 3.13. Notice that again the two factors are image bandwidth and rotation of image away from the center. Because of this, for small rotations, the SNR result for rotation multiplexing is approximately the same as for  $\theta$  multiplexing.

### 3.3.4 k-space for Wavelength Multiplexing

Wavelength multiplexing in the pure reflection geometry is shown in Figure 3.8. A single grating is recorded by two plane waves of a given wavelength interfering inside the material as shown in Figure 3.14 part (a). Again, the grating vector is the thicker vector that connects the reference and signal plane waves and is defined as  $\vec{K} = \vec{k}_s - \vec{k}_r$ . To reconstruct the hologram, the reference beam with the original wavelength is input, diffracts off the grating, and reconstructs the original signal, since  $\vec{K} + \vec{k}_r = \vec{k}_s$ .

Just like in  $\theta$  multiplexing, since the material is of finite thickness one can imagine that instead of a single grating the actual recording is a spectrum of gratings that have an amplitude distribution of a sinc function as shown in Figure 3.14 part (a).

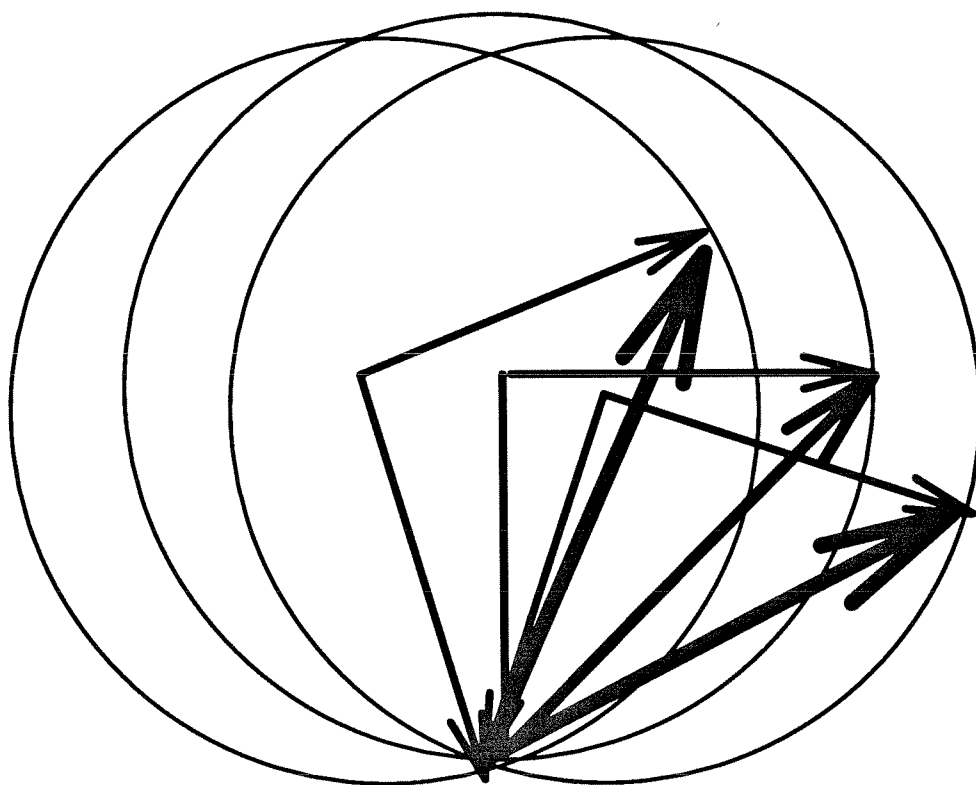


Figure 3.13: **K**-space diagrams for rotation multiplexing.



To  $\lambda$  multiplex another grating, both the reference and signal beam wavelengths are changed and a new grating is recorded on a different normal surface, since the radius of the normal surface depends in the wavelength of the light as in part (b). To find the cross-talk between the gratings, simply place the second grating vector at the tip of the first vector and see where the resultant vector's sinc function crosses the normal surface. Notice that for two plane waves it again is possible to place the resultant in the minimum of the sinc function. In other words, it is possible to record two plane waves with minimal cross-talk. Unfortunately, as in the angle multiplexing case, images are composed of many plane wave components and it is **not** possible to make all of these components fall in the minima of the sinc functions. To better see what factors influence the cross-talk for wavelength multiplexing, it is again advantageous to look at the grating relationships in grating space.

### 3.3.5 K-space for Wavelength Multiplexing

The same procedure of placing the ends of the gratings together is used to construct the **K**-space picture for  $\lambda$  multiplexing. Figure 3.15 shows the grating space picture for three  $\lambda$  multiplexed holograms. For clarity, the grating vectors are not shown. Again, notice that as the angular spread of the images ( $\Delta\theta_i = d/2F$ ) is increased the distance between the gratings is changed. Unlike in the  $\theta$  multiplexing case, the images do not rotate or move relative to each other as more holograms are stored. This explains why the SNR saturates to a fixed value for a large number of  $\lambda$  multiplexed holograms.

The closed form solution for the SNR for  $\lambda$  multiplexing holograms in a pure reflection geometry was given in Chapter 2 as

$$\text{SNR} = \frac{8F^2}{d^2} = 4\left(\frac{1}{\Delta\theta_i}\right)\left(\frac{1}{\Delta\theta_i}\right). \quad (3.6)$$

We see that for  $\lambda$  multiplexing, the SNR goes as one over the bandwidth of the image

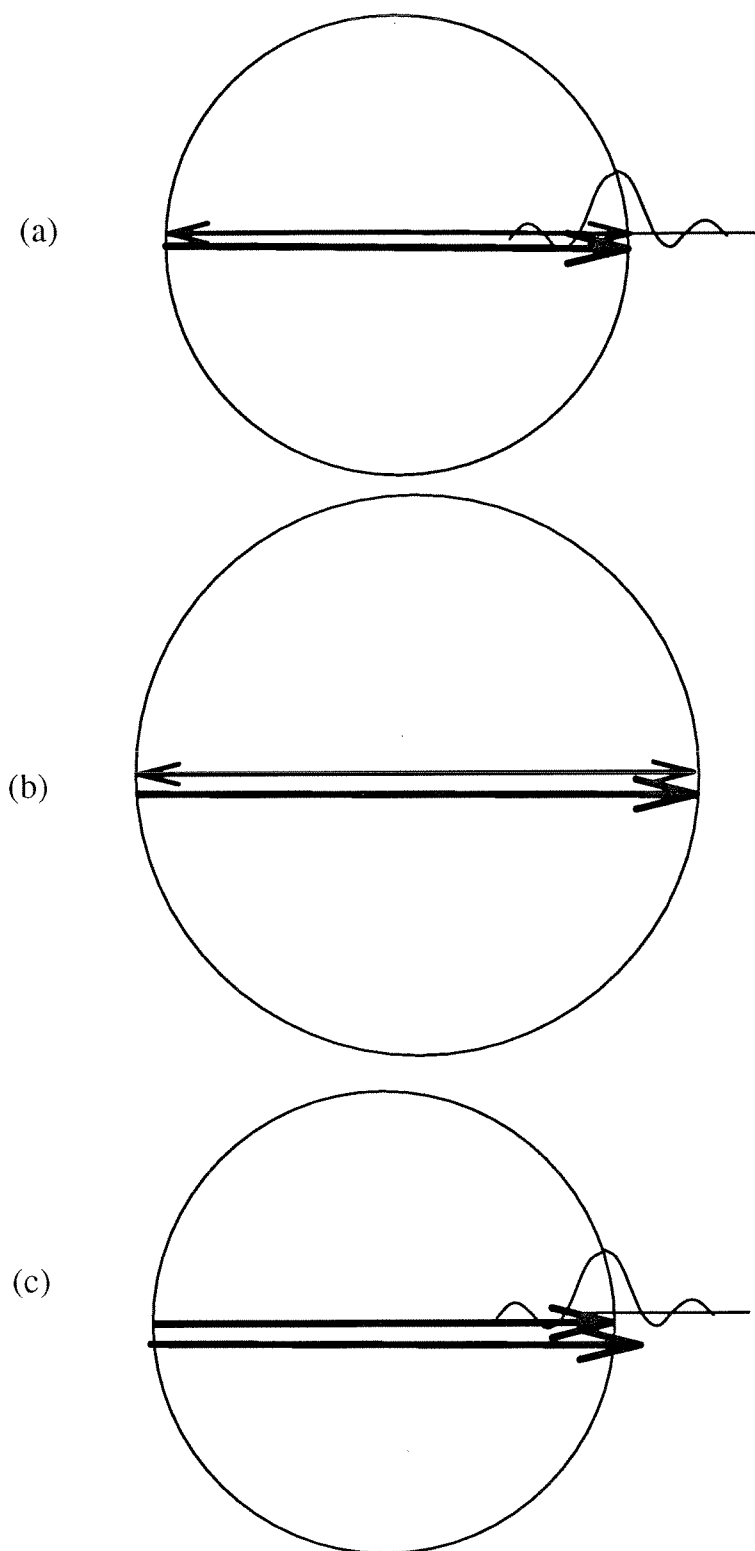


Figure 3.14: k-space diagram for reconstruction (a) single grating selectivity, (b) another grating angle multiplexing, (c) placing subsequent gratings in nulls of first grating.

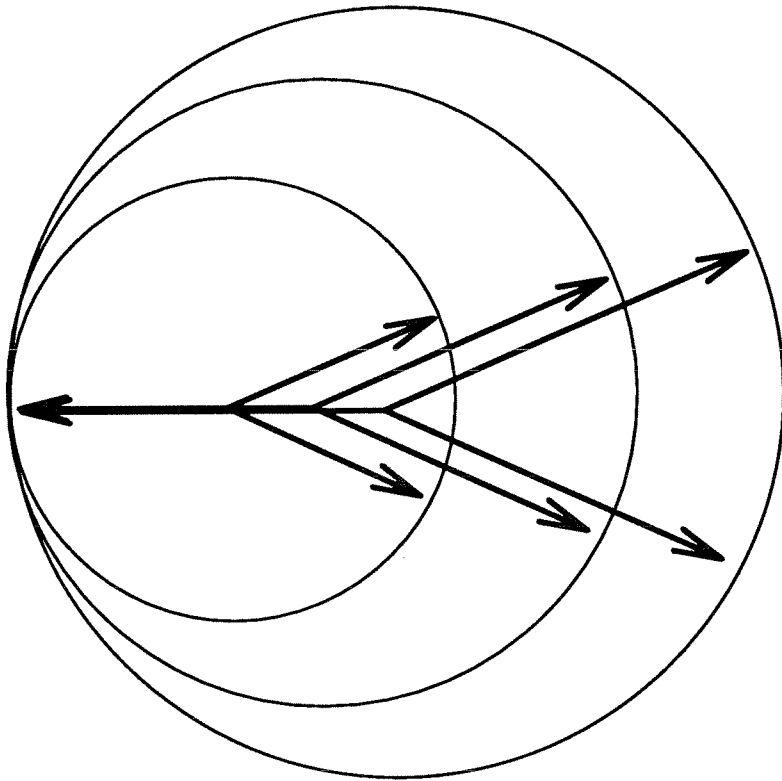


Figure 3.15:  $\mathbf{K}$ -space diagrams for wavelength multiplexing.

beam squared.

### 3.3.6 Comparison of Methods

For angle multiplexed Fourier transform holograms in a  $90^\circ$  configuration the minimum SNR is given by

$$SNR_\theta = \frac{2tf}{\lambda dN}, \quad (3.7)$$

where  $t$  is the interaction length,  $f$  is focal length,  $\lambda$  is the wavelength,  $d$  is the full length of the output/image plane, and  $N$  is the number of holograms. For wavelength multiplexed Fourier-transform holograms in a  $180^\circ$  configuration, the minimum SNR is given by

$$SNR_\lambda = \frac{8f^2}{d^2}, \quad (3.8)$$

where  $f$  is focal length, and  $d$  is the full length of the output/image plane. Therefore, for  $\lambda$  multiplexing, the SNR (for  $N >$  a few holograms) is independent of the number of holograms.

The SNR expressions can be compared to derive some general trends of how to best use each multiplexing method. Looking at the closed form solution we can see that the SNR depends on the bandwidth of the image beam or the reference beams. Looking at the functional behavior, the SNRs can be written as

$$\begin{aligned} SNR_\theta &\sim \frac{1}{dN_\theta} \\ SNR_\lambda &\sim \frac{1}{d^2}. \end{aligned} \quad (3.9)$$

Therefore, when  $\theta$  multiplexing one wants to store large images and fewer holograms. When  $\lambda$  multiplexing one should store many small images. Thus, increasing the page size favors  $\theta$  multiplexing, but increasing the number of holograms favors  $\lambda$  multiplexing.

We can divide Equation 3.8 by Equation 3.7 and get the ratio of the two SNR's. The number of holograms that can be stored is proportional to  $\frac{t}{\delta}$ , where  $\delta$  is the

resolution of the material. This fact follows from a degrees of freedom argument and is relatively easy to prove for  $\theta$  multiplexing. We will briefly show how it is also true for  $\lambda$  multiplexing. The largest optical frequency that we can use is  $2c/\delta$  and the smallest is  $2c/t$ . The selectivity of the  $\lambda$  multiplexed holograms to a change in the optical frequency is  $\delta\nu = c/t$ . The number of holograms that can be  $\lambda$  multiplexed is therefore  $c(\frac{1}{\delta} - \frac{1}{t})/(c/t) \approx \frac{t}{\delta}$ . Using this upper bound for  $N$ , the ratio of SNRs becomes

$$\frac{SNR_\lambda}{SNR_\theta} = \frac{4\lambda f N}{dt} \approx \frac{4\lambda f}{d\delta}. \quad (3.10)$$

Using typical experimental parameters of  $f/d = 5$  and  $\lambda \approx \delta$  results in an SNR ratio  $\sim 10$  when storing the maximum possible (geometrically limited) number of holograms. Notice that for thicker holograms,  $\theta$  multiplexing becomes more favorable since for a given number of holograms the needed  $\Delta\theta_r$  is smaller which reduces the rotation of the images in  $\mathbf{K}$ -space for the case for angle multiplexing.

Another way to compare the two methods is by solving Equation 3.7 for  $N$ , and then using the  $SNR_\lambda$  from Equation 3.8 to obtain an expression for the number of holograms that can be  $\theta$  multiplexed with the same SNR as that obtained by  $\lambda$  multiplexing. This results in an expression for the number of holograms that can be angle multiplexing as

$$N_\theta = \frac{td}{4\lambda f}. \quad (3.11)$$

Using typical parameters ( $t \approx 1\text{cm}$ ) results in  $N_\theta \sim 10^3$ . Let  $\Delta\nu$  be the spectral width of the light source and/or the spectral sensitivity of the material. Then the number of holograms that can be stored with  $\lambda$  multiplexing is

$$N_\lambda = \frac{\Delta\nu}{\delta\nu} = \frac{\Delta\nu t}{c}. \quad (3.12)$$

We can now obtain an expression for the ratio of the number of holograms that can be stored with the same SNR for  $\lambda$  and  $\theta$  multiplexing due to system constraints:

$$R = \frac{N_\theta}{N_\lambda} = \frac{1}{8} \frac{\frac{d}{f}}{\frac{\Delta\nu}{\nu}} = \frac{1}{4} \frac{\Delta\theta_i}{\frac{\Delta\nu}{\nu}}. \quad (3.13)$$

Therefore, it is the **fractional spectral bandwidth of the tunable laser versus the bandwidth of the optical system that carries the stored images** that determines which method can store more holograms in the same volume and with the same SNR. Using  $\Delta\theta_i \sim 1/5$  and  $\frac{\Delta\nu}{\nu} \sim 1/10$  results in a ratio on the order of one. Thus, the number of holograms that can be stored with the same SNR and in the same volume using either method is roughly equivalent.

### 3.4 Dynamic Range – When Does Crosstalk Dominate?

There are many sources of noise in a practical holographic memory system such as scattering, fanning, spatial light modulator uniformity, mechanical instabilities, and detector noise. To answer how these noise sources compare to cross-talk, we must consider how the grating strengths change as more holograms are recorded. Given that we want the diffraction efficiencies of the individual holograms to be the same, we know that the individual grating amplitudes vary as  $1/M$  where  $M$  is the number of holograms multiplexed in the same volume [155]. Therefore, the individual diffraction efficiency goes as  $1/M^2$ . The total NSR can now be written as

$$NSR = \frac{\alpha N_{ct}/M^2 + N_o}{\alpha/M^2}, \quad (3.14)$$

where  $N_{ct}$  is the normalized cross-talk noise,  $\alpha$  is the reconstructed intensity in  $\text{W}/\text{cm}^2$  that is obtained when a single hologram is recalled, and  $N_o$  is the noise from all sources that are independent of the number of holograms stored.

The crossover between cross-talk dominated noise and domination by other noise sources can be found by setting the two terms equal and solving for  $M$ . This results

in

$$M = \sqrt{\frac{NSR_x \alpha}{N_o}}, \quad (3.15)$$

where  $NSR_x$  is the cross-talk NSR. Using the closed-form asymptotic expressions for  $NSR_x$  with typical experimental parameters that we see in the lab, such as  $\alpha = .1 \text{ W/cm}^2$ ,  $N_o = 1 \text{ nW/cm}^2$  and the parameters mentioned previously, we find the crossover points to be  $M_\lambda \sim 350$  and  $M_\theta \sim 700$ . This means that if we  $\lambda$  multiplex more than 350 holograms,  $N_o$  will dominate since its contribution increases relative to the cross-talk noise. The crossover number for  $\phi$  multiplexing is even smaller than for  $\theta$  multiplexing, since  $\phi$  multiplexing has significantly better SNR. The exact value of the crossover depends on the parameters used; however, we can still conclude that for all three multiplexing methods in the optimal geometries **cross-talk is only important for a small number of holograms**, and SNR will be dominated by other noise sources in any holographic memory that multiplexes a large number of holograms in the same volume. Moreover, if the total number of holograms is small then the SNR is large and cross-talk is not the limiting factor in performance.

## Chapter 4

# Holographic Storage Using DuPont Photopolymer

### 4.1 Introduction to Photopolymers

#### 4.1.1 Background

In the early 1970's, while photorefractives were initially being developed, photopolymers were also being pursued for the purpose of holographic storage [100]–[110]. These photopolymers consist of monomers and a light sensitive dye inside a polymer matrix. Two plane waves interfere inside the material and sets up a sinusoidal intensity pattern as shown in Figure 4.1(a). The light excites the dye and initiates a free radical reaction. This locally polymerizes the monomers in the illuminated regions and sets up gradients in monomer concentration inside the polymer as in Figure 4.1(b). The monomers then diffuse into these regions and are polymerized, which increases the density of the film in the illuminated regions. This density change directly relates to a change in the index of refraction as in Figure 4.1(c). Thus, like photorefractives, these polymers record phase gratings. Unlike photorefractives, these films inherently record permanent gratings since the polymerization reaction is irreversible. Figure 4.1(d) diagrams the entire process with monomer diffusion and local polymerization. After the holograms are stored, the film is uniformly exposed to complete polymerization



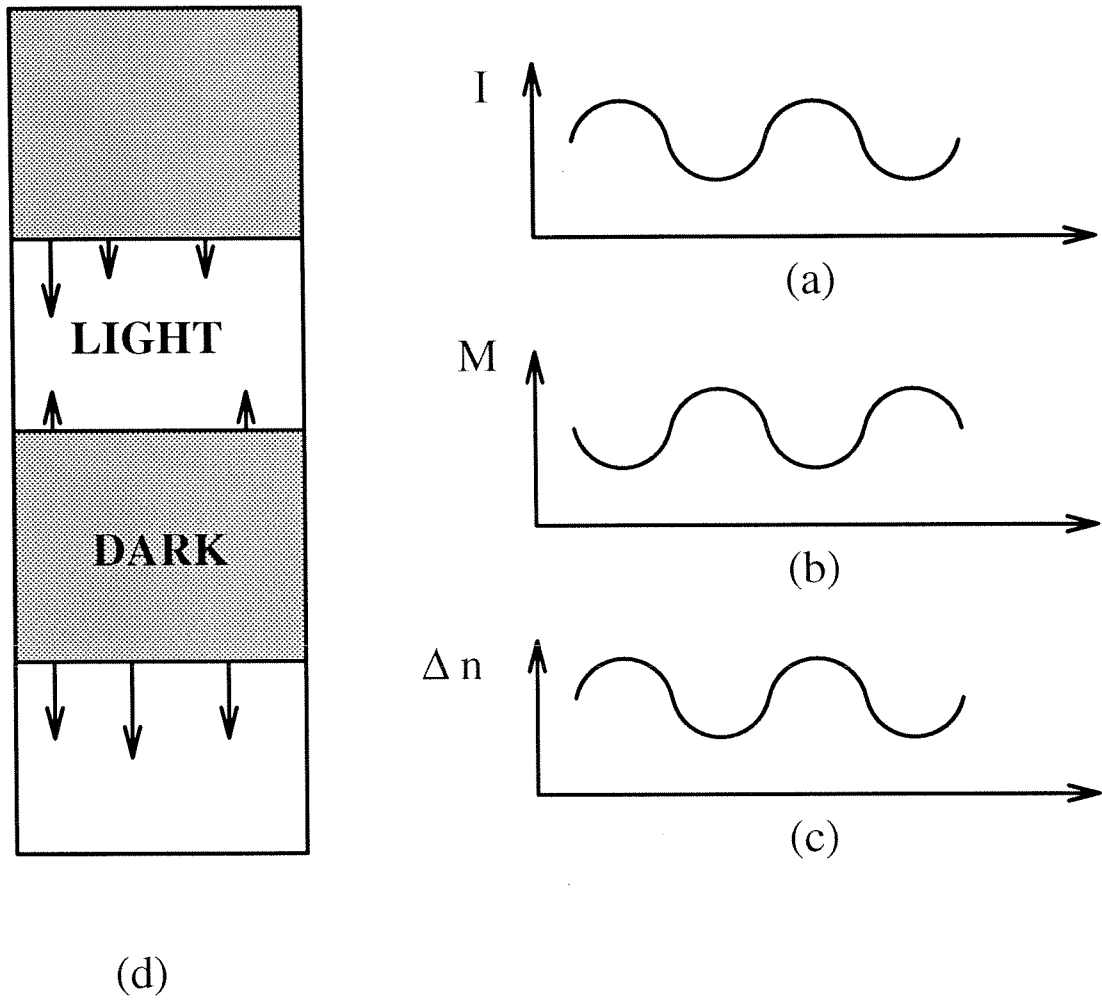


Figure 4.1: Light induced polymerization: (a) light intensity (b) monomer concentration after local polymerization (c) resulting index grating (d) diagram of the entire grating formation process.

and ensures the stability of the gratings and the film.

In 1976, Bartolini et al. [111] angle multiplexed 550 holograms in one centimeter thick photopolymer film. Using similar material, DuPont marketed a commercial holographic photopolymer product in the 1970's. Recently, DuPont, Polaroid, and others have photopolymer films [112]– [117] that have good sensitivity, resolution, and relatively large index changes. These films have been used for filters [118, 119, 120], holographic storage [121, 122], and the verification of drivers licenses and credit cards (although, the holograms on these cards are mechanically, not holographically reproduced). In addition, the theory of polymerization and monomer diffusion in photopolymer films has been investigated [123, 124].

## **4.2 Characterization of the DuPont Photopolymer for Holographic Storage**

### **4.2.1 Introduction**

DuPont's HRF-150 photopolymer film is investigated for use in 3-D holographic memories. Measurements of sensitivity, hologram persistence, the lateral spread of the photo-initiated reaction, and the variation of diffraction efficiency with modulation depth, spatial frequency/tilt angle, and intensity are reported. 3-D holographic disk systems can be used as memories and correlation devices with very high storage density and correlation rates [196, 120]. DuPont's photopolymer HRF-150 has excellent sensitivity and good resolution for transmission holograms recorded with blue-green light [112, 194, 195]. Thus it is a likely candidate for use as a holographic storage material. This section presents results on sensitivity, hologram persistence, the lateral spread of the photo-initiated reaction, and the variation of the diffraction efficiency with modulation depth, spatial frequency/tilt angle and intensity. These are all important characteristics for evaluating these photopolymer films as media for

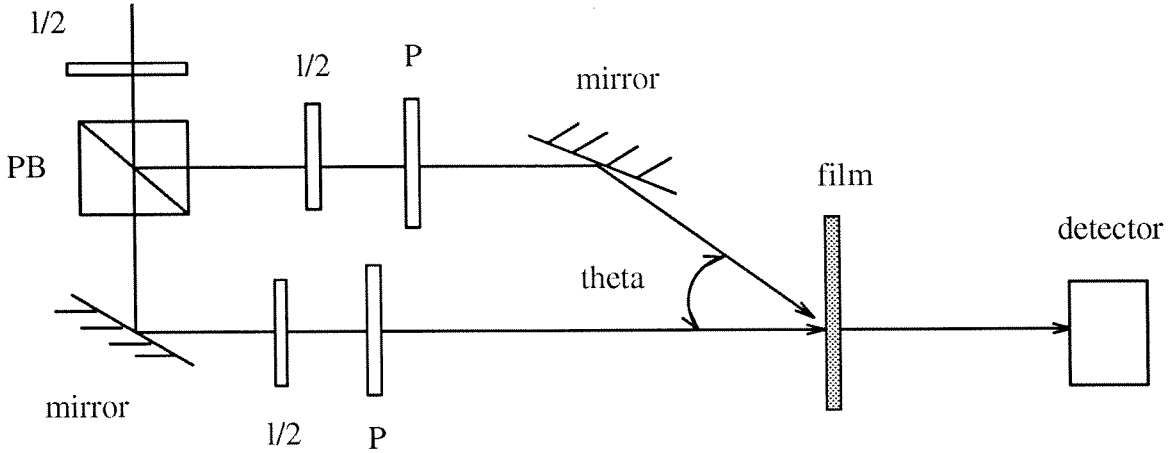


Figure 4.2: Recording geometry used for measurements with the elements given as: P – polarizer,  $1/2$  – half-wave plate, PB – polarizing beam splitter.

holographic storage applications.

#### 4.2.2 Recording Characteristics

HRF-150 photopolymer films consist of  $38\mu\text{m}$  thick photopolymer film sandwiched between two thin sheets of mylar. Hologram storage was performed using the setup shown in Figure 4.2. The object beam is incident normal to the surface, while the reference beam is incident at an angle  $\theta$ . Both beams are plane waves at 488 nm and the film is mounted on a glass substrate. After recording, the hologram is fixed by exposing the film to uniform UV light for 45 seconds. We define diffraction efficiency as incident intensity divided by diffracted intensity after the reference beam is re-Bragg matched to adjust for film shrinkage [122].

The sensitivity curve of the material is shown in Figure 4.3 where the diffraction efficiency of the first diffracted order is plotted against the total exposure energy. The intensity of each beam was kept constant at  $2\text{ mW/cm}^2$  with  $\theta \approx 18^\circ$  and the exposure time varied to obtain the data in Figure 2. When the diffraction efficiency exceeded 20%, a considerable amount of energy was diffracted into the 2 and -1 orders. Notice

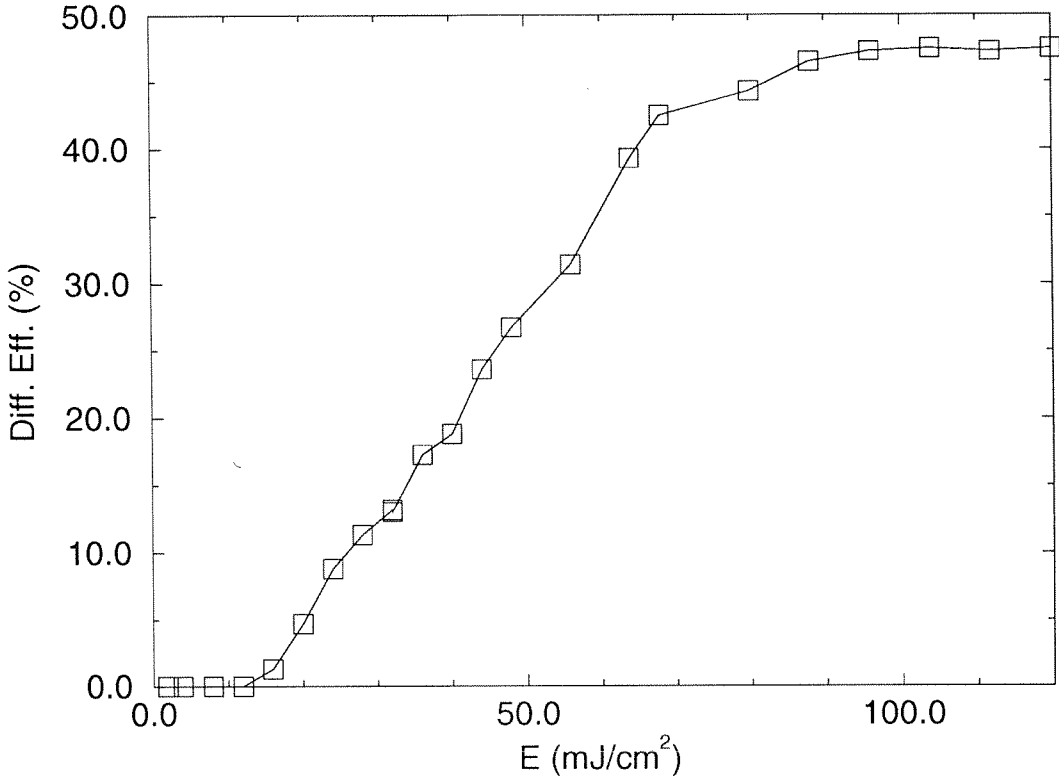


Figure 4.3: Diffraction efficiency vs exposure energy.

that the saturation exposure was approximately 80 mJ/cm<sup>2</sup>.

We also measured the effect of the intensity on the photopolymer's sensitivity. Keeping the intensities of the two beams equal,  $\theta \sim 20^\circ$ , and the total exposure of 20 mJ/cm<sup>2</sup>, the diffraction efficiency was measured as the total intensity of the two beams was changed. Figure 4.4 shows that as the intensity is increased, the film becomes less effective. Since the photo-initiated reaction is a free radical reaction, it can be slowed or stopped by inhibitors. Free radicals themselves are very efficient inhibitors, therefore, increasing intensity, which generates more free radicals, can inhibit the reaction.

The modulation depth is defined as  $m = 2RS/(R^2 + S^2)$  where  $R$  and  $S$  are the amplitudes of the reference and the signal beams. Figure 4.5 shows the diffraction efficiency versus  $m$  for  $\theta \sim 20^\circ$  with the total exposure kept constant at 20 mJ/cm<sup>2</sup>. The curve shows that if  $m < 0.2$ , the diffraction efficiency is small. This is probably

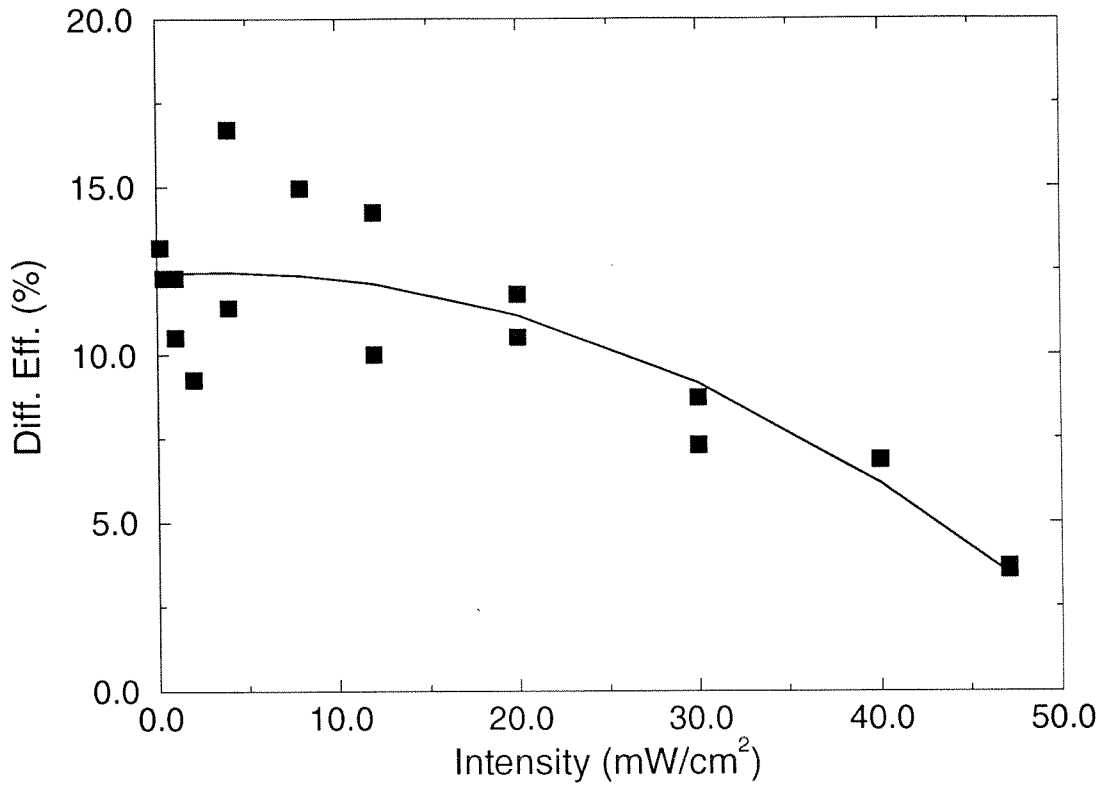


Figure 4.4: Diffraction efficiency vs intensity for  $E = 20 \text{ mJ}/\text{cm}^2$ .

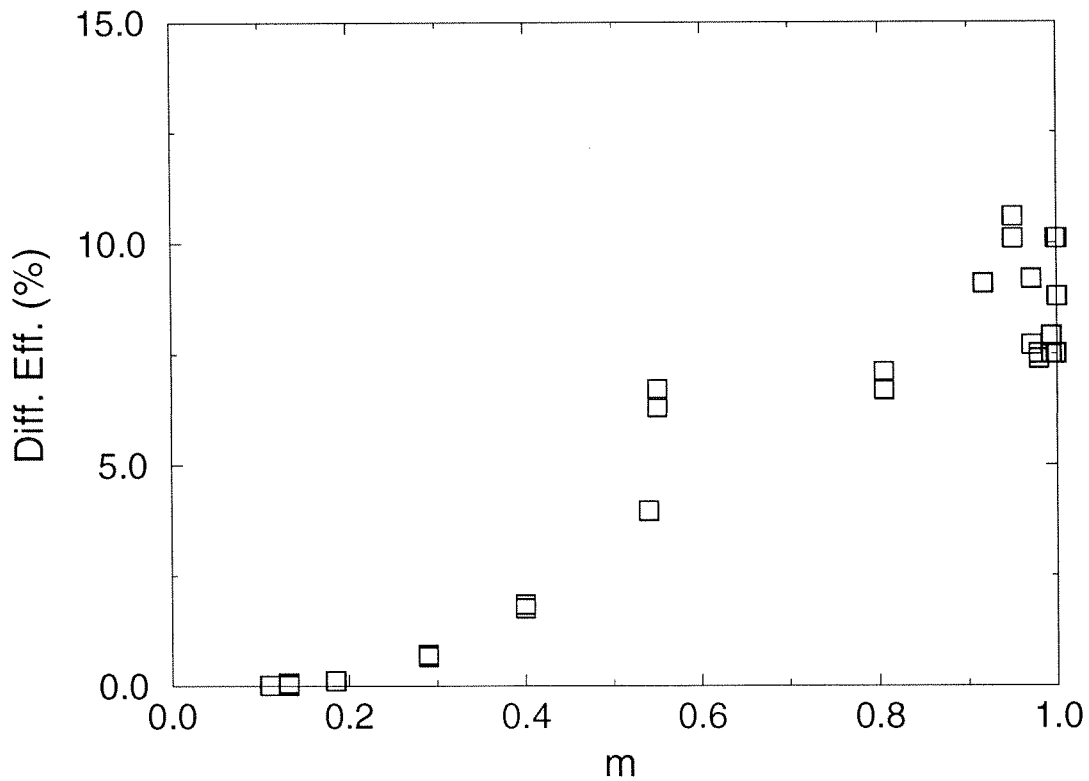


Figure 4.5: Diffraction efficiency vs modulation depth.

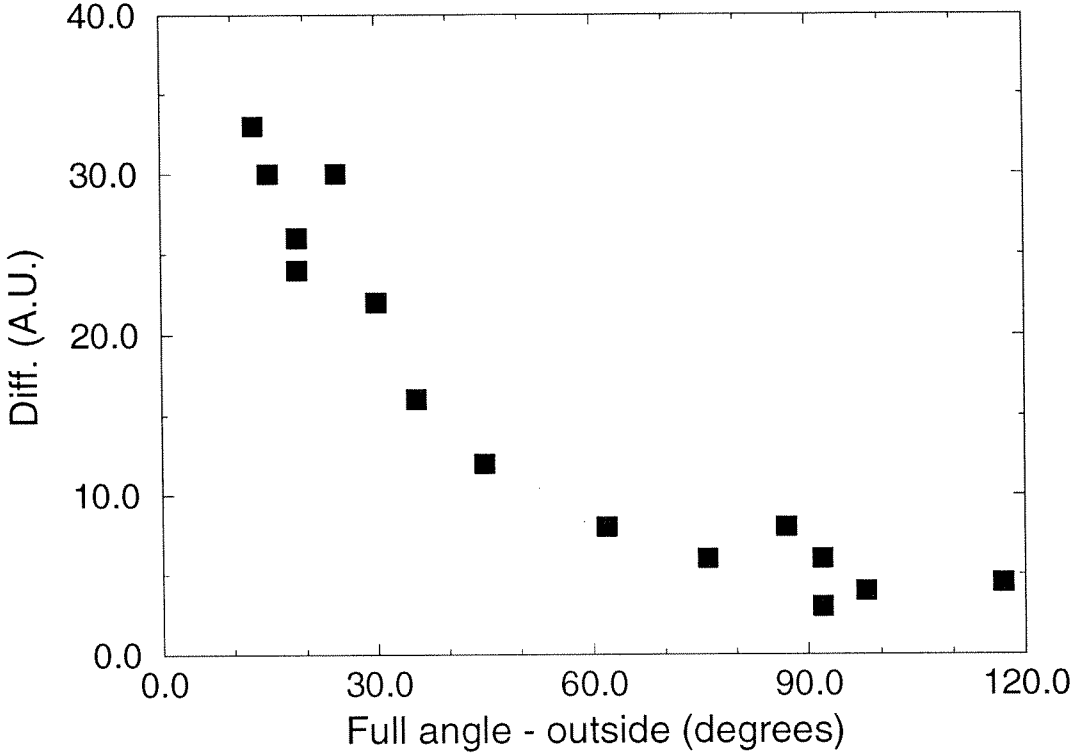


Figure 4.6: Diffraction efficiency vs full angle between the beams outside the material for  $E = 20 \text{ mJ/cm}^2$  and no grating tilt.

due to the fact that at low modulation the large background intensity polymerizes the material uniformly which impedes monomer diffusion. The curve also shows a saturation behavior around  $m = 1$ . This saturation is not due to the normal diffraction from a strong grating because we have a relatively weak modulation ( $\frac{2\pi\Delta nL}{\lambda n} < .33$ ). Therefore, we attribute this saturation to the monomer diffusion mechanism in the film.

The film's response to grating frequency is measured by recording the diffraction efficiency as a function of the angle between the beams while keeping the bisector of the angle perpendicular to the film's surface. This geometry ensures that the fringes are always perpendicular to the film surface. Figure 4.6 shows the result for  $E = 20 \text{ mJ/cm}^2$ . The efficiency is plotted against the full angle between the beams to make it

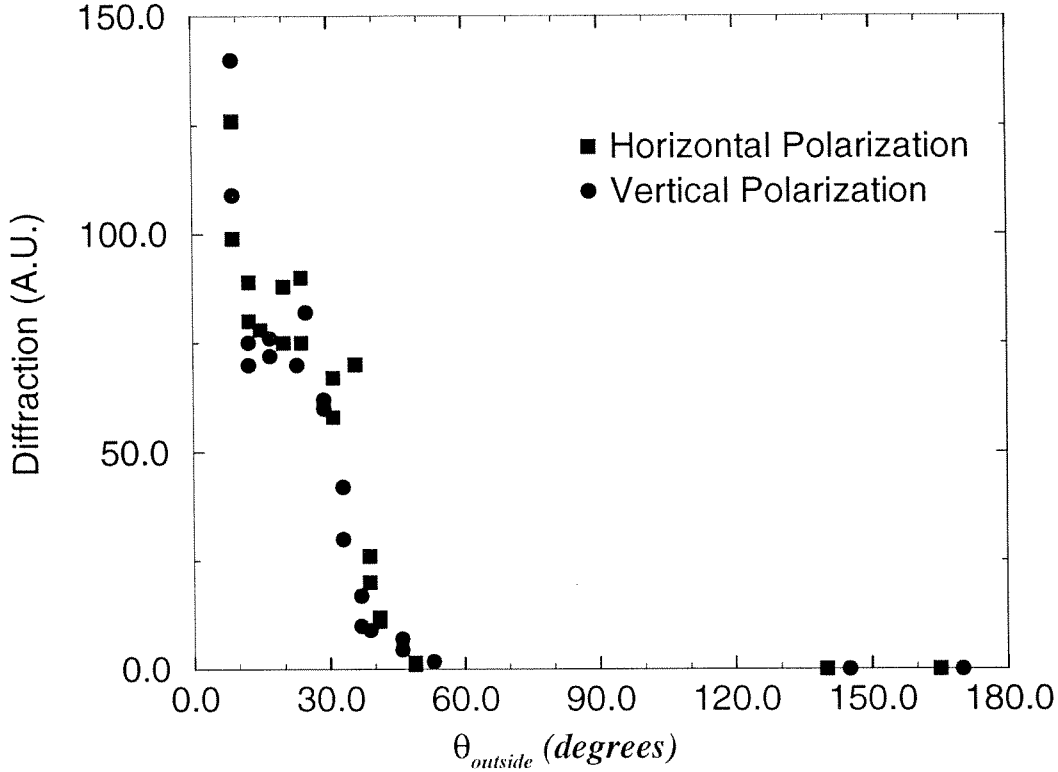


Figure 4.7: Diffraction efficiency vs angle outside the material for  $E = 20 \text{ mJ/cm}^2$  image beam incident normal to the surface.

easier to understand the next two results. Notice that the film prefers lower frequency gratings but the curve becomes flatter around  $90^\circ$  between the beams. The 3db point of the frequency response is at approximately  $30^\circ$ , corresponding to approximately 1,000 cycles/mm.

Figure 4.7 shows the diffraction efficiency for a total exposure of  $20 \text{ mJ/cm}^2$  as a function of the reference angle  $\theta$  for both polarizations with the signal beam incident normal to the surface. The intensity of each beam is  $2.2 \text{ mW/cm}^2$ . The results for each polarization are similar. HRF-150 is designed by DuPont as a transmission film and, as Figure 4.6 shows, it does not effectively record reflection holograms. Comparing this result with Figure 4.6, we see that the drop in efficiency for angles larger than  $30^\circ$  is not due entirely to the change in spatial frequency. It appears that the film does not effectively record gratings that have a large tilt angle inside the material. This is probably due to the film's thickness being only  $38\mu\text{m}$ .

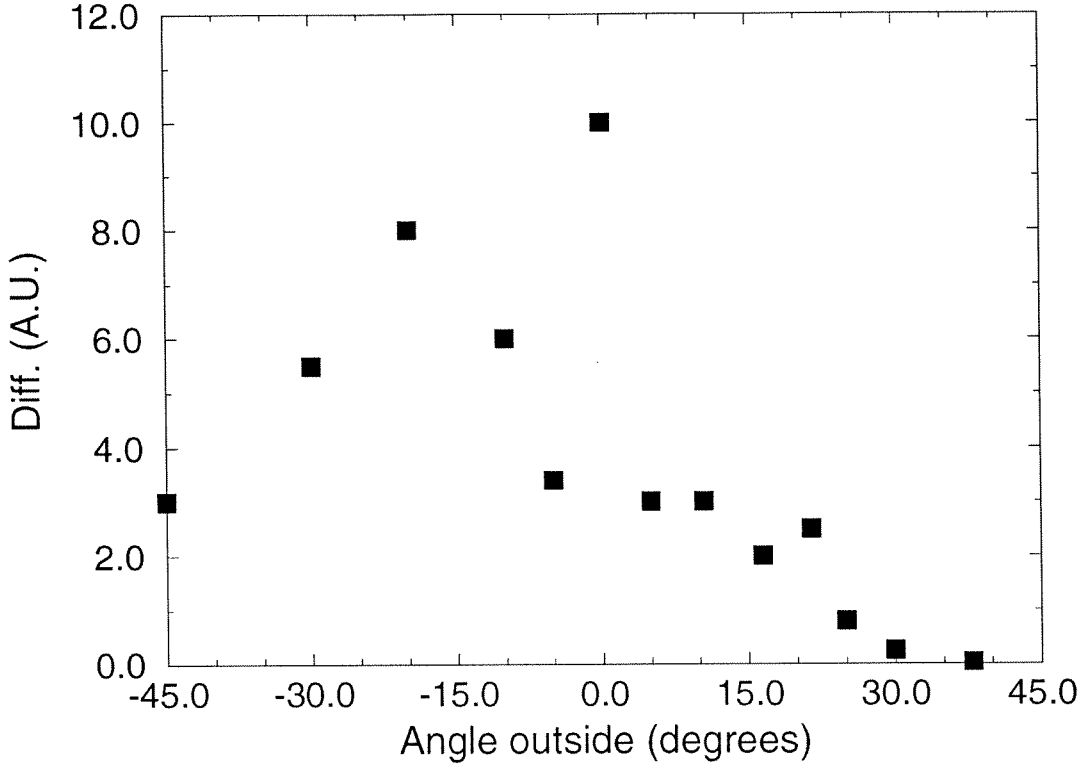


Figure 4.8: Diffraction efficiency vs angle outside the material for  $E = 20 \text{ mJ/cm}^2$  with  $90^\circ$  degrees between the  $0^\circ$  reference beam and the signal.

To evaluate this effect further, we again measured the film's response as the reference angle is changed, but this time with the center angle between the beams being  $90^\circ$  where we know (from Figure 4.6) that the film's frequency response is relatively flat. Figure 4.8 plots diffraction efficiency for a total exposure of  $20 \text{ mJ/cm}^2$  as a function of the reference angle  $\theta$ . In this plot, an angle of  $0^\circ$  refers to  $90^\circ$  angle between the beams with no tilt in the gratings. Notice that at  $-45^\circ$  the diffraction is less than at  $-20^\circ$  even though we know that the spatial frequency at  $-45^\circ$  is preferred by the film when the grating is not tilted. This loss in sensitivity is attributed to the grating tilt. The maximum fringe tilt that the film can tolerate is approximately  $10^\circ$ .

If the material is going to be used as a storage element, the holograms must be able to be recalled nondestructively after they are fixed with UV light. Figure 4.9 shows the continuous readout of a hologram for 100 hours. The reference intensity was  $\sim 3 \text{ mW/cm}^2$  with a diffraction efficiency of about 1.5%. The initial increase,



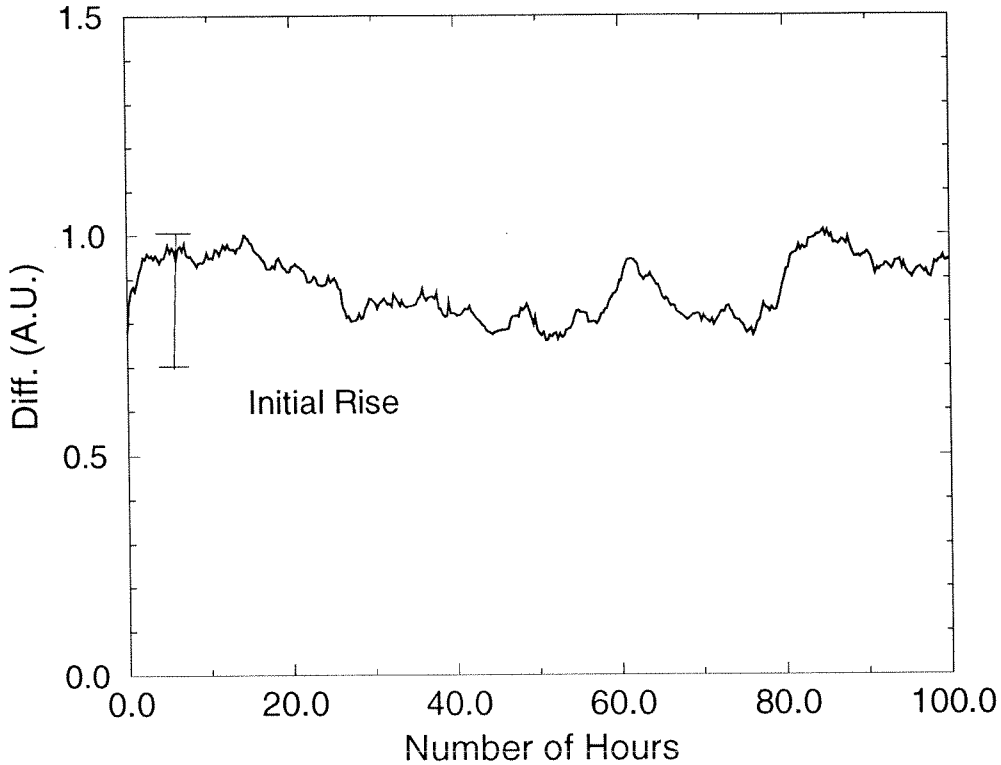


Figure 4.9: Diffraction vs readout time demonstrating nondestructive recall.

shown on the graph, is due to the bleaching of the material with light exposure. The grating was recorded with the object beam incident at  $-45^\circ$  and the reference beam incident at  $45^\circ$  from surface normal. Multiplexed image plane holograms recorded more than a year ago and stored at room temperature can still be recalled.

Whether or not the reaction spreads laterally (in the plane of the film) is important if spatial multiplexing is desired. To test if the reaction spreads, a slit about 1mm across and 1cm long was cut in tin-foil to create a mask. The photopolymer was exposed with a single normal incident beam through this mask ( $E \approx 300 \text{ mJ/cm}^2$ ). The sample was then left in the dark, giving the reaction time to polymerize the material in the exposed region and possibly laterally spread to beyond the illuminated area. The mask was then removed and a hologram recorded in a large area around the slit's location on the film. The reconstructed hologram consists of bright areas where the film was unaffected by the first exposure, and dark areas where the film had already

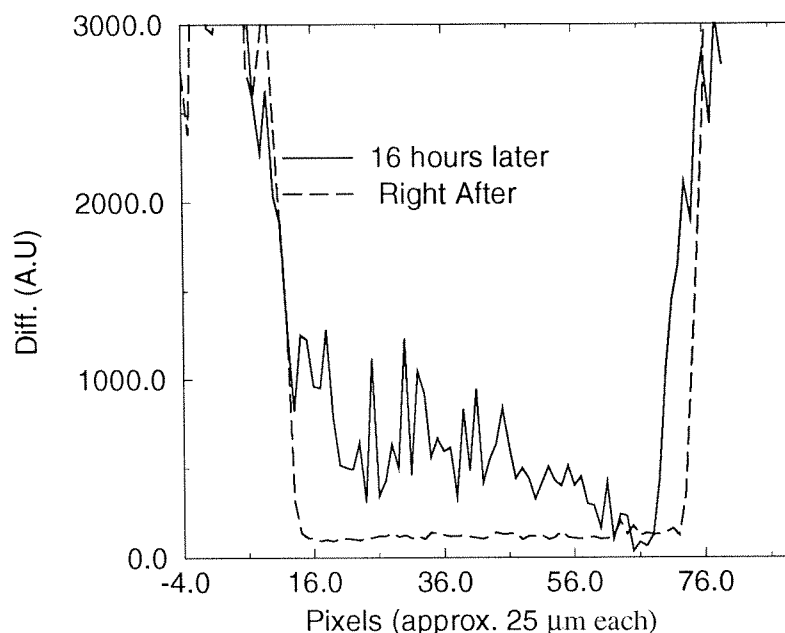


Figure 4.10: Cross sections of two reconstructed holograms across pre-exposed areas.

been exposed. Any lateral spreading of the reaction would cause a gradual loss of efficiency as the reaction spreads out and uses up the film's dynamic range. When the second exposure immediately followed the first, no appreciable lateral spread was observed, as shown in Figure 4.10. The curve marked "Right After" shows a cross section of the intensity of the reconstruction across the dark area. The experiment was repeated, except this time the sample was left in the dark for 16 hours before the holographic exposure was done. Again, a dark slit was seen in the hologram, and a cross section marked "16 hours later" is shown in Figure 4.10. Notice that at the transition from dark to light the slope of the two curves is approximately the same. Therefore, within a measurement error of  $\sim 100\mu\text{m}$ , the reaction does not laterally spread. The 16 hour curve is slightly narrower width because it was taken at a different place along the slit.

In summary, we have observed that the sensitivity of the HRF-150 photopolymer

decreases with increases in intensity and grating tilt angle. The holograms can be nondestructively reconstructed for long periods of time at room temperature. The photo-initiated reaction was seen to spread less than  $100\mu\text{m}$  over periods of many hours. Overall, the HRF-150 has many of the desirable characteristics needed for a practical holographic storage material. The major improvement that is needed is an increase in film thickness to  $200\mu\text{m}$  or more. This would allow more holograms to be multiplexed at a single location and also it would probably alleviate the problem with tilted gratings that is reported in this paper.

## 4.3 Phase Grating Profiles in Photopolymers

### 4.3.1 Introduction

Thick volume gratings have been investigated since the beginning of the century in connection with X-ray diffraction, light diffraction by acoustic waves, and holography. The simplest theory [26] takes into account two waves (the diffracted wave and the transmitted wave) diffracting off a pure sinusoidal phase grating. Considerable work has been done on calculating the effect of higher order modes for arbitrary grating shapes at on-Bragg angle incidence [28, 197, 198, 199]. This was expanded for pure sinusoidal gratings to off-Bragg angles of incidence by B. Benlarbi et al. [200]. Recently, for photorefractives, Kukhtarev's equations have been solved for higher modes at on and off-Bragg angles of incidence [201, 202]. In addition to photorefractives, there is reason to believe that photopolymers form gratings that are nonsinusoidal [124]. Solutions to the nonlinear diffusion equation that governs photopolymer polymerization are given in reference [124], and they demonstrate that nonsinusoidal gratings are possible.

The criteria for the validity of the two mode theory is related to two quantities  $F$

and  $Q'$  [27], where  $F$  and  $Q'$  are defined as

$$\begin{aligned} F &= \frac{16n^2 \sin^2 \theta}{\epsilon_{r1}} \\ Q' &= \frac{2\pi\lambda L}{\Lambda^2 \cos \theta} . \end{aligned} \quad (4.1)$$

In the above equations  $\lambda$  is the wavelength of the radiation,  $L$  is the interaction length,  $\theta$  is the interior angle of incidence measured from the perpendicular,  $\epsilon_{r1}$  is the relative permittivity of the perturbation,  $n$  is the index of refraction of the material, and  $\Lambda$  is the grating period. For sinusoidal gratings at on and off-Bragg angles of incidence,  $Q' \gg 1$  and  $F \gg 1$  is a sufficient condition for two mode theory to be adequate. By recording holograms in a geometry such that either  $Q'$  or  $F$  is small, the angular selectivity and diffraction efficiencies of the modes are determined by the grating profile.

Analyzing the grating profile and its effect on diffraction efficiency and angular selectivity is important for the effective use of photopolymers in optical systems that operate in the small  $F$  or small  $Q'$  regimes. For example, using thermoplastic plates, a system that uses the second-order mode instead of the first-order mode has been demonstrated [203]. In addition, if the gratings are similar to what is predicted by the nonlinear diffusion equation, this would indicate that diffusion is the dominant process in photopolymer grating formation. For these reasons, determination of the grating profile is the subject of this section.

### 4.3.2 Theoretical Model

We shall consider a one-dimensional, lossless, unslanted phase grating with the setup shown in Figure 4.11. Because the grating formation conditions are symmetrical around the peak of the sinusoidal intensity, we will look for symmetrical solutions

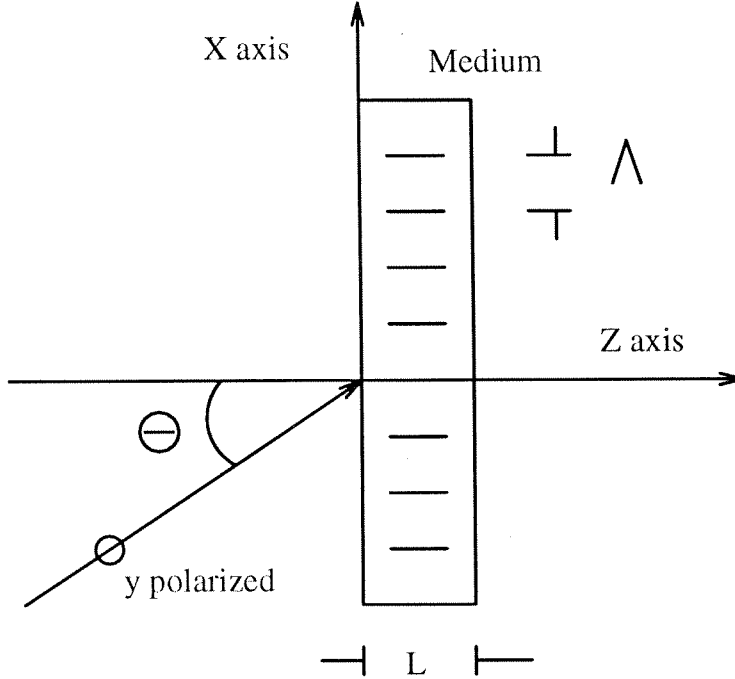


Figure 4.11: Geometry for diffraction from unslanted gratings.

without a phase shift. Therefore the relative permittivity is defined as

$$\epsilon_r(x) = \epsilon_{r0} + \sum_{i=1}^{\infty} \epsilon_{ri} \cos(iK_h x) \quad (4.2)$$

where  $K_h = 2\pi/\Lambda$ , and  $\Lambda$  is the grating period. Note that  $\Delta n_i$  is approximately equal to  $\epsilon_{ri}/2n$ . Using this information and the quantities defined in Figures 4.11 and 4.12, the problem is reduced to solving the scalar wave equation for  $E_y$ .

$$\frac{\partial^2}{\partial x^2} E_y + \frac{\partial^2}{\partial z^2} E_y - \mu_0 \epsilon_0 \epsilon_r(x) \frac{\partial^2}{\partial t^2} E_y = 0 \quad (4.3)$$

Now we assume a solution of the form of:

$$E_y(x, z) = e^{j\omega t} \sum_{m=-\infty}^{\infty} A_m(z) e^{j\phi_m} \quad (4.4)$$

where

$$\phi_m = \frac{2\pi n}{\lambda} (\sin(\theta)x + \cos(\theta)z) + mK_h x. \quad (4.5)$$

Plugging equation 4.4 into equation 4.3 and assuming the interchange of energy between modes is slow with respect to a wavelength (so that the second order derivatives can be neglected) results in a series of coupled differential equations for the mode amplitudes  $A_m(z)$ ,

$$\sum_{m=-\infty}^{\infty} \frac{\partial}{\partial z} A_m(z) + j \frac{m\beta}{L} \left( m + \frac{2n \sin \theta \Lambda}{\lambda} \right) A_m(z) - j \sum_{i=1}^{\infty} \alpha_i (A_{m+i}(z) + A_{m-i}(z)) = 0 \quad (4.6)$$

where  $m$  is the number of the mode, and  $\beta$  and  $\alpha_i$  are defined as

$$\beta \equiv \frac{\pi \lambda L}{n \cos \theta \Lambda^2} \quad (4.7)$$

$$\alpha_i \equiv \frac{\pi \epsilon_{ri}}{2n \cos \theta \lambda} . \quad (4.8)$$

The initial conditions for the modes are

$$A_0(0) = 1 \quad A_m(0) = 0 \quad \text{for } m \neq 0 . \quad (4.9)$$

These coupled differential equations were solved numerically by using a fourth order Runge-Kutta algorithm [204] to numerically integrate the equations. Seven modes were used (-3 to +3). When converting from exterior to interior angles, Snell's Law was used with  $n = 1.525$  as the index of refraction of the material. The plots were graphed using

$$\gamma = \left| \frac{A_{-1}(\theta)}{A_{-1}(\theta_{Bragg})} \right|^2, \quad (4.10)$$

where  $A_{-1}(\theta_{Bragg})$  is the amplitude of the -1 mode at Bragg incidence.

### 4.3.3 Experimental Results

To record a hologram that has low  $F$  and  $Q'$  values requires a very small angle between the recording beams. The geometry used is shown in Figure 4.12. The thickness of the sample was measured with a second hologram, recorded in the regime that Kogelnik's theory is valid (large  $F$  and  $Q'$ ). The two plane wave holograms were sequentially recorded in HRF-150 photopolymer using 488 nm light. Both holograms

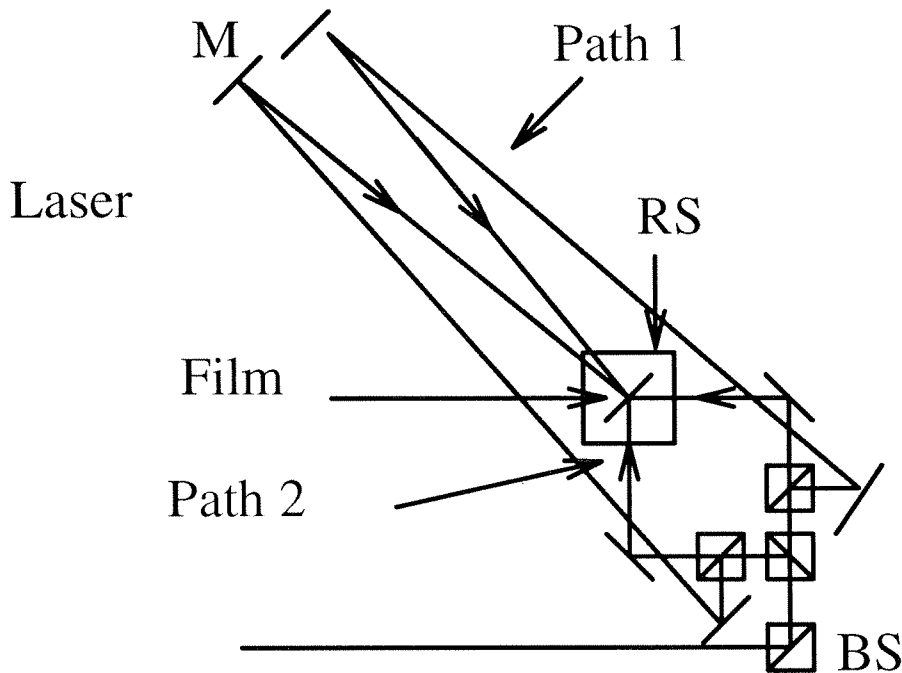


Figure 4.12: Experimental Setup. BS - beam splitter, RS - rotation stage, M - mirror, BE - beam expander, L - lens.

record unslanted gratings by aligning the reflections of a HeNe laser before making exposures with the blue light (equal and opposite angles from material's normal). First the hologram of interest (low spatial frequency) was recorded using Path 1 on Figure 4.12. Path 2 was blocked during this time. The exposure time for the first hologram was 50 seconds with a reference beam's intensity of  $104 \mu\text{W}/\text{cm}^2$  and an object beam's intensity of  $73 \mu\text{W}/\text{cm}^2$ . The -1 mode's diffraction efficiency was 9.17%. This was measured by taking the diffracted intensity of the -1 mode and dividing by the sum of the diffracted intensities of all modes plus the undiffracted light. The second hologram was recorded with 90 degrees between the beams using Path 2 while Path 1 was blocked. The exposure time was 30 seconds with a reference beam intensity of  $77 \mu\text{W}/\text{cm}^2$  and an object beam intensity of  $77 \mu\text{W}/\text{cm}^2$ . Its diffraction efficiency was 10.1%. The sample was then fixed by exposure to UV light.

The thickness of the sample was given by DuPont as nominally  $38 \mu\text{m}$ . By mea-

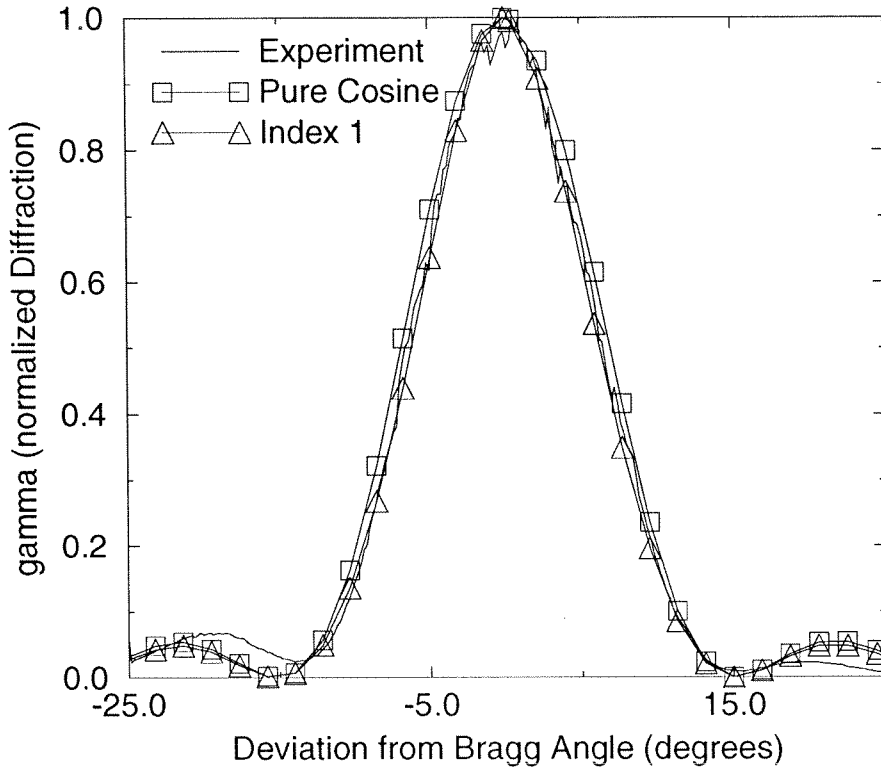


Figure 4.13: Theoretical and experimental -1 order mode angular selectivity.

measuring the angular selectivity of the second hologram the thickness can be accurately measured. Angular selectivity was determined by measuring the diffracted intensity while rotating the sample and dividing by the diffracted intensity at Bragg incidence giving a  $\gamma$  vs angle curve. In this way the thickness of the sample was measured to be 38.4  $\mu\text{m}$ . The angular selectivity of the first hologram was measured in the same way and produced the experimental curve in Figure 4.13. The first hologram was recorded with a half angle between the beams of 2.1 degrees. This was measured by replacing the sample with a mirror normal to one beam and using the rotation stage to rotate the mirror to reflect the beam back along the other beam. The experimental diffraction efficiencies of the various modes measured at the -1 mode's Bragg angle are:  $\eta_{-1} = 9.17\%$ ,  $\eta_1 = 6.80\%$ , and  $\eta_{-2} = 1.0\%$ .



### 4.3.4 Comparison

The pure cosine curve in Figure 4.13 was generated using the parameters of  $n = 1.525$ ,  $\Delta n = 1.281 \times 10^{-3}$ , thickness  $L = 38.4 \text{ } \mu\text{m}$ , external wavelength  $\lambda = 488 \text{ nm}$ , and  $\Lambda = 6.6587 \times 10^{-6} \text{ meters}$ . This corresponds to  $F = 5.5$  and  $Q' = 1.74$ . The  $\Delta n$  was found by matching diffraction efficiency of the -1 mode at the Bragg angle. Notice that the curve is wider than the experimental curve. The predicted diffraction efficiencies at the -1 mode's Bragg angle are  $\eta_{-1} = 9.17\%$ ,  $\eta_1 = 7.17\%$ , and  $\eta_{-2} = 0.20\%$ . This also differs from the experimental results.

By adding another coefficient in the cosine expansion of the perturbation, the angular selectivity could be dramatically and apparently uniquely altered. By varying this parameter to find a good fit, a second order correction term to the grating profile can be determined. The "Index1" curve on Figure 4.13 is a better match to the experimental data. It was calculated using the same parameters but with  $\Delta n_1 = 1.21 \times 10^{-3}$  and  $\Delta n_2 = 4.423 \times 10^{-4}$ . The predicted diffraction efficiencies of  $\eta_{-1} = 9.16\%$ ,  $\eta_1 = 6.59\%$ , and  $\eta_{-2} = 0.74\%$  also agree better with the experimental results. "Index1" is compared to a pure cosine in Figure 4.14. It also is similar to the predicted solutions to the nonlinear diffusion equation given in Reference [124].

### 4.3.5 Conclusions

The angular selectivity of the -1 mode, and the diffraction efficiencies of the modes depend on the grating profile when  $F$  and/or  $Q'$  are small. By measuring these quantities and comparing to theory, a second order correction term was determined. Since experimentally both  $F$  and  $Q'$  were small, a small change in the second order coefficient produced noticeable changes in the predicted angular selectivity curve. The relatively small second order coefficient and small changes in the angular selectivity means that the material in this geometry records fairly linearly. The postulated in-

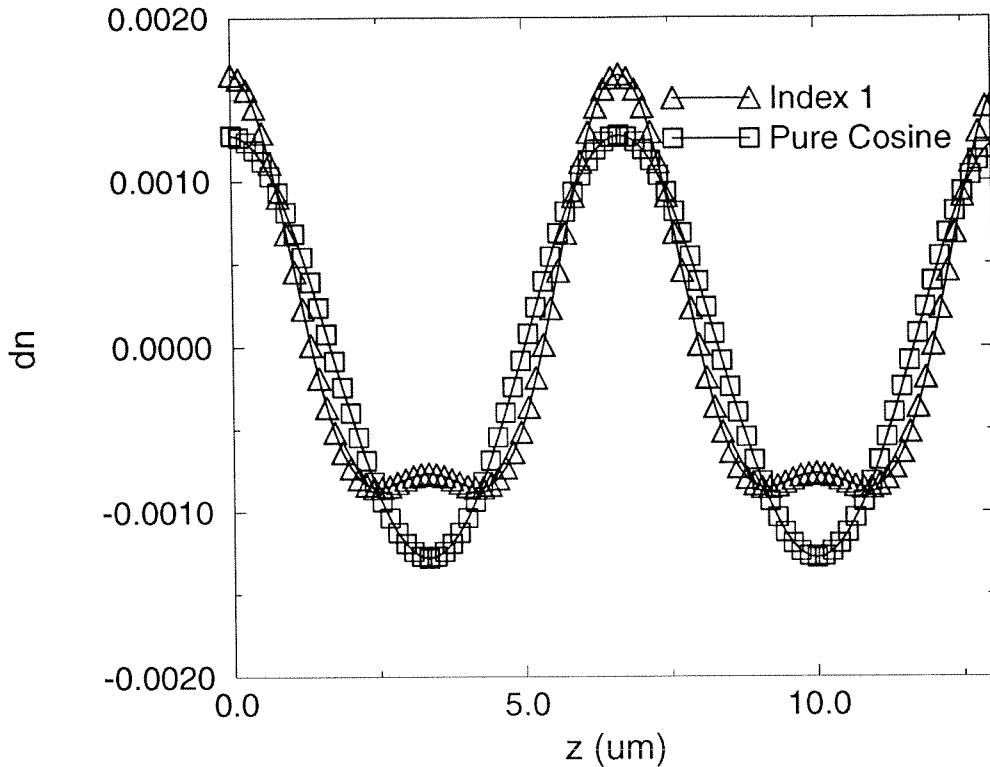


Figure 4.14: Index profiles.

dex profile from these measurements resembles the possible solutions of the diffusion model for photopolymer grating formation given in reference [124]. This agreement indicates that diffusion could be an important factor in grating formation as opposed to local effects. Since the grating profile determines the angular selectivity and diffraction efficiencies of the modes, it is an important parameter if systems that use these other modes are to be implemented using photopolymers.

## 4.4 Recording of Multiple Holograms in Photopolymer Films

### 4.4.1 Introduction

Angle multiplexed holograms have potential applications as optical memories, information processors, and interconnections. Important considerations include sensitivity, resolution, and the ability to record permanent holograms. DuPont's photopolymer

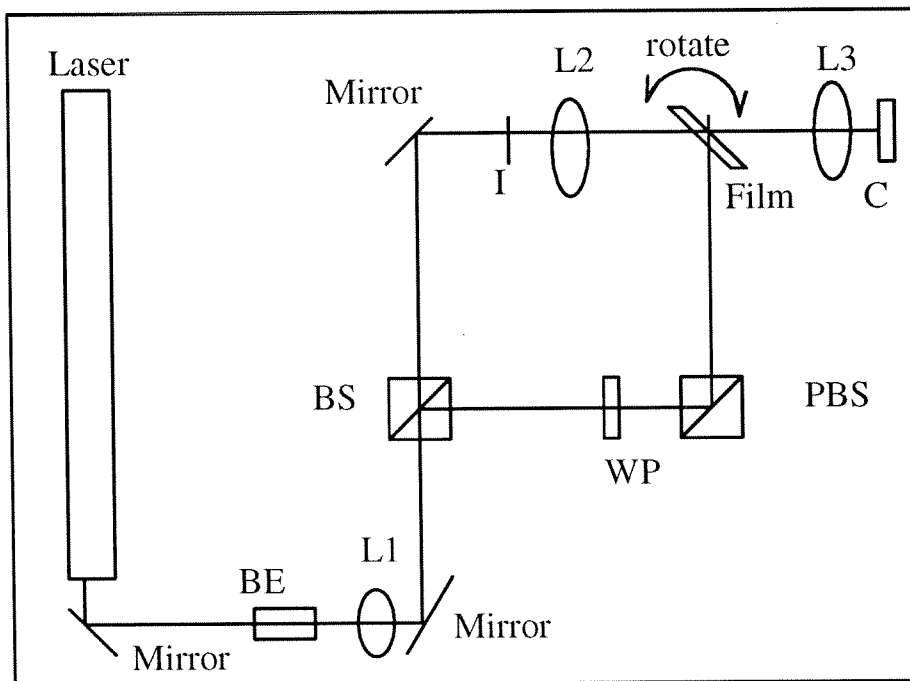


Figure 4.15: Geometry: BE - beam expander, L1 - lens, BS - beam splitter, L2,L3 - imaging lenses, WP - 1/2 waveplate, PBS - polarizing BS, I - image, C - camera.

HRF-150 has already been shown to have excellent sensitivity and resolution for transmission holograms recorded with blue-green light. With this in mind, HRF-150 is a good candidate for a practical, optical storage medium in which multiple holograms can be recorded. In this section we report the experimental demonstration of multiple holograms recorded in HRF-150.

#### 4.4.2 Experiment

Ten images were recorded using the geometry shown in Figure 4.15. A sheet of DuPont HRF-150 photopolymer was taped to a glass plate with a centered window cut out of it and mounted on a rotation stage. The holograms were recorded with visible light ( $\lambda = 488\text{nm}$ ) with 90 degrees between the writing beams and then fixed with UV light. First the angular selectivity of the material was determined at 488 nm. The material was approximately 38 microns thick, and had angular selectivity

( half width, measured at the first minimum) of one degree in very good agreement with Kogelnik's two wave theory [26] for diffraction from thick phase gratings. The equation for the diffraction efficiency from an unslanted, lossless grating is

$$\eta = \frac{\sin^2(\nu^2 + \xi^2)^{1/2}}{(1 + \xi^2/\nu^2)}, \quad (4.11)$$

where

$$\nu = \pi \Delta n d / \lambda \cos \theta, \quad (4.12)$$

and

$$\xi = (K \cos(\pi/2 - \theta) - K^2 \lambda / 4\pi n) d / 2 \cos \theta. \quad (4.13)$$

In the above equations,  $\eta$  is the diffraction efficiency,  $\Delta n$  is the index perturbation,  $n$  is the index of the material,  $\lambda$  is the wavelength,  $\theta$  is the angle inside the material measured from the normal to the surface,  $K$  is the grating vector's magnitude, and  $d$  is the interaction length. Good agreement between experiment and theory for a single plane wave hologram recorded with 90 degrees separation between beams outside the material ( $\pm 45$  deg from the normal to the film surface) is shown in Figure 4.16. The theoretical plot was calculated using  $d = 39$  microns,  $n = 1.525$ ,  $K = 1.82 \times 10^7$  rad/m,  $\Delta n = 1.51 \times 10^{-3}$ ,  $\theta_{Bragg} = 27.62$  degrees, and  $\lambda = 488$  nm. The angles were adjusted according to account for refraction at the surfaces. The agreement of the experiment and theory, indicates two things. First, it indicates that the gratings are not due to surface deformations which has been observed in similar materials [102] for low spatial frequency holograms. Also it indicates that the recorded hologram is most likely a phase grating rather than an absorption grating since the diffraction efficiency of 17.1 percent would be too large for an absorption grating.

Figure 4.17 shows the diffraction efficiency for 10 holograms of the same image stored at 10 different angular setting of the recording medium while keeping the angles between the recording beams the same. The holograms were recorded at

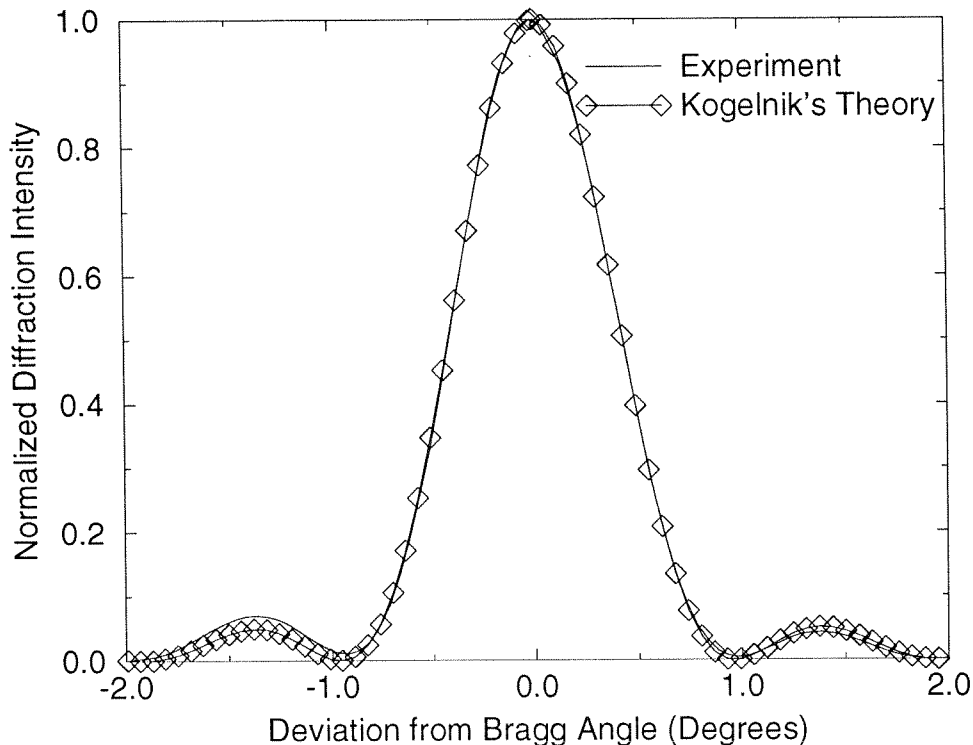


Figure 4.16: Angular selectivity of plane wave hologram.

intervals of two degrees or at the second minima of the angular selectivity curve of the adjacent holograms. The holograms were read out by blocking the object beam and detecting the reconstructed image while rotating the sample. Notice that the first two holograms ( at  $-10^\circ$  and  $-8^\circ$ ) are weaker than the others, which indicates that exposure to light enhances the sensitivity. Figure 4.18 shows 10 image holograms of roughly equal strength that were recorded with a five second pre-illumination pulse with  $1.85 \text{ mW/cm}^2$  intensity ten seconds before the first hologram was recorded. Both Figure 4.17 and 4.18 holograms were recorded using the setup in Figure 4.15, and, except for the pre-illumination pulse, identical recording parameters were used. The total recording energy per hologram was  $1.88 \text{ mJ/cm}^2$  (one second exposure per hologram) with an object-to-reference beam ratio of approximately  $1/4.6$ . The time between holograms is 10 seconds for a total experiment run time of 120 seconds. The diffraction efficiency shown is calculated by dividing the diffracted light by the incident

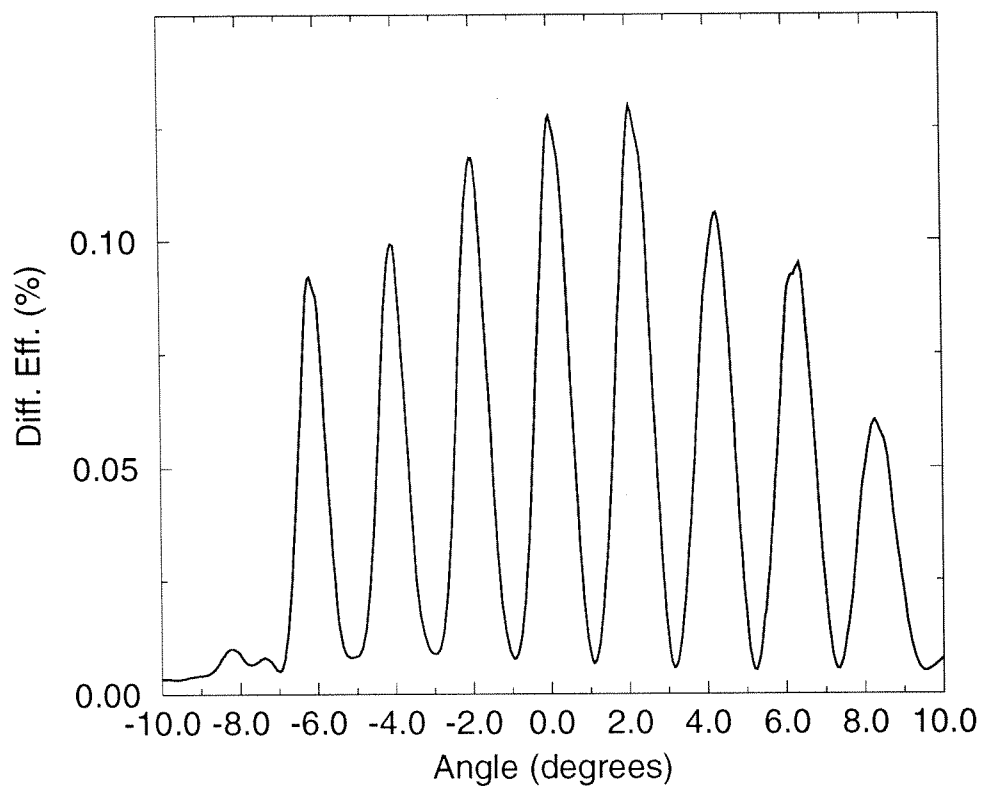


Figure 4.17: 10 holograms, no pre-exposure.

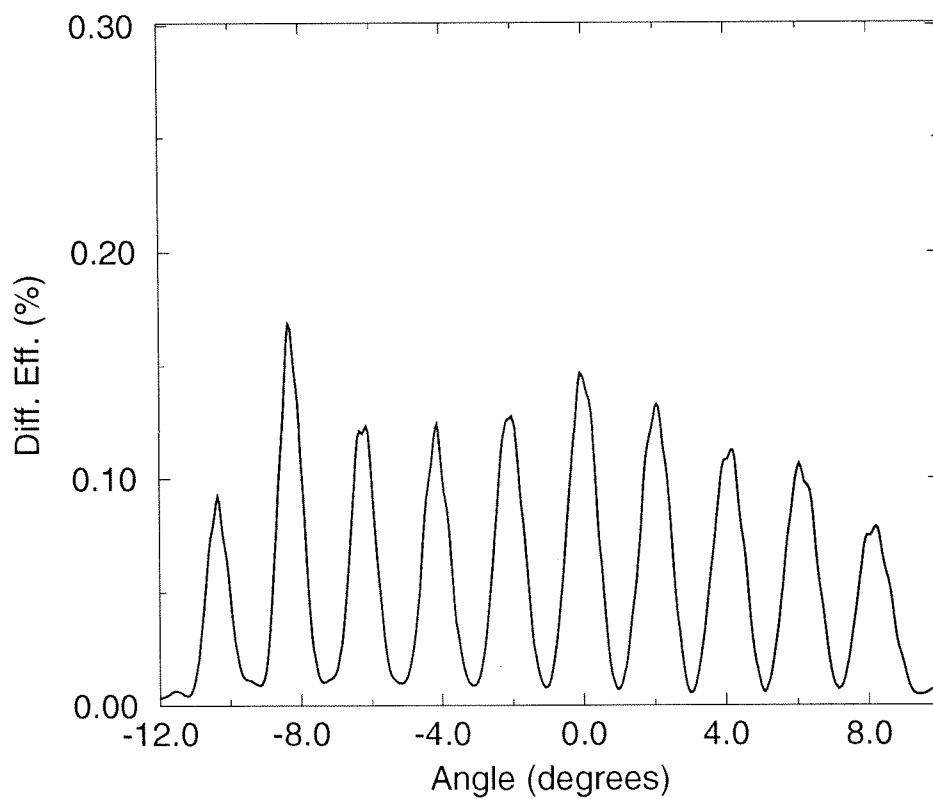


Figure 4.18: 10 image plane holograms.



Figure 4.19: Reconstructed image.

minus reflected light intensities, and multiplied by 100 to convert to a percentage. The pre-illumination pulse sensitizes the material and is necessary if the first hologram is to be recorded effectively. Since the material has a finite dynamic range, too long of a pre-illumination pulse wastes dynamic range causing the later holograms to be weaker than previous ones. Too short of a pre-illumination pulse fails to sensitize the material and the first few holograms are lost as shown in Figure 4.17.

The optical quality of the material is good, and Figure 4.19 shows a reconstructed image (hologram number 6 in Figure 4.18- at  $0^\circ$ ) with no noticeable distortion. Figure 4.20 shows the index perturbation vs exposure and demonstrates that the pre-illumination pulse does use up dynamic range and increases the sensitivity for small exposures. The index perturbation was calculated by measuring the Bragg matched diffraction efficiency ( $\xi = 0$ ) and solving for  $\Delta n$  in equation 4.11. In this case the diffraction efficiency used for  $\eta$  is diffracted light divided by transmitted light intensity

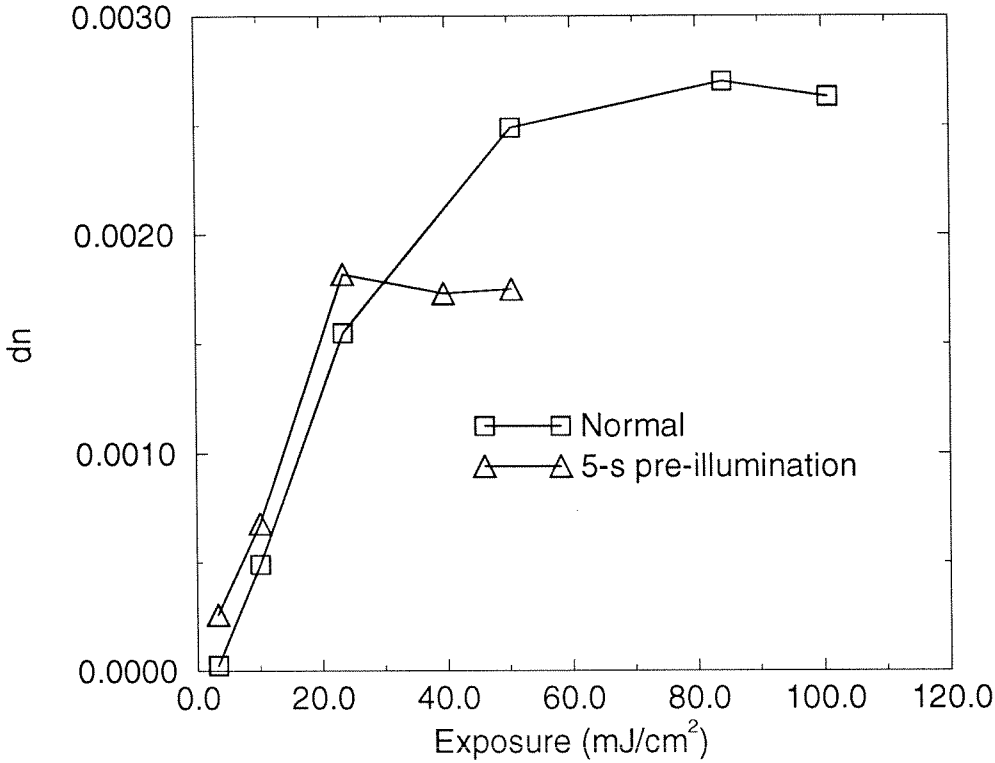


Figure 4.20: Index perturbation vs exposure.

to eliminate the effects of reflection from the surface and absorption. Notice also that the saturation exposure was about 20 mJ/cm<sup>2</sup>; therefore, each of the ten holograms should be given about 2 mJ/cm<sup>2</sup>. This is close to the actual value of 1.88 mJ/cm<sup>2</sup> per hologram used. Without the pre-illumination pulse the maximum  $\Delta n$  obtained was  $\Delta n \approx .0028$ . The  $\Delta n$  we measured is lower than what is reported in reference [112] ( $\Delta n \approx .008$ ), most likely because of the much higher spatial frequency recorded in our experiment (2897 cycles/mm versus 1007 cycles/mm). With the pre-illumination pulse, the saturation exposure for this material is 20 mJ/cm<sup>2</sup> for a  $\Delta n_{max} = 1.7 \times 10^{-3}$ . In comparison, for a thick sample of BaTiO<sub>3</sub> [205] the experimentally measured saturation exposure is approximately 100 mJ/cm<sup>2</sup>. Recent experiments with thin samples of BaTiO<sub>3</sub> [206] and doped SBN [207] indicate sensitivities and  $\Delta n$ 's roughly equivalent to the photopolymer.

Figure 4.21 shows the effect of extending the total run time by increasing the



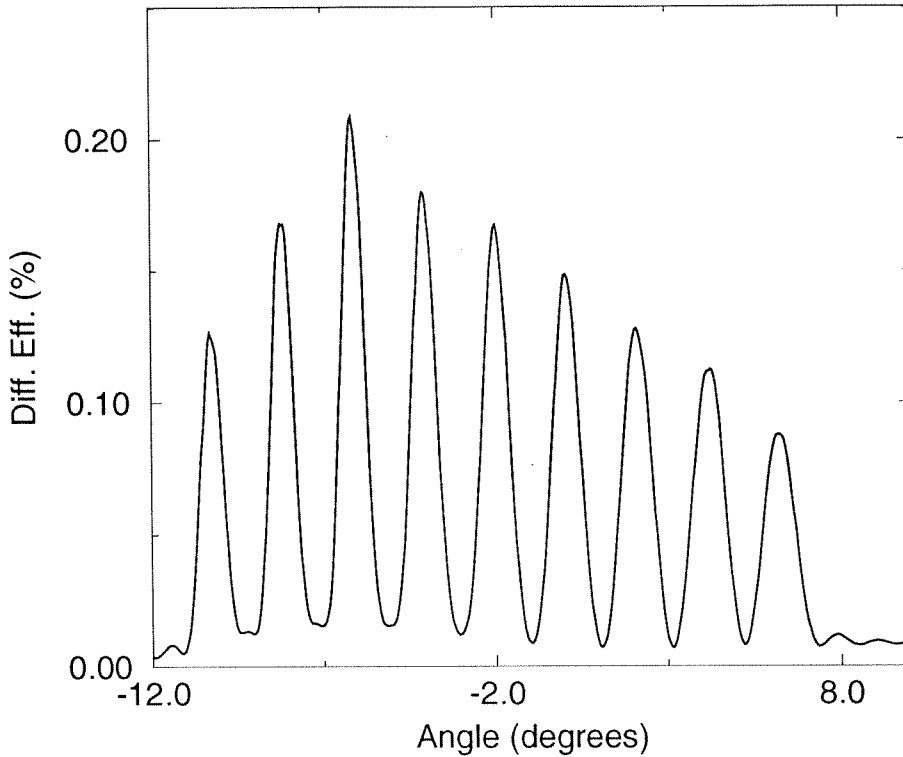


Figure 4.21: 10 image plane holograms, 1 minute apart.

time between holograms from 10 seconds as in Figure 4.18 to one minute. As the exposure time increases the dynamic range of the material is also used up as the residual monomer is polymerized. The decrease in hologram strength due to loss of dynamic range is evident in Figure 4.21. Therefore, not only is angular selectivity and exposure time important parameters for recording, but total run time is also. In order to maximize the number of holograms that can be stored, the total run time ( $T_{run}$ ) should be minimized.  $T_{run}$  depends on  $t_{setup}$  (the time to present an image to the system and change the angle) and  $t_e$  (the exposure time) which varies as  $t_o/N$  where  $t_o$  is the single hologram saturation time and  $N$  the number of holograms.  $T_{run}$  can be written as

$$T_{run} = t_{setup} \times N + t_e \times N = t_{setup} \times N + t_o. \quad (4.14)$$

Using  $N=1000$ ,  $t_{setup} = .1$  seconds, and  $t_o = 10$  seconds results in a total run time of 110 seconds. For a large number of holograms, the setup time dominates  $t_o$ .

However since  $t_{setup}$  can be equal to .1 seconds or less, the total run time even for 1000 holograms is below the 120 seconds that was experimentally demonstrated. Therefore, in practice the total run time is not the limiting factor in the number of holograms that can be stored.

The factor that does limit the number of holograms that can be stored is the thickness of the current samples. Increasing the thickness of the material, increases the angular selectivity and hence the storage capacity of the material. Scattering may eventually limit the maximum useful thickness of the material. As an example, let us assume that we can fabricate material that is 4 mm thick with sufficiently low scatter level. Then using the method presented in Reference [208] and with a maximum index perturbation of  $1 \times 10^{-3}$ , we predict that around 2,800 holograms each with .01 percent diffraction efficiency can be superimposed in this material. However fabricating thick samples with good optical quality is not going to be trivial.

#### 4.4.3 Conclusions

High quality, multiple holographic storage has been demonstrated in DuPont's HRF-150 photopolymer. Angular multiplexing qualities have been shown to be predictable using Kogelnik's two wave diffraction theory. In addition, a pre-illumination pulse can be used to record equal strength holograms with the same exposure energy per hologram. Uniform diffraction efficiency for multiple holograms over a long run time can be obtained at the expense of hologram strength due to the finite dynamic range of the material. Therefore, total run time, for sequential exposures, is an important recording parameter and needs to be minimized. The thickness of the material is currently the limiting factor in the number of holograms that can be stored. With a large dynamic range, if it is possible to increase the thickness of the material without significantly degrading the optical quality, it might be possible for thousands of permanent holograms to be recorded with adequate diffraction efficiency.

## 4.5 Thick Samples of DuPont Photopolymer

### 4.5.1 Why Thickness is Important?

The number of holograms that can be angularly multiplexed in a given holographic system is primarily a function of two parameters – the system’s angular bandwidth and the material’s dynamic range. From the section above, we were able to conclude that the photopolymer had sufficient dynamic range for more holograms than could be recorded due to angular bandwidth restrictions. Making the photopolymer films thicker can help both the dynamic range and the required bandwidth resulting in an increased system capacity.

The angular selectivity ( $\Delta\theta(L)$ ) of the holograms goes linearly with the material’s thickness. The total number of holograms that can be stored due to bandwidth restrictions is given by

$$\text{Number of Holograms} = \left( \frac{\theta_r}{\Delta\theta(L)} \right), \quad (4.15)$$

where  $\theta_r$  is the total bandwidth of system (total angle that the reference beam can be changed, and  $\Delta\theta(L)$  is the individual holograms selectivity (how much the reference angle must be changed before another hologram can be recorded). The total system bandwidth can also be increased by using fractal sampling grids [43]. By doubling the thickness of the material, the number of holograms that can be stored is doubled.

In addition to increasing the bandwidth limited number of holograms, increasing the thickness will increase the dynamic range of the material which allows more holograms to be stored. For transmission holograms the diffraction efficiency scales as the thickness squared. Thus increasing the thickness of the film allows for a larger number of holograms to be stored with the same diffraction efficiency. The practical demonstration of this idea is given by Burr et al. [50] in their long interaction length architecture (LILA). Using this increased interaction length idea, they demonstrated

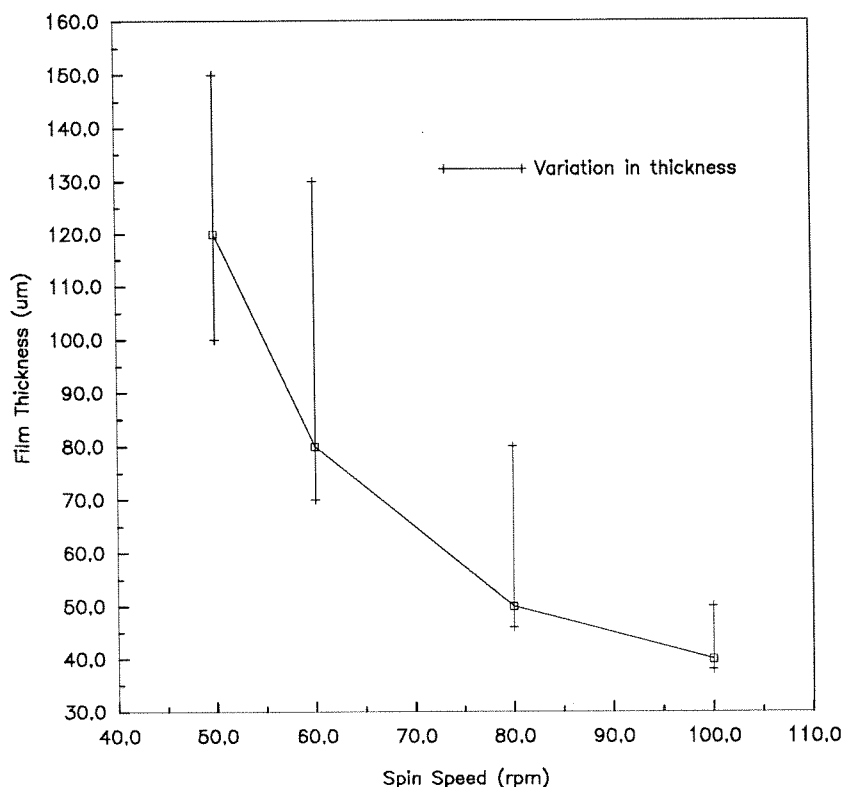


Figure 4.22: Film thickness in  $\mu\text{ms}$  vs spin speed in rpms.

the storage of 10,000 holograms in a single iron doped  $\text{LiNbO}_3$  crystal. The trade off for this increase in capacity is a less sensitive material (if the overall absorption is kept relatively low) resulting in slower recording times. Eventually other qualities of the recording material such as absorption or trap/dye density will prevent the capacity of the material from being raised arbitrarily high. Still, increasing the thickness of film from its present  $38\mu\text{m}$  thickness should increase both the film's dynamic range and, of particular interest, its angular selectivity.

### 4.5.2 Spin Coating and Casting Photopolymer Films

An obvious way to make thicker films is to spin coat glass substrates with the photopolymer. Figure 4.22 shows the film thickness vs revolutions per minute for HRF-150 photopolymer with a 50% solid solution. The solution is mixed with the sensi-

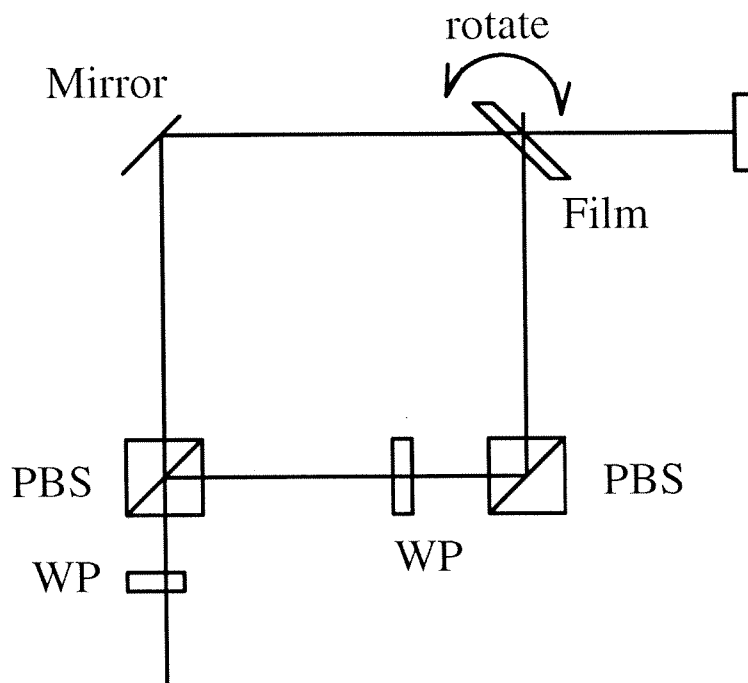


Figure 4.23: Setup for testing films: PBS – polarizing beam splitter, WP –  $1/2$  waveplate.

tizing dye and allowed to sit in the dark in order to get the bubbles out. Then the material is spun at the desired rpm for several minutes to coat the glass evenly and to dry the film. The coated film is left in the dark overnight. The long spin and waiting period is necessary to evaporate the solvents in the polymer solution. Exposing “wet” film results in either a white non-transparent plastic being formed during recording or, if the film is just slightly “wet,” the angular selectivity of the film is less than what it should be for its given thickness. Films have been made up to  $200\text{-}250\ \mu\text{m}$  thick with either single or double coating. Coating the film more than twice was not possible because the effects of surface imperfections or dust accumulating with each layer.

The films were tested using the setup shown in Figure 4.23 with plane waves incident on the film. Figure 4.24 show the angular selectivity function of a  $120\ \mu\text{m}$  thick film that was spun at 50 rpm for one hour and then dried over night.

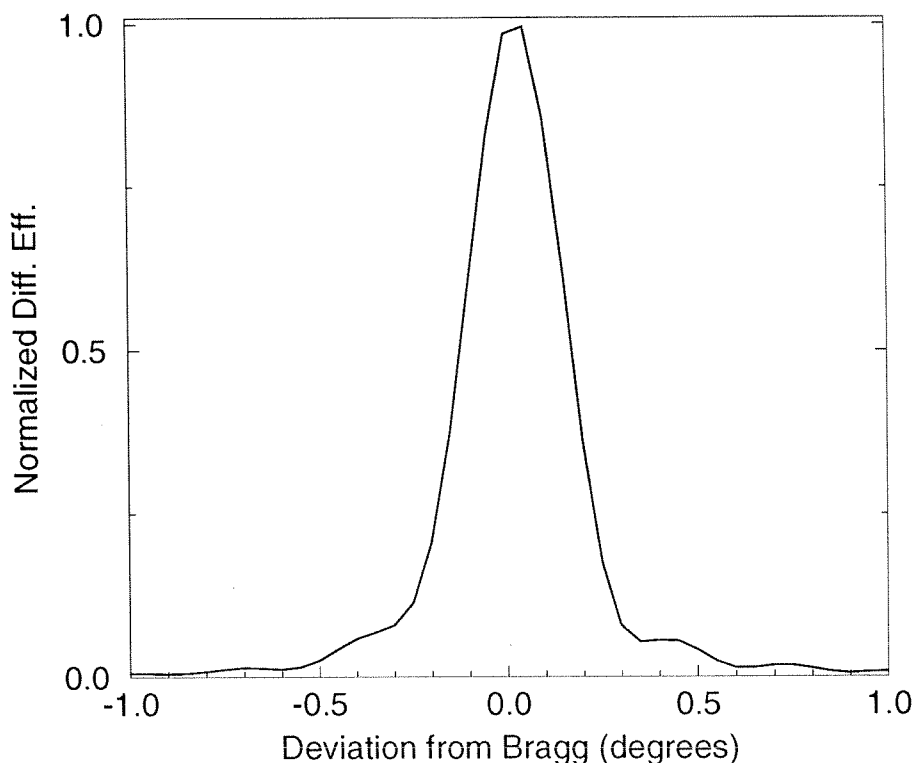


Figure 4.24: Angular selectivity for 120  $\mu\text{m}$  thick film.

The diffraction efficiency of this film was 92.8%. The biggest problems with spin coating were thickness variations and optical quality. Two films using the exact same procedure and processed one after another would result in films of different thickness. Some of this variation in the thickness is due to evaporation of the solution over time which change the film viscosity. The change in thickness is important because the film's angular selectivity and sensitivity can change which makes experiments difficult to repeat. The optical quality got worse with increasing thickness. If the films have good optical quality, such as the standard films that are purchased directly from DuPont, they can be stuck together to double the thickness. Figure 4.25 shows a plot of diffraction efficiency for 50 holograms multiplexed in an approximately 200  $\mu\text{m}$  thick sample prepared by spin coating two layers at 60 rpms. The separation between holograms is 0.4 degrees. The first couple are not present due to the fact that the pre-illumination was not sufficient to get the film into the linear region of its

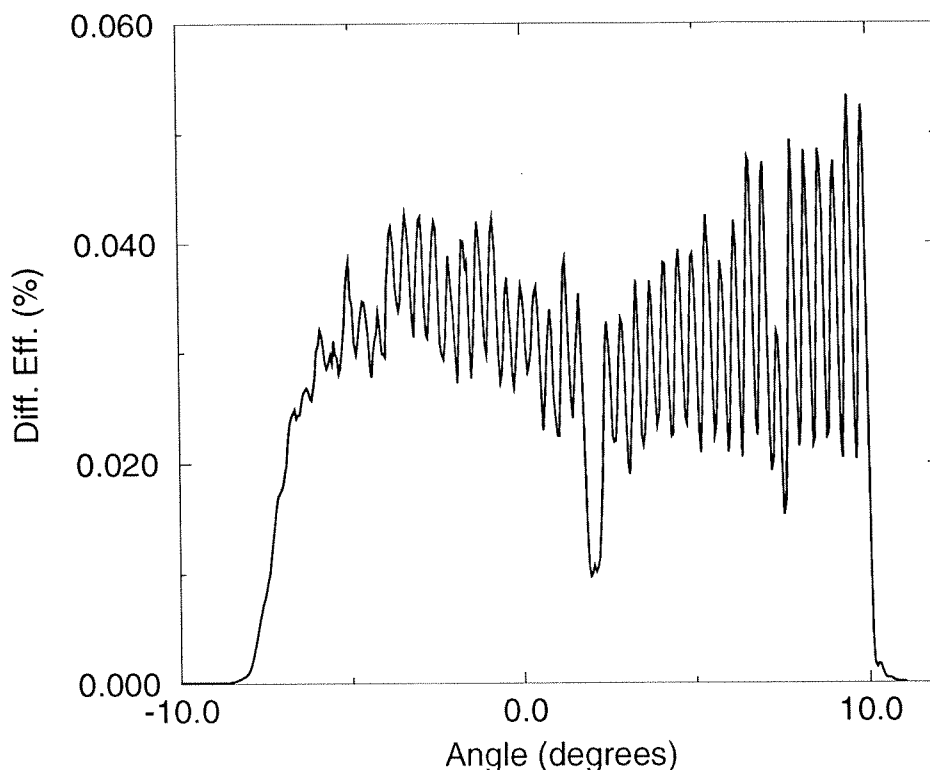


Figure 4.25: 50 holograms in a  $\sim 200 \mu\text{m}$  thick film.

exposure curve.

The other method developed for manufacturing thick films is called casting. The solution is prepared as before, but this time it is poured onto a glass plate with an o-ring glued to it to prevent the solution from spreading. The film is then left in the dark to dry. The samples can be periodically weighed to estimate when they are dry. The center of such samples will have good optical quality on both sides. This technique was developed here and later DuPont was persuaded to experiment with it. DuPont produced a  $200 \mu\text{m}$  film that looked prefect. Unfortunately, after a uniform UV exposure DuPont measured 12% scattering in the film. The scattering (DuPont calls it haze) in the standard HRF-150 films is less than 0.1%.

Figure 4.26 show the real-time development of the grating in the  $200 \mu\text{m}$  thick film that DuPont produced. Notice that the film is less sensitive since it takes more than  $30 \text{ mJ}/\text{cm}^2$  to get to the linear region on the curve compared to about  $15 \text{ mJ}/\text{cm}^2$

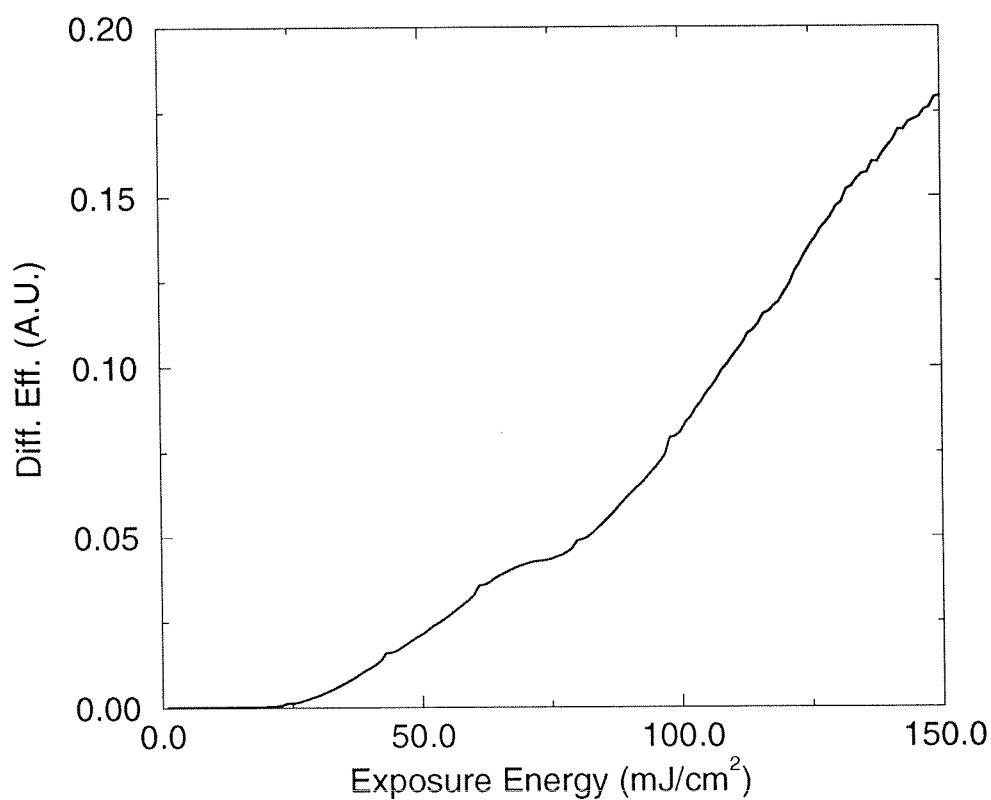


Figure 4.26: Diffraction vs exposure energy for 200  $\mu\text{m}$  thick film.

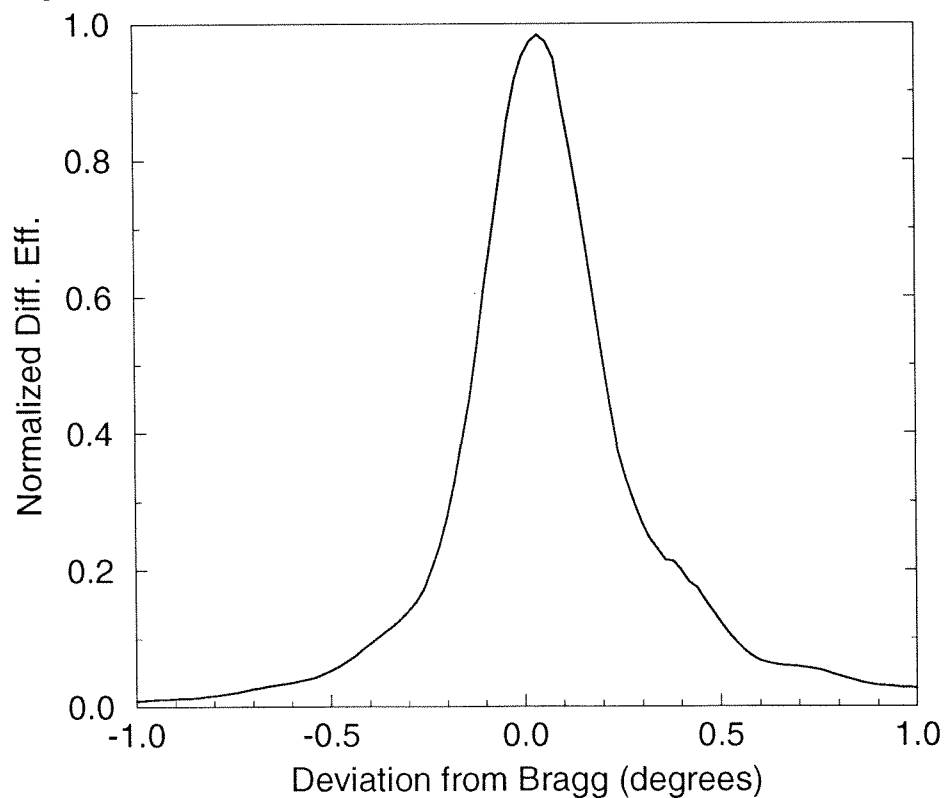


Figure 4.27: Angular selectivity for 200  $\mu\text{m}$  thick cast film.



for the standard films. Figure 4.27 shows the angular selectivity for the 200  $\mu\text{m}$  thick film. Notice that the angular selectivity is about equivalent to a sample that is about 100  $\mu\text{m}$  thick. The really bad news is scattering makes this film unusable.

Using any method, the optical quality of thick samples of HRF-150 is not satisfactory. At most, about 150  $\mu\text{m}$  thick films can be made with acceptable optical quality. It appears that it will require some work on DuPont's part to change the chemical composition of the films to made solution based photopolymers thick. Other polymer work has been done with thick materials – Bartolini et al. [111]. These use materials that have already undergone a free radical reaction to form the thick solid matrix. The photopolymer reaction then uses the residual monomers. This method results in thick material at a significant loss in sensitivity, dynamic range, and repeatability. Thus, the solution based polymers have substantially better recording characteristics.

### 4.5.3 Stratified Media

Another method of making thick materials from thin layers is to make a stratified medium. In such a medium, the thin layers are surrounded by buffer material which does not interact with the light. These thin layers of material and buffer are then stacked up to make a thick medium – sort of like a sandwich. This has been proposed before [209, 210, 211, 212], and the theory of diffraction from these stacks of material has been worked out. This particular photopolymer was used to make a stratified medium and compared to theory by Nordin et al. [213]. We also made stratified media out of the photopolymer to evaluate its potential.

Figure 4.28 shows two films stuck together with a thin mylar sheet in between. Figure 4.29 shows two films stuck together with a thick ( $\sim 50\mu\text{m}$ ) mylar sheet in between. In both cases, the overall selectivity is the same as the thin layer, but the

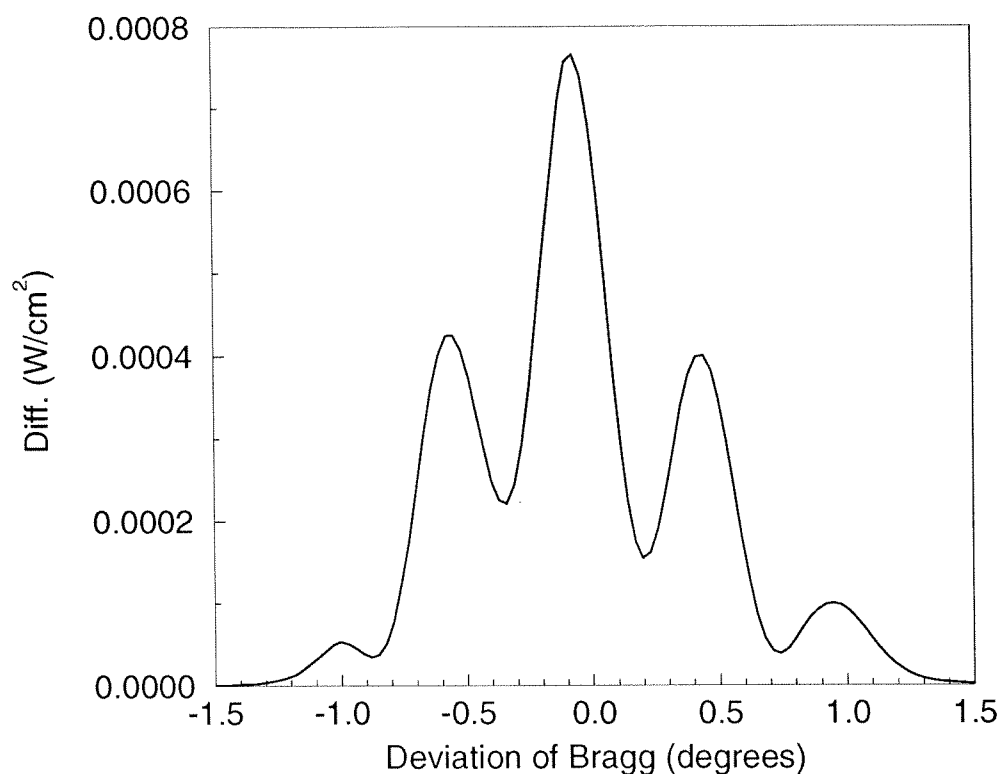


Figure 4.28: Angular selectivity 2 films stuck together with a thin mylar sheet in between.

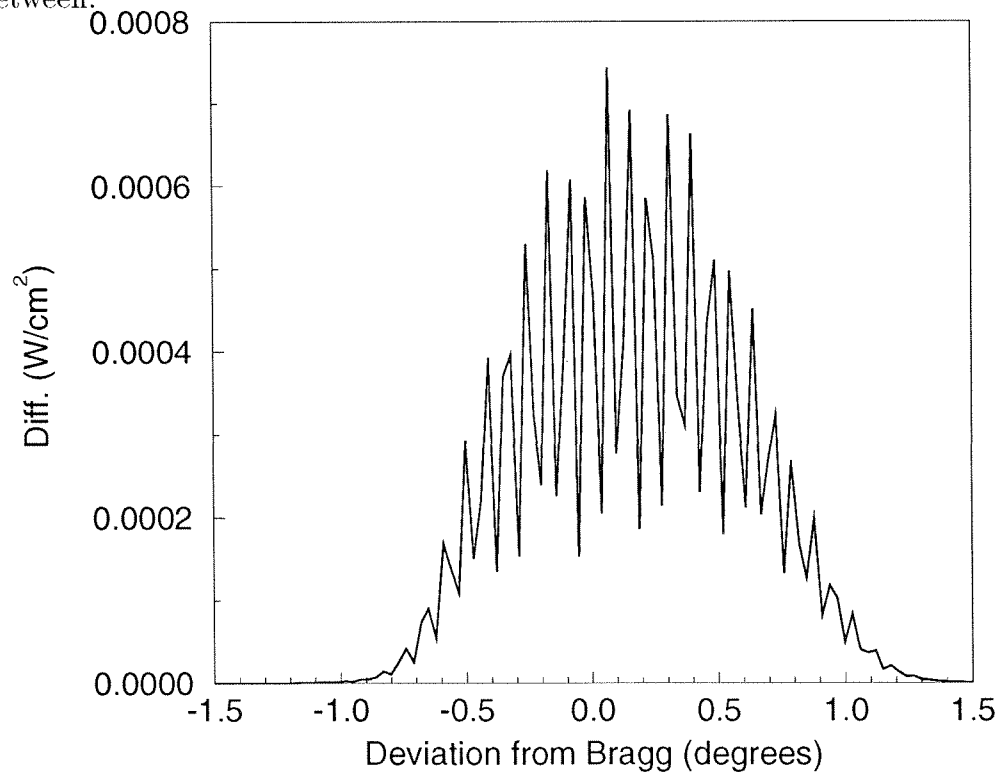


Figure 4.29: Angular selectivity 2 films stuck together with a thick mylar sheet in between.

main and sidelobes of this sinc function have nulls in them. The spacing of these nulls is very dependent on the buffer layer thickness. To try to multiplex holograms in a nulls is possible, but it is very difficult. The spacing must be maintained carefully since the other holograms also have this same selectivity function. Also the cross-talk for holograms spaced in the inner minimums of the main lobe will be much greater. Basically, stratified medium is useful for increasing the dynamic range or the material, but it probably is not a practical solution to the angular bandwidth limited problem.

## 4.6 A Method for Holographic Storage Using Peristrophic Multiplexing

### 4.6.1 Introduction

In a previous section we reported 10 angle multiplexed holograms in a  $38\mu\text{m}$  thick film [121] with diffraction efficiency of  $10^{-3}$ . Since we can typically work with holographic diffraction efficiencies on the order of  $10^{-6}$ , we have sufficient dynamic range to record significantly more than 10 holograms. The angular bandwidth limitation can be alleviated by making the film thicker [113] but scattering increases rapidly with thickness in these materials. Another method that has been previously used to increase the utilization of the available bandwidth of the system is fractal sampling grids [214, 43].

In this section we describe peristrophic (Greek for rotation) multiplexing as a solution to the bandwidth limited capacity problem. With this method the hologram is physically rotated with the axis of rotation being perpendicular to the film's surface every time a new hologram is stored. The rotation does two things. It shifts the reconstructed image away from the detector allowing a new hologram to be stored and viewed without interference, and it can also cause the stored hologram to become non-Bragg matched. In addition, peristrophic multiplexing can be combined with

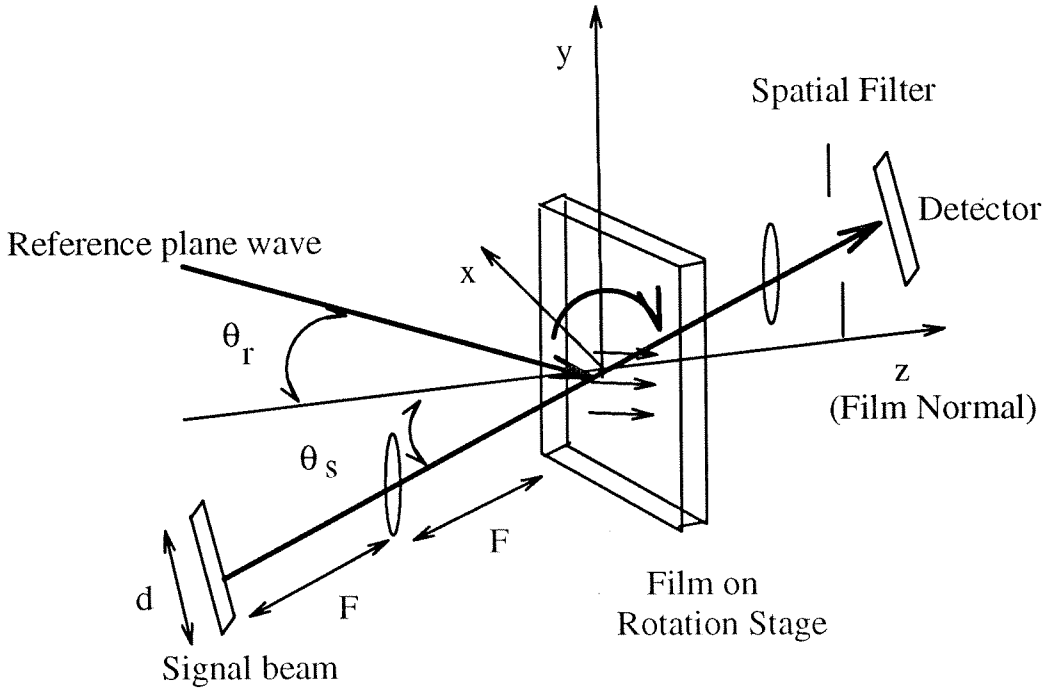


Figure 4.30: Setup for peristrophic multiplexing.

other multiplexing techniques such as angle or wavelength multiplexing to increase the storage density and with spatial multiplexing to increase the storage capacity of the system.

### 4.6.2 Theory

The setup for peristrophic multiplexing is shown in Figure 4.30. The reference plane wave (R) is incident at an angle  $\theta_r$  and the signal beam (S) is incident at an angle  $\theta_s$ , both angles measured with respect to the film's normal. Taking the center pixel of the image as the signal and neglecting any effects due to hologram thickness, the hologram transmittance can be written as

$$R^* S = e^{-i2\pi \frac{\sin \theta_r}{\lambda} x} e^{-i2\pi \frac{\sin \theta_s}{\lambda} x} . \quad (4.16)$$

The hologram is then rotated by  $d\theta$  about the center of the  $x$ - $y$  plane as shown in Figure 4.30. Assuming the rotation is small, this results in the coordinates being

transformed according to:  $x' \approx x - yd\theta$ , and  $y' \approx y + xd\theta$ . Substituting these relations into Eq. 4.16, the hologram be expressed in terms of the unrotated coordinates  $(x, y)$

$$R^*S = e^{-i\frac{2\pi \sin \theta_r x}{\lambda}} e^{-i\frac{2\pi \sin \theta_s x}{\lambda}} e^{-i\frac{2\pi(\sin \theta_s + \sin \theta_r)d\theta y}{\lambda}}. \quad (4.17)$$

After multiplying by  $R$  and Fourier transforming, the last term in Eq. 4.17 results a shift in the image. The rotation required to translate the image out of the detector aperture is approximately given by,

$$d\theta \geq \frac{\frac{d}{F}}{\sin \theta_s + \sin \theta_r}, \quad (4.18)$$

where  $d$  is the size of the image at the detector plane and  $F$  is the focal length of the lens used. For image plane holograms, the expression is

$$d\theta \geq \frac{\frac{2\lambda}{\delta}}{\sin \theta_s + \sin \theta_r}, \quad (4.19)$$

where  $1/\delta$  is the highest spatial frequency in the image. For image plane holograms the filtering must be done at the Fourier plane of the system. Similar expressions can be found for Fresnel holograms [215]. Notice that this method can be combined with other volumetric multiplexing methods to further increase the storage capacity.

The Bragg selectivity, assuming the reference is given by  $R = e^{-i(\frac{2\pi \sin \theta_r}{\lambda}x + \frac{2\pi \cos \theta_r}{\lambda}z)}$  and the signal given by  $S = e^{i(\frac{2\pi \sin \theta_r}{\lambda}x + \frac{2\pi \cos \theta_r}{\lambda}z)}$ , can be calculated using the Born and paraxial approximations and integrating over the volume of the hologram. Assuming that the transverse  $(x, y)$  dimensions of the film are much larger than the bandwidth of the images, the Bragg selectivity can be shown to be

$$d\theta = \sqrt{\frac{2\lambda}{t} \left( \frac{\cos \theta_s}{\sin \theta_r (\sin \theta_s + \sin \theta_r)} \right)}, \quad (4.20)$$

where  $t$  is the thickness of the material. Using  $\lambda = 488\text{nm}$ ,  $t = 38\mu\text{m}$ , and  $\theta_s = \theta_r = 30^\circ$  results in a selectivity of about  $10^\circ$ . The Bragg matching requirement is the dominant effect if  $\frac{d}{F} > \sqrt{2\lambda \cos \theta_s (\sin \theta_s + \sin \theta_r) / t \sin \theta_r}$ . For most material thicknesses,

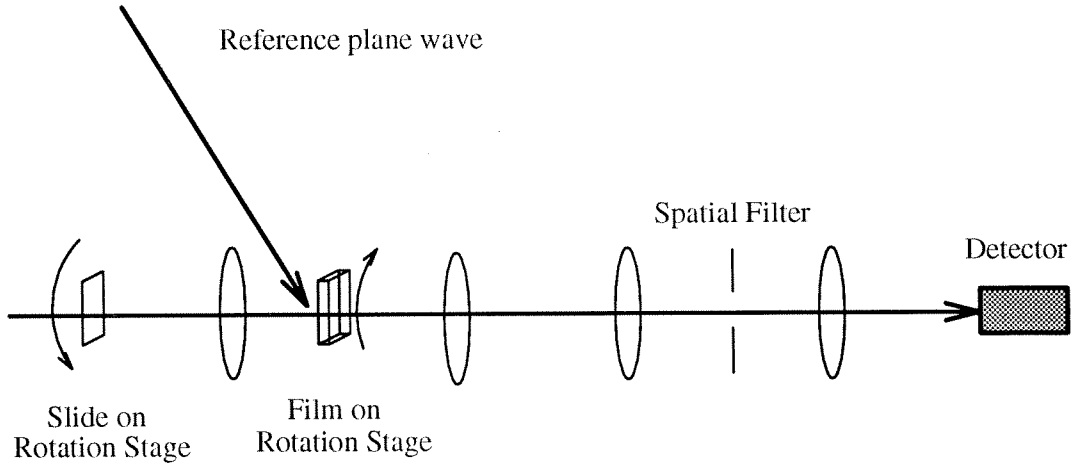


Figure 4.31: Experimental setup for peristrophic multiplexing.

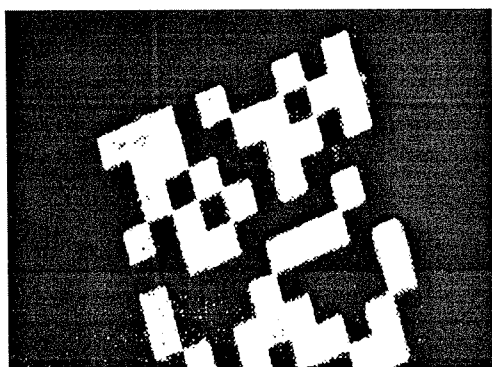
the Bragg matching criterion determines the required rotation for peristrophic multiplexing. In our experiments, because the thickness of the film is only  $38\mu\text{m}$ , the image could be filtered out before the gratings become non-Bragg matched.

### 4.6.3 Experimental Results

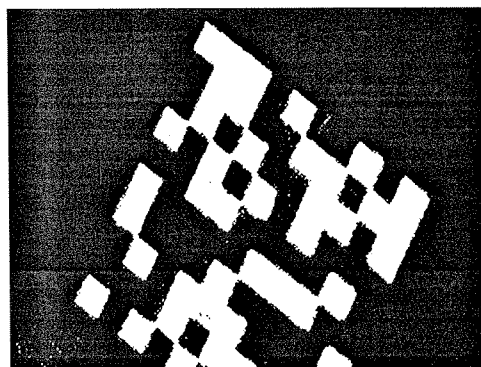
Peristrophic multiplexing was demonstrated using DuPont's HRF-150 photopolymer. The setup in Figure 4.31 was used to record 100 holograms in the  $38\mu\text{m}$  thick film. The image stored was a random black and white pattern on a slide. The slide was placed on a rotation stage and rotated in the opposite direction that the film was rotated to distinguish individual images from each other. The reference plane wave was incident at  $32^\circ$  from normal and the signal at  $29^\circ$  from the normal. The reference beam intensity was  $3.76\text{ mW/cm}^2$  and the signal beam had  $449\text{ }\mu\text{mW}$  in it. The film was rotated by  $1.7^\circ$  degrees between holograms to enable the other holograms to be filtered out. The image was rotated by  $-2^\circ$  for each hologram. The exposure time per hologram was 0.12 seconds and there was a two second delay between holograms to allow the rotation stages to rotate to a complete stop. Figure 4.32 shows some of the reconstructed holograms, and Figure 4.33 shows their diffraction efficiency. The

variations in diffraction efficiency are due in part to exposing the material through a mylar cover sheet. The mylar is weakly birefringent and therefore slightly depolarizes the beams by an amount that depends on orientation of the film.

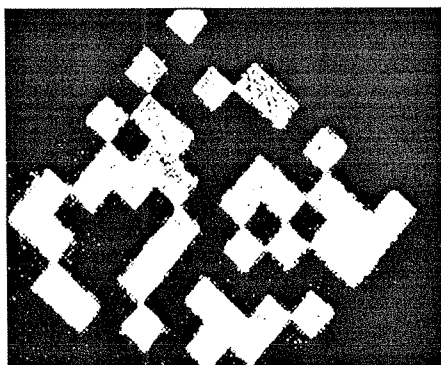
The experimental setup shown in Figure 4.34 was used to demonstrate both peristrophic and angle multiplexing. Notice that a second rotation stage was added to rotate the film around a vertical axis as well as around the film's normal. This makes it possible to combine peristrophic and angle multiplexing. The film was located a significant distance from the Fourier plane so that the signal beam was approximately uniform. For each peristrophic position, multiple holograms are stored using standard angle multiplexing by rotating the medium. A spatial light modulator (SLM) was used to present images (cartoons) to the system. Each frame is numbered according to the sequence in which they were stored. The reference and signal beams were initially incident at  $\pm 30^\circ$  from the film's normal. The reference beam intensity was  $1.1 \text{ mW/cm}^2$  and the signal beam had  $300 \text{ } \mu\text{W}$  in about a  $1 \text{ cm}$  by  $0.5 \text{ cm}$  area. The film was rotated in-plane by  $3^\circ$  between each set of angle multiplexed holograms to enable the other holograms to be filtered out. Eq. 4.18 predicts a required rotation of about  $9^\circ$  for Fourier plane hologram while Eq. 4.19 predicts about  $1.7^\circ$  rotation for image plane. The  $3^\circ$  was experimentally observed for the in-between (Fresnel) case we used. Each angle multiplexed hologram was also separated by  $3^\circ$ . The initial exposure time was 0.11 seconds, but starting at hologram number 26, each hologram was exposed for 0.005 seconds longer than the previous hologram to correct for the lost sensitivity due to run time [2]. There was a 1.5 second delay between holograms to allow the rotation stages to completely stop. Figure 4.35 shows five of the 295 holograms stored in the polymer by peristrophic multiplexing 59 times and storing 5 angle multiplexed holograms with each peristrophic position. The diffraction efficiency of the 295 holograms is plotted in Fig. 4.36. The average efficiency was  $\sim 4 \times 10^{-6}$  and



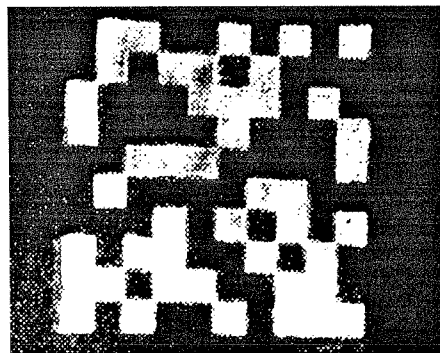
1<sup>st</sup> Hologram



25<sup>th</sup> Hologram



75<sup>th</sup> Hologram



100<sup>th</sup>

Figure 4.32: Some of the 100 reconstructed holograms.



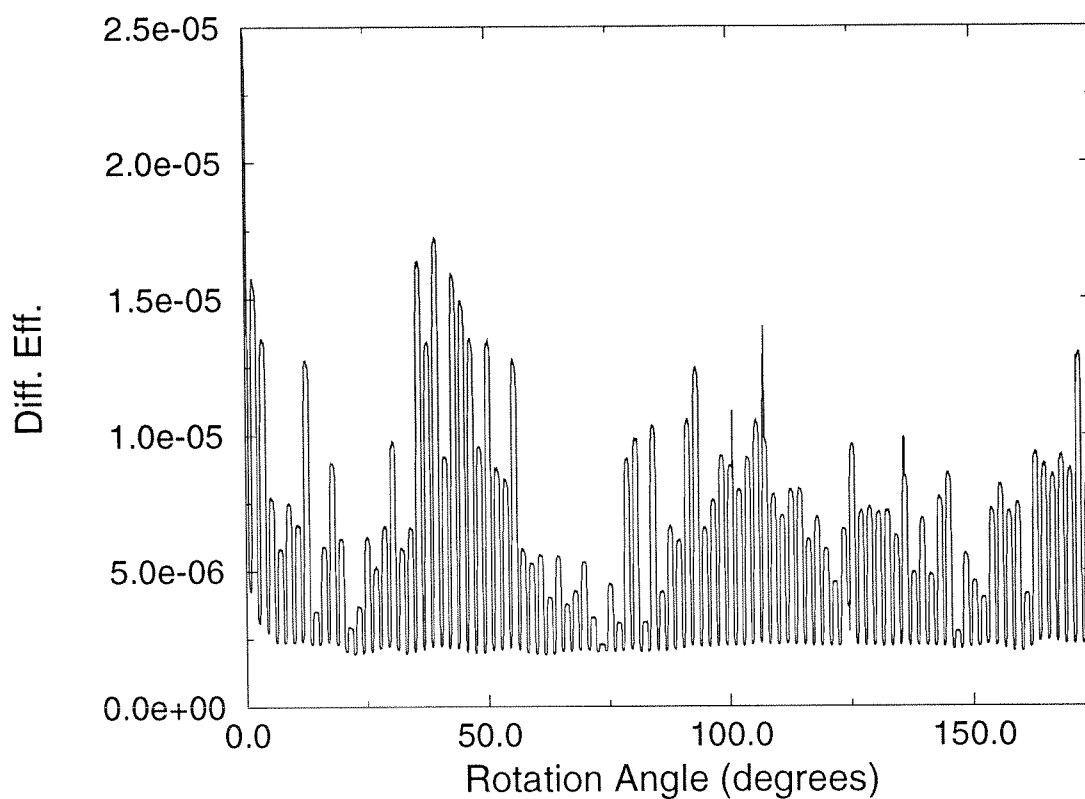


Figure 4.33: Diffraction efficiency of 100 holograms stored in  $38\mu\text{m}$  thick film.

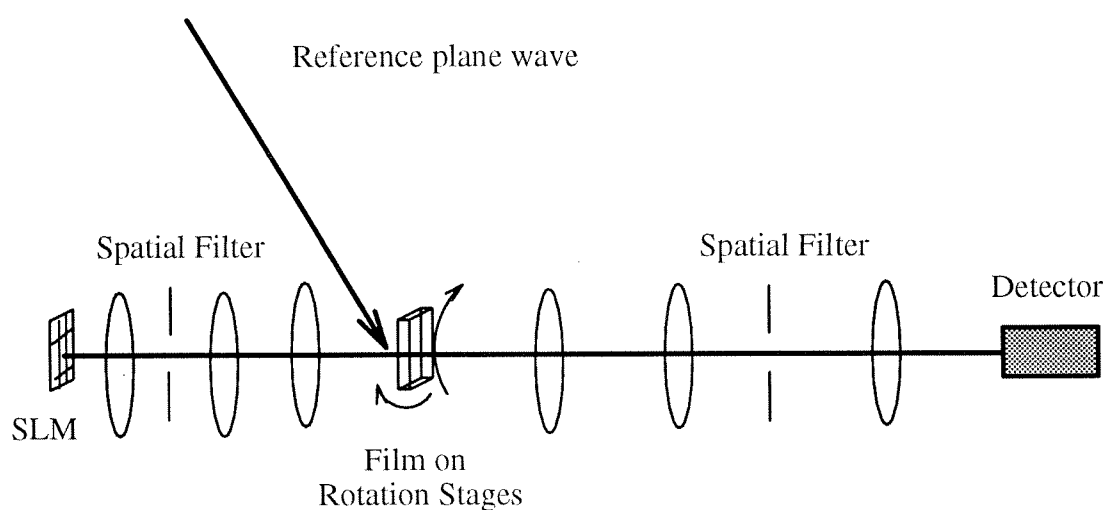


Figure 4.34: Setup for both peristrophic and angle multiplexing.

the variations are primarily due to variation in the average intensity of the frames. In separate experiment, we stored equal amplitude plane wave holograms and observed a decrease in diffraction efficiency proportional to  $1/M^2$  [216].

Previously we stored  $M = 10$  holograms with roughly  $10^{-3}$  diffraction efficiency [121] limited by the angular bandwidth of the optical system. Peristrophic multiplexing made it possible to store  $M = 295$  holograms with a diffraction efficiency of  $\sim 4 \times 10^{-6}$ . Assuming one bit per pixel, the above experiment results in a storage density of roughly 1 bit/ $\mu\text{m}^2$ . Thus, peristrophic multiplexing allowed for almost two orders of magnitude increase in the storage capacity of the DuPont photopolymer and changed the limiting factor from the angular bandwidth of the optical system to the dynamic range of the material.

1<sup>st</sup> Hologram103<sup>th</sup> Hologram160<sup>th</sup> Hologram219<sup>th</sup> Hologram295<sup>th</sup>

Figure 4.35: Some of the 295 reconstructed holograms.

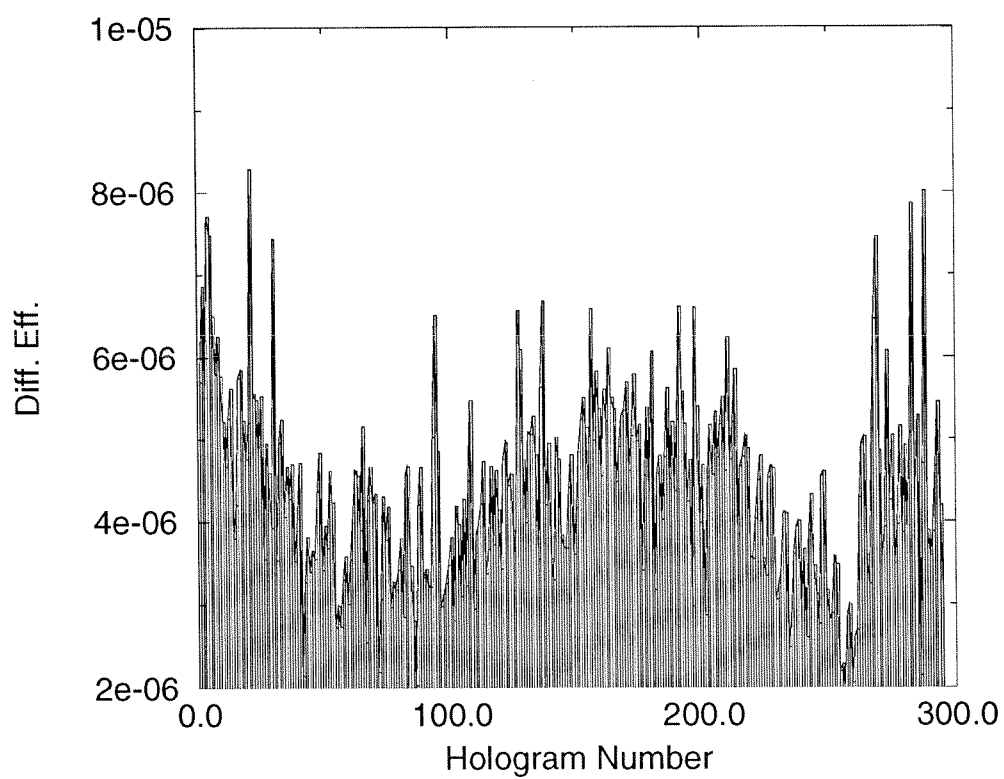


Figure 4.36: Diffraction efficiency of 295 holograms stored in  $38\mu\text{m}$  thick film.

## Chapter 5

# 3-D Photopolymer Disks for Correlation and Data Storage

## 5.1 Correlator

### 5.1.1 Introduction

In recent years we have seen the implementation of several pattern recognition systems that use arrays of optical correlators to realize a variety of algorithms (e.g., template matching, radial basis functions, and neural networks). An array of correlators can be implemented by spatially multiplexing different filters on a 2-D medium, typically in a disk configuration [217, 218, 219, 220]. Alternatively, 3-D holographic storage can be used by superimposing multiple holograms utilizing either wavelength [176] or angular [221, 222] multiplexing to form the correlation between the input and any stored reference. Finally, spectral and time domains can be used to store multiple filters in 2-D spectral hole burning media [223]. In this chapter we present a method that combines angularly multiplexed storage in 3-D media with spatial multiplexing to form a 3-D disk [196]. The 3-D disk allows us to store a very large number of gray scale templates (in excess of 100,000 ) whereas the disk rotation proves convenient for implementing the necessary shift in one of the two dimensions for the calculation of the image correlation function. The architecture we describe can be

implemented with any holographic storage medium that can be fabricated into a thick slab (to allow volume holography) with sufficient area (to make a sizable disk). For instance, photorefractive crystals can be used. The 3-D disk used in the experimental demonstration in this paper, is constructed with the HRF-150 Dupont photopolymer [112].

### 5.1.2 Theory of Operation

When a thick hologram is used to store the Fourier transform filter in a VanderLugt correlator, the shift invariance is destroyed in one of the two dimensions [224]. As a result, the output of such a system is simply one line of the 2-D correlation function. We can use angularly multiplexed holograms in a 3-D volume to store multiple templates and use the second dimension of the output plane to display in parallel one line of each of the 2-D correlation patterns between the input and the stored templates. By translating the input image, the 2-D correlation patterns are generated line by line with time [218, 221, 222]. Here we describe an alternative solution in which we translate the holographic medium instead. The reference holograms are stored by Fourier transforming in one direction and imaging in the other. Since the input is imaged onto the hologram in one dimension, translating the hologram produces the desired relative shift between input and reference that is necessary to compute the correlation function in 2-D. If the holographic medium is a 3-D disk, then this translation is provided by the disk rotation. Moreover, in this architecture, we can increase the number of templates by storing multiple sets of angle multiplexed holograms at different locations on the disk. Disk rotation can also be conveniently used to access the different locations in sequence.

The optical setup is shown in Figure 5.1. The image  $f(x,y)$  to be stored is presented to the system on a SLM ( an Epson LCTV). The first lens takes the Fourier transform of the image presented on the SLM. At the Fourier plane the image is filtered both

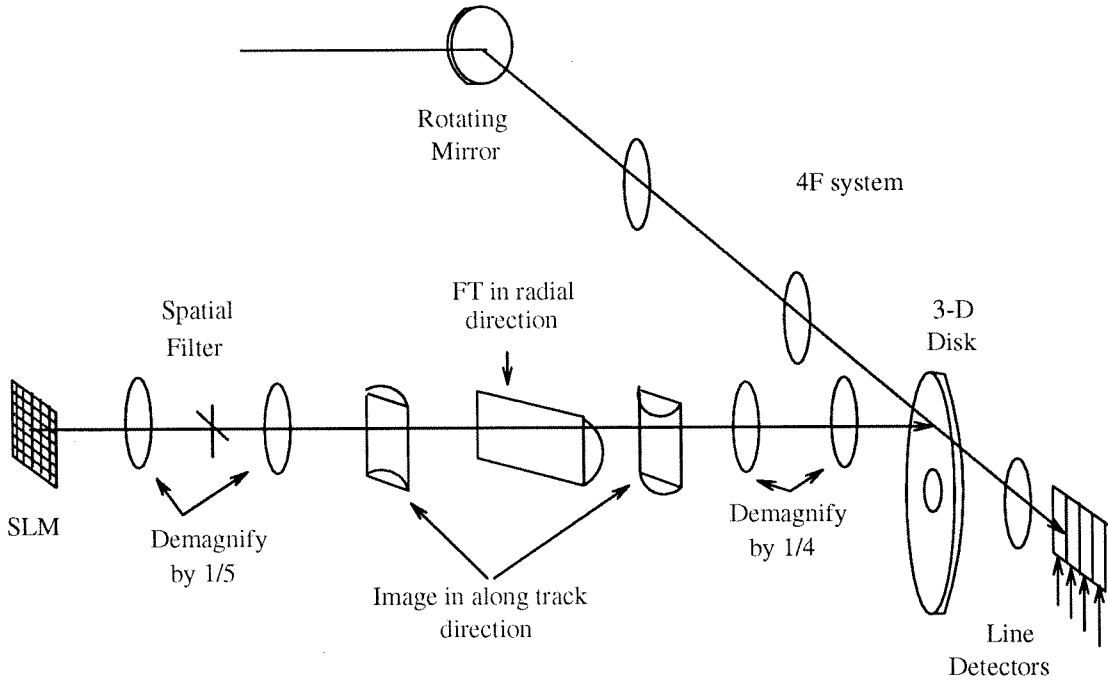


Figure 5.1: Optical setup for 3-D disk-based correlator.

to edge enhance and to remove the diffraction orders of the SLM. The second lens transforms this filtered signal back into a filtered and demagnified version of the original image. Using three cylindrical lenses, the image is imaged in the along “track” ( $x$ ) direction and Fourier transformed in the radial ( $y$ ) direction. The spherical lens, after the cylindrical lenses, demagnify the 1-D Fourier transform and 1-D image further. This pattern is then interfered with a plane wave with spatial frequency  $\mu$  and stored in an optically thick photopolymer film on a glass optical disk. This results in the image of  $f$  in the  $x$  direction and complex conjugate of Fourier transform of  $f$  in the  $y$  direction ( $f^*(M_1 x', M_2 v)$ ) being stored, where  $v$  is the spatial frequency in the  $y$  direction. More holograms can be multiplexed at this location on the disk by changing  $\mu$  using the rotating mirror and a 4F system of lenses to change the incident angle of the reference beam without changing the beam location. The hologram can

be written as

$$T(x', y') \approx f^*(M_1 x', M_2 v) e^{j2\pi \mu x'}, \quad (5.1)$$

where the  $M$ 's are magnification factors.

To generate the correlation function between  $f$  and another function  $g$ , input  $g(x, y)$  on the SLM. After going through the same optics before the disk,  $g(x, y)$  is mapped into  $g(M_1 x', M_2 v)$ . The signal after the disk is given by

$$g(M_1 x', M_2 v) f^*(M_1 x', M_2 v) e^{j2\pi \mu x'}. \quad (5.2)$$

This signal is then Fourier transformed by a spherical lens, with focal length  $F$ , and detected a distance  $2F$  away from the disk. Ignoring the reconstructed plane wave that ends up uniquely shifting (for each angle multiplexed hologram) the result on the detector plane, the signal - to within angular selectivity limits - is given by

$$S(x_d, y_d) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(M_1 x', M_2 v) f^*(M_1 x', M_2 v) e^{-j\frac{2\pi}{\lambda F}(x' x_d + y' y_d)} dx' dy'. \quad (5.3)$$

Detecting the signal at  $x_d = 0$  and taking the Fourier transform in the  $y$  direction results in the signal becoming:

$$S(O, y_d) \approx \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(M_1 x', y'') f^*(M_1 x', y'' - \frac{y_d}{M_2}) dx' dy''. \quad (5.4)$$

This is an inner product in the  $x$  direction and a correlation in the  $y$  direction or one line of the full 2-D correlation function of  $f$  and  $g$ .

Using disk rotation as approximately a linear shift ( $\theta(t)$ ) in  $x$  direction generates the full 2D correlation function with time.

$$S(O, y_d, t) \approx \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(M_1 x', y'') f^*(M_1 x' + \theta(t), y'' - \frac{y_d}{M_2}) dx' dy''. \quad (5.5)$$

Since intensity at detector is  $I(0, y_d, t) = SS^*$ , the intensity detected is equal to the square of the correlation function. With a thick medium, multiple holograms can be recorded at a single spot on the disk using angle multiplexing. Therefore, different



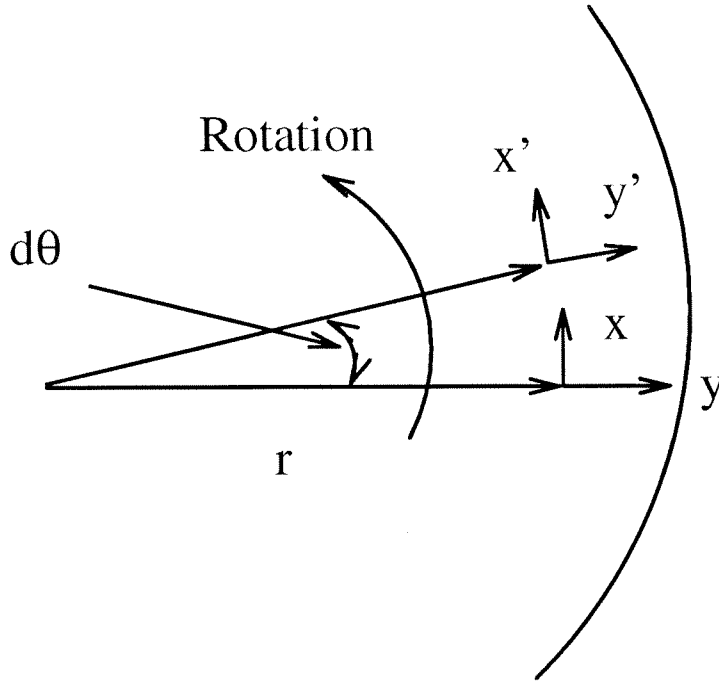


Figure 5.2: Coordinates for disk rotation by  $d\theta$  in counterclockwise direction.

2-D correlation functions for each hologram multiplexed at a spot can be generated one line at a time and detected in parallel.

### 5.1.3 Disk Rotation Effects

Hologram rotation during reconstruction results in a shift at the correlation detection plane. To analyze this effect, a cylindrical wave with origin at  $z = 0$  and  $x_o$  is the signal beam. The coordinates for disk rotation are shown in Figure 5.2. This signal beam is further simplified to a plane wave  $S \sim e^{-j2\pi\beta y}$  with spatial frequency  $\beta = \frac{\sin\theta_s}{\lambda}$ , since the cylindrical wave in the  $x$  direction is a point on the film that forms the inner product ignoring the effects of thickness. This signal beam is stored with a plane wave  $R = e^{j2\pi\alpha x}$  with spatial frequency of  $\alpha = \frac{\sin\theta}{\lambda}$  in the  $x$  direction. In the expressions above,  $\theta$  is the angle between the normal to the disk and the reference beam, and  $x, y$  are the coordinates at the disk ( $x$  along “track” direction,  $y$  radial

direction). The recorded hologram is

$$RS^* = e^{j2\pi\alpha x} e^{j2\pi\beta y}. \quad (5.6)$$

The hologram is then rotated by  $d\theta$  about a center of rotation in the  $x$ - $y$  plane a distance  $r$  away from the center of the  $x, y$  coordinates as shown in Figure 5.2. This rotation results in the coordinates being transformed according to the formulas:

$$\begin{aligned} x' &\doteq x - (y + r)d\theta, \\ y' &\doteq y + xd\theta. \end{aligned} \quad (5.7)$$

Substituting these into equation 5.6, the hologram can now be written as

$$RS^* = e^{j2\pi\alpha(x-(y+r)d\theta)} e^{j2\pi\beta y} e^{j2\pi\beta d\theta x}. \quad (5.8)$$

Multiplying by  $S$  to generate the 1-D inner product and 1-D correlation, expand and simplify, results in

$$SRS^* = e^{j2\pi\alpha x} e^{-j2\pi\alpha d\theta y} e^{j2\pi\beta d\theta x}. \quad (5.9)$$

In the above expression after the Fourier transforming lens, the factors are:

$e^{j2\pi\alpha x}$  is the reconstructed plane wave without rotation – the desired signal.

$e^{-j2\pi\alpha d\theta y}$  corresponds to a shift in  $y$  direction on the correlation plane.

$e^{j2\pi\beta d\theta x}$  corresponds to a shift in  $x$  direction on the correlation plane.

Since in the experimental setup  $\alpha \gg \beta$ , we expect the correlation to shift in the  $y$  direction on the detector as the disk is rotated.

#### 5.1.4 Experimental Correlation

Using the system shown in Figure 5.1, 300 transmission holograms were stored on one 5 cm radius “ring” of a holographic (6 cm radius) 3D disk. Using DuPont’s HRF-150 film and exposing with 488 nm light, three holograms (image plane in the

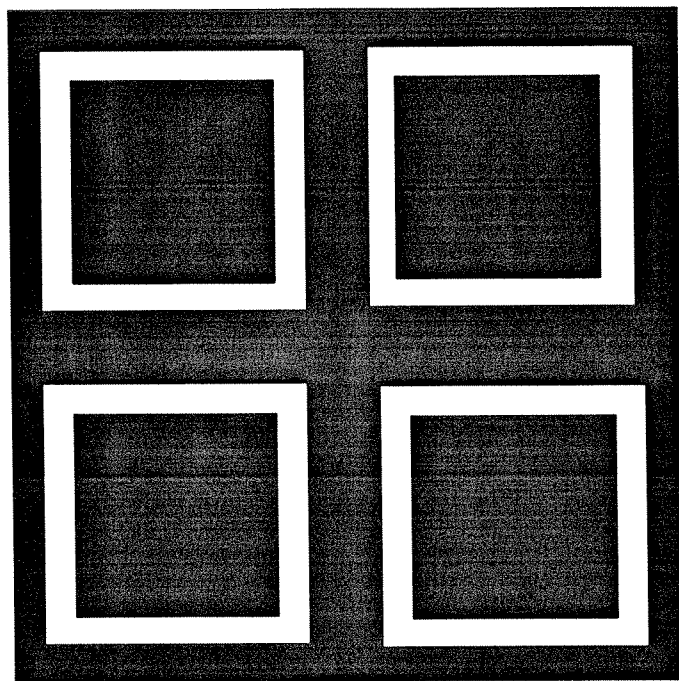


Figure 5.3: Image of four boxes that were used to test the correlator.

along “track” direction and Fourier transformed in the radial direction) were angle multiplexed at a given spot on the disk using the methods presented in Chapter 4. The disk was then rotated to expose a total of 100 different spots; at each spot, three holograms were recorded for a total of 300 holograms stored. The object size was approximately .9x.6 mm with 13  $\mu\text{W}$  intensity recorded with a 1 cm x 1.5 mm plane wave reference beam with 300  $\mu\text{W}$  intensity. After each spot had been sensitized with pre-illumination with the reference beam for 25 seconds, each hologram was exposed for 10 seconds. Disk rotation between spots was  $3.5^\circ$  and the angular separation of the three multiplexed holograms was  $1.5^\circ$ . The holograms were made with an image of four boxes as shown in Figure 5.3. The peak diffraction efficiency for the three hundred holograms is shown in Figure 5.4. The 3-D disk was made by pressing (laminating) the 38  $\mu\text{m}$  thick photopolymer film onto the glass disk. In addition, spin coating the photopolymer onto the disk is possible.

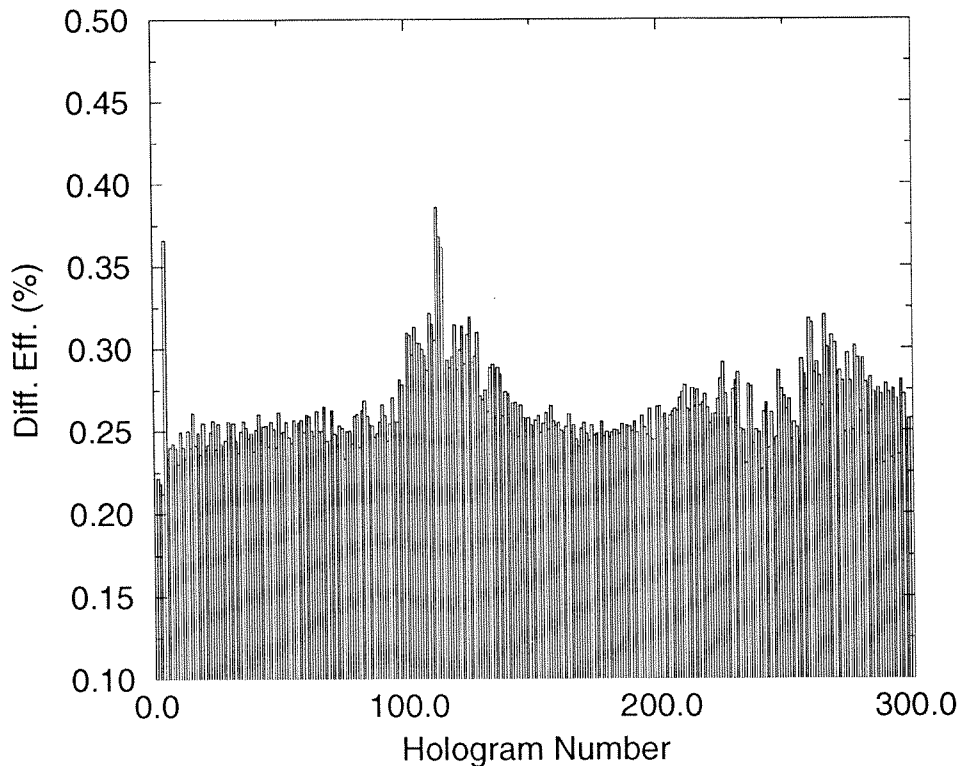


Figure 5.4: Peak diffraction efficiency of 300 holograms – 3 holograms at a spot with 100 spots on a ring of the 3-D disk.

To test the correlator, an experimental autocorrelation squared of the four boxes was measured using the correlator setup in Figure 5.1. Figure 5.5 is an experimental plot of the autocorrelation function squared of the four boxes of Figure 5.3 and roughly agrees with the theoretical autocorrelation of this image shown in Figure 5.6 (both plots are linear).

Figure 5.5 was generated by storing the image on the disk as explained above, and then presenting the image to the system. The correlation plane was sampled by a line detector. The detector was read out, the disk was rotated, and the detector was read out again. This procedure was repeated until the entire 2-D autocorrelation function squared was generated. Since the boxes occupy a large portion of the input plane (SLM), Figure 5.5 demonstrates that the correlation ability exists across most of the input plane. In addition, Figure 5.7, is the gray scale image looking straight down from the top of Figure 5.5. Notice the shift in peak location due to disk/hologram

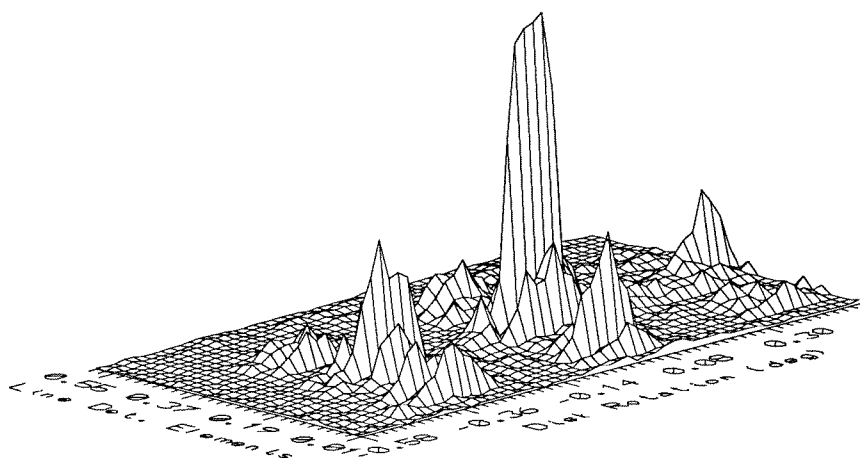


Figure 5.5: Experimental autocorrelation squared of the test image generated using the optical disk-based correlator.

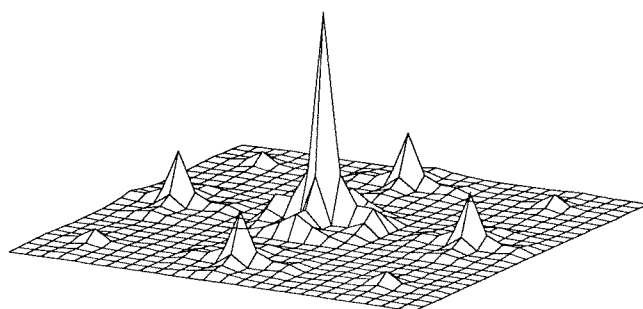


Figure 5.6: Theoretical autocorrelation squared of the test image.

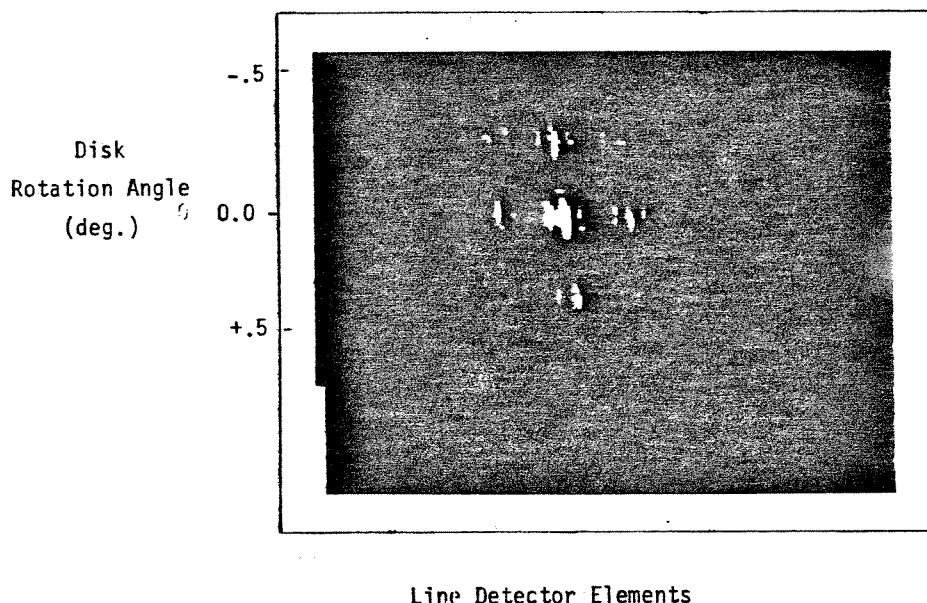


Figure 5.7: Experimental gray scale representation of autocorrelation squared.

rotation. The shift produced by the experimental setup was in the  $y$  direction as predicted.

The correlator described above eliminates the need for fast SLM's; instead, the processing bottle neck is shifted to the detectors. The detectors need to read out the correlation peaks. The largest peak then needs to be found and the address associated with it (pixel and line detector location) reported out. This means that a tremendous amount of information must be quickly read out and moved off chip. To alleviate this problem, we have tested prototype chips that do winner-take-all over multiple line detectors and report out the winner's address and magnitude. Figure 5.8 shows a picture of a chip that consists of a row of eight pixels on the left that go to a winner-take-all circuit (in the center of the picture) and then to a static encoder (far right) to report out the winning pixel address and magnitude. This chip was built through a  $2\mu\text{m}$  p-well MOSIS process and implements the desired functions for

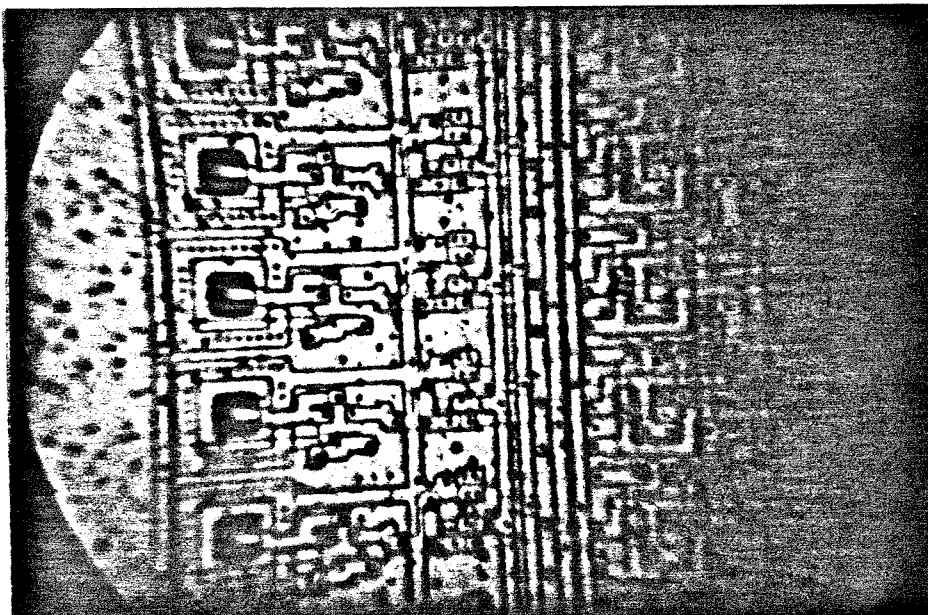


Figure 5.8: Picture of detector chip for correlator.

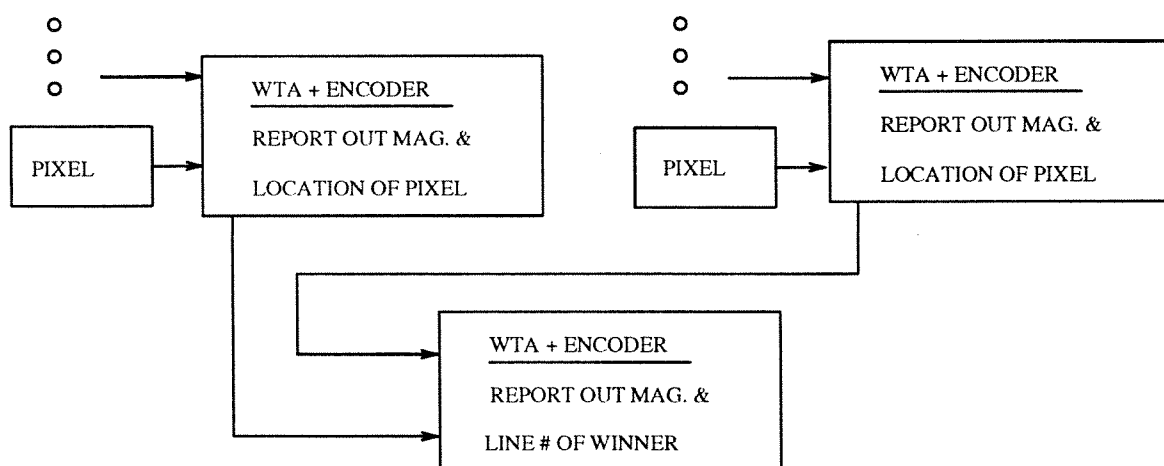


Figure 5.9: Box diagram of detector chip for correlator.

a single line of pixels. Figure 5.9 shows a box diagram of the detector circuit used for peak detection. A chip that does both pixel and multiple line WTA and then reports out the address and the magnitude of the winner was also built. These chips have the potential to increase the correlation rate by an order of magnitude.

### 5.1.5 Correlator Summary

We have demonstrated a multi-channel disk-based optical correlator using both angular and spatial multiplexing of volume holograms. The full 2-D correlation function of a presented image against stored images is generated by using disk rotation to shift one image over the other, thus reading out one line of each correlation function in parallel. 300 holograms, in the required 1-D image and 1-D Fourier transformed format, were stored on a single ring of an optical disk. The correlation generation was experimentally verified and compared to theory. Also, custom detector arrays were built and tested that find the peak on chip and report out its location which dramatically reduces how fast information is required to be read off chip.

The current system stores three holograms at a single spot and uses most of one ring around the disk resulting in 300 holograms stored. The effective thickness of the material was about 38  $\mu\text{m}$ . Using the spin coating technique, a photopolymer film thick enough ( $\approx 350 \mu\text{m}$ ) to store more than 50 holograms in one spot has been made. Using the whole area of the disk and leaving enough blank space between spots, approximately 3000 spots can be stored on a 5 cm radius disk. This results in a  $\sim 150,000$  hologram storage capacity. The speed of the current correlator is limited by the speed of the linear detectors. A line detector is needed for each angularly multiplexed hologram position. For example, if 50 holograms are multiplexed at each spot then 50 line detectors are needed. Table 5.1 shows the capacity and speed of a disk that can be realized with off the shelf technology. For example, small parallel detector arrays, SLMs, and photopolymer films that can use existing red light sources,



Characteristic	Value	Limiting Factor
Pixels ( $N$ )	500 x 500	SLM
# Holograms/spot	60	Thickness (L)
Total stored images	$\sim 3 \times 10^5$	12cm diameter & L
Correlation Rate ( $R$ )	$\sim 10^4$	30 MHz detectors & L
Integer Ops/sec ( $N^2 \times R$ )	$10^{14} - 10^{15}$	detectors & L

Table 5.1: Disk-based Correlator Parameters

are all currently available. Using the custom detector arrays similar to what has already been demonstrated, the correlation rate could be increased by an order of magnitude.

## 5.2 Data Storage

### 5.2.1 Introduction

Holographic storage of data in 3-D materials such as photorefractives or photopolymers can provide large capacities and fast parallel readout of the stored information. These memories were investigated in the 1960's and 1970's, but they never found commercial applications. Recently, with improvements in materials, light sources, detectors, and the invention of spatial light modulators, holographic storage has again arisen as a possible mass storage candidate. The theoretical upper limit on the storage density is  $V/\lambda^3$ , where  $V$  is the volume of the media and  $\lambda$  is the wavelength of the light used. This limit suggests that densities of  $\sim 10^{12}$  bits per  $\text{cm}^3$  are possible. Unfortunately, due to the finite numerical aperture or the finite dynamic range of the material, practical systems are limited to approximately  $10^9$ – $10^{10}$  bits per  $\text{cm}^3$ . For example,  $10^3$  holograms each with  $10^6$  pixels has a total capacity of  $10^9$  bits. With the increase in storage densities of both magnetic and RAM memories, storage technologies that are limited to these capacities are not of commercial interest. In order to increase the capacity to useful levels it is necessary to spatially multiplex the

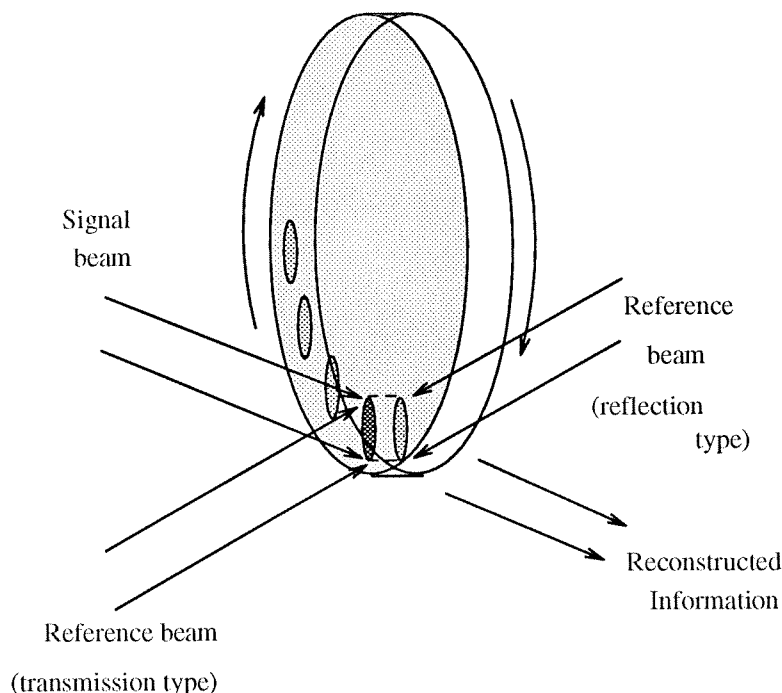


Figure 5.10: Using the disk as a storage device by presenting the reference plane wave for reconstruction of the stored data.

holograms as well as volumetrically multiplexing them. Spatial multiplexing simply means to use new, “unexposed” material to record more holograms.

The storage architectures presented here accomplishes the spatial multiplexing by shaping the medium in the form of a disk and rotating the disk or moving the location of the beams. At each spot on the holographic disk (HD), holograms are multiplexed by angle, wavelength, phase code, or in-plane rotation. The disk/head motion is used to access new spots to record the information just like in the correlator in the preceding section. To get the data out, the correct plane wave reference illuminates the correct spot on the disk. Figure 5.10 shows a disk system where holograms are multiplexed in either transmission or reflection geometries. In this section, we will calculate the geometrically limited storage capacity of such a system for angle, wavelength, and peristrophic plus angle multiplexing schemes. The derivation for angle and wavelength follows the derivation presented in reference [225] where the

signal beam is incident normal to the disk surface.

### 5.2.2 Angle Multiplexed Holographic Disk

The capacity of the angle multiplexed HD is

$$N = N_p^2 N_\theta N_s, \quad (5.10)$$

where  $N_p^2$  is the total number of pixels per hologram,  $N_\theta$  is the number of multiplexed holograms, and  $N_s$  is the number of spatial locations used. The number of angularly multiplexed holograms per location that one can store if the holograms are stored in the second minimum of the adjacent holograms sinc function is given by

$$N_\theta = 1 + \left( \frac{nL}{2\lambda} \right) |\cos \theta_1 - \cos \theta_2|. \quad (5.11)$$

In relation 5.11,  $n$  is the index of the material,  $L$  is the thickness of the material,  $\cos \theta_1$  is the smallest reference angle inside the material, and  $\cos \theta_2$  is the largest reference angle inside the material.

To calculate the number of unique locations in a given area  $A$ , we need to find the area of the individual spots. Figure 5.11 shows the geometry of the recording areas. For angle multiplexing there are two widths. One due to angular spread of the image, and one due to the angular spread of the reference beams. Using geometrical optics the highest frequency component of the image is traveling at an angle  $\theta = \sin^{-1}(\lambda/n\delta)$ , where  $\delta$  is the resolution or pixel spacing of the focused image on the disk.  $\delta$  in terms of system parameter is given by

$$\delta = \lambda \sqrt{4F^\# + 1}, \quad (5.12)$$

where  $F^\#$  is the F-number of the optical system. The width of the image inside the material is then

$$w = N_p \delta + L \tan \theta = N_p \delta + \frac{L}{\sqrt{(n\delta/\lambda)^2 - 1}}. \quad (5.13)$$

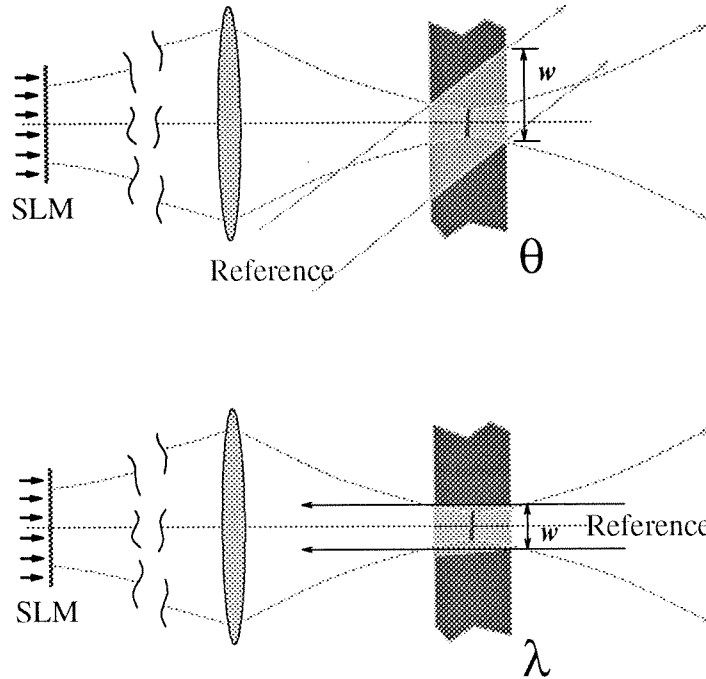


Figure 5.11: Area considerations for density calculation for both  $\theta$  and  $\lambda$  multiplexing.

As shown at the top of Figure 5.11, the reference beams occupy more area so the width in the direction of reference and signal intersection is given by

$$w' = w + L \tan \theta_2. \quad (5.14)$$

Therefore, the number of spots available is  $N_s = A/ww'$ .

The density of the information is given by dividing out the area from the total capacity. The expression is shown below.

$$N/A = N_p^2 \frac{1 + \left(\frac{nL}{2\lambda}\right) |\cos \theta_1 - \cos \theta_2|}{\left(N_p \delta + \frac{L}{\sqrt{(n\delta/\lambda)^2 - 1}}\right) \left(N_p \delta + \frac{L}{\sqrt{(n\delta/\lambda)^2 - 1}} + L \tan \theta_2\right)} \quad (5.15)$$

This equation has two optimal regions depending on  $L$ . Figure 5.12 shows the theoretical data storage limit for a photopolymer disk calculated using the Eq 5.15 in the thin regime. Also plotted is the number of holograms that have to be stored at a given location on the disk to get the density shown. Figure 5.12 was calculated with

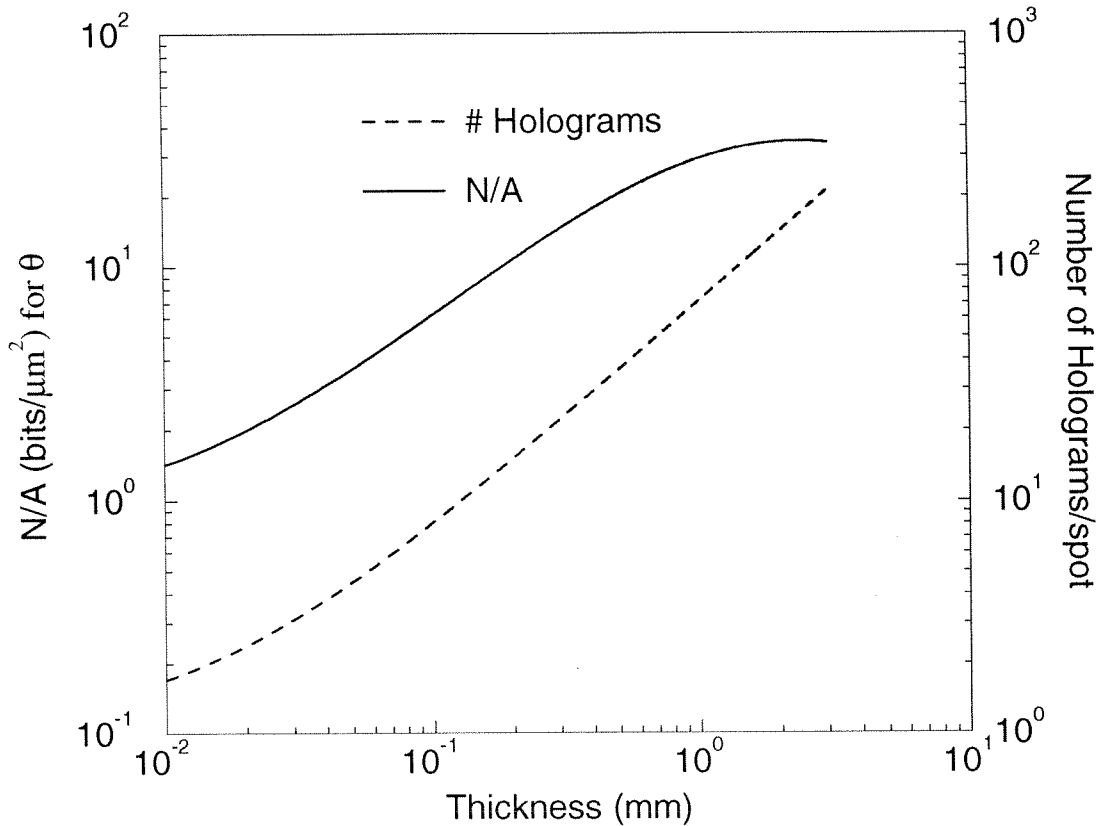


Figure 5.12: Geometrical storage limit of angle multiplexed HD.

$\lambda=488\text{nm}$ ,  $\theta_1 = 10^\circ$ ,  $\theta_2 = 20^\circ$ ,  $N_p = 1000$ ,  $n = 1.525$  and  $F^\# = 1$ . Notice that for a  $40\mu\text{m}$  thick film the optimal number of holograms is about 20 holograms/spot. Therefore, our demonstrated 10 angle multiplexed holograms would work out to about 6 bits/ $\mu\text{m}^2$ . In addition, angle multiplexing can be done on the other side of the signal beam and in reflection geometry which increases the storage density by a factor of four. Since current 2-D disks have densities on the order of 1 bit/ $\mu\text{m}^2$ , 3-D photopolymer disks can have densities that are about 100 times that of current disks depending on the film thickness. The density for Fourier multiplexed holograms is equivalent to the density for image plane holograms calculated here given the systems' bandwidths are equal.

### 5.2.3 Wavelength Multiplexed Holographic Disk

The capacity of a wavelength multiplexed HD is

$$N = N_p^2 N_\lambda N_s, \quad (5.16)$$

where  $N_p^2$  is the total number of pixels per hologram,  $N_\lambda$  is the number of multiplexed holograms, and  $N_s$  is the number of spatial locations used. The number of wavelength multiplexed holograms per location that one can store if the holograms are stored in the fourth minimum of the adjacent holograms sinc function is given by

$$N_\lambda = 1 + \frac{nL}{2} \left( \frac{1}{\lambda_1} - \frac{1}{\lambda_2} \right), \quad (5.17)$$

where the wavelength was swept from  $\lambda_1$  to  $\lambda_2$  to multiplex the holograms.

As the bottom of Figure 5.11 shows, wavelength multiplexing's area per location is smaller since the reference beam is interfered in reflection geometry. The maximum area is just the signal beam area with  $\lambda = \lambda_2$ . The number of spots on the disk for lambda becomes

$$w = N_p \delta + \frac{L}{\sqrt{(n\delta/\lambda_2)^2 - 1}}. \quad (5.18)$$

Thus the density for wavelength multiplexed HD can be written as

$$N/A = N_p^2 \frac{1 + \frac{nL}{2} \left( \frac{1}{\lambda_1} - \frac{1}{\lambda_2} \right)}{\left( N_p \delta + \frac{L}{\sqrt{(n\delta/\lambda_2)^2 - 1}} \right)^2}. \quad (5.19)$$

Figure 5.13 shows the theoretical data storage limit for a photopolymer disk calculated using the Eq 5.15 in the thin regime. Also plotted is the number of holograms that have to be stored at a given location on the disk to get the density shown. Figure 5.13 was calculated with  $\lambda_1 = 500\text{nm}$ ,  $\lambda_2 = 550\text{nm}$ ,  $N_p = 1000$ ,  $n = 1.525$  and  $F^\# = 1$ . Notice that the density for wavelength multiplexing is higher than the transmission angle multiplexed HD since the spot area is less for wavelength.

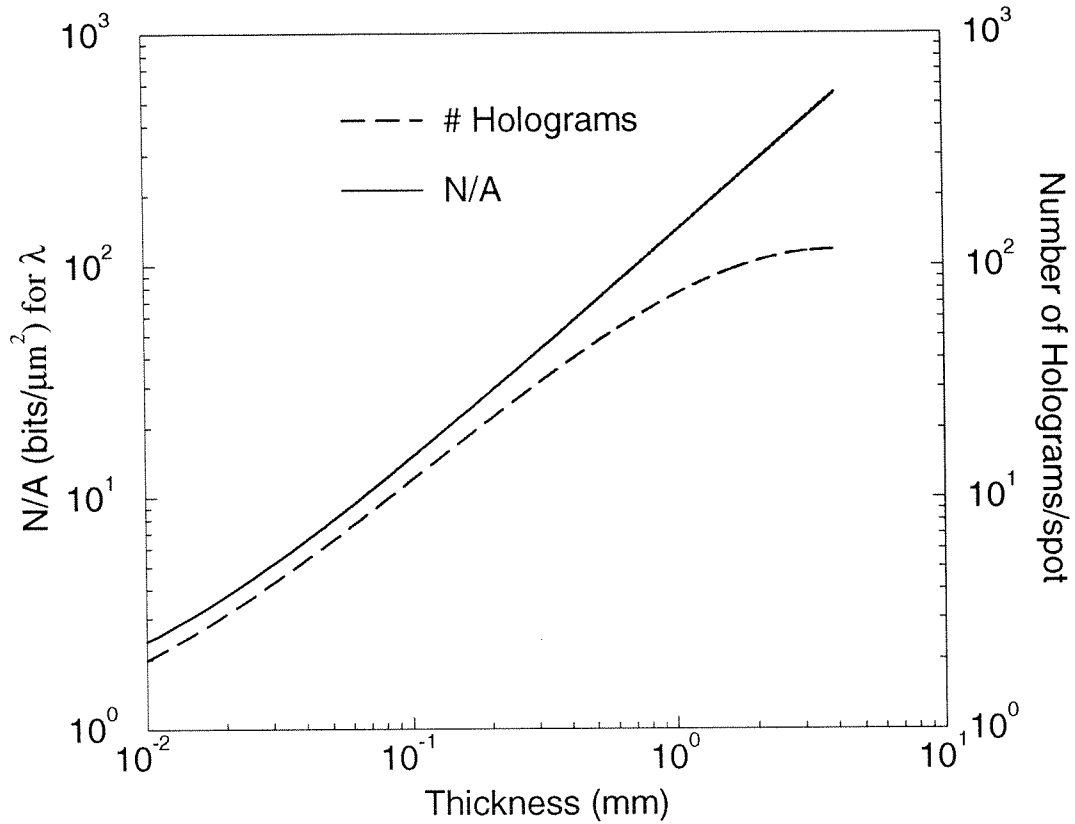


Figure 5.13: Geometrical storage limit of wavelength multiplexed HD.

### 5.2.4 Peristrophic and Angle Multiplexed Holographic Disk

The capacity of the peristrophic and the angle multiplexed HD is

$$N = N_p^2 N_\theta N_r N_s, \quad (5.20)$$

where  $N_p^2$  is the total number of pixels per hologram,  $N_\theta$  is the number of multiplexed holograms,  $N_r$  is the number of peristrophic multiplexed holograms possible, and  $N_s$  is the number of spatial locations used. For each peristrophic position, a new set of holograms can be angle multiplexed which makes the total number of holograms multiplexed at a given spot approximately equal to  $N_r \times N_\theta$ . The number of angularly multiplexed holograms per location that one can store if the holograms are stored in the second minimum of the adjacent holograms sinc function is again given by

$$N_\theta = 1 + \left( \frac{nL}{2\lambda} \right) |\cos \theta_1 - \cos \theta_2|. \quad (5.21)$$

In relation 5.21,  $n$  is the index of the material,  $L$  is the thickness of the material,  $\cos \theta_1$  is the smallest reference angle inside the material, and  $\cos \theta_2$  is the largest reference angle inside the material. To calculate the number of unique locations in a given area  $A$ , we need to find the area of the individual spots. For peristrophic and angle multiplexing the area of recording is a circle, and the number of spots available is given by

$$N_s = \frac{A}{\sqrt{3}} \frac{1}{[N_p \delta + \frac{L}{\sqrt{(\delta n/L)^2 - 1}} + 2L \tan \theta_2]^2}. \quad (5.22)$$

With peristrophic multiplexing, the number of holograms can depend on the bandwidth of the signal or the thickness of the film. I will use the Bragg condition worked out in the previous chapter with  $\theta_2 = 0^\circ$ . The amount of rotation required to place the holograms in the second minimum is

$$d\theta = \frac{2}{\sin \theta_1} \sqrt{\frac{2\lambda}{Ln}}. \quad (5.23)$$

The total rotation possible is  $\pi$ , since greater than  $\pi$  rotation results in grating degeneracy. Thus the number of holograms that can be rotation multiplexed is

$$N_r = \frac{\pi}{d\theta} = \frac{\pi \sin \theta_1}{2} \sqrt{\frac{Ln}{2\lambda}}. \quad (5.24)$$

The storage density of the information is given by dividing out the area from the total capacity. The expression is shown below.

$$N/A = N_p^2 \frac{\pi \sin \theta_1}{2\sqrt{3}} \sqrt{\frac{Ln}{2\lambda}} \frac{(1 + \left(\frac{nL}{2\lambda}\right) |\cos \theta_1 - \cos \theta_2|)}{[N_p \delta + \frac{L}{\sqrt{(\delta n/L)^2 - 1}} + 2L \tan \theta_2]^2}. \quad (5.25)$$

Figure 5.14 shows the theoretical data storage limit for a photopolymer disk calculated using the Eq 5.25. Also plotted is the number of holograms that have to be stored at a given location on the disk to get the density shown and the density for just angle multiplexing without peristrophic. Figure 5.14 was calculated with  $\lambda=488\text{nm}$ ,  $\theta_1 = 20^\circ$ ,  $\theta_2 = 30^\circ$ ,  $N_p = 1000$ ,  $n = 1.525$ , and  $F^\# = 1$ . In addition, angle multiplexing



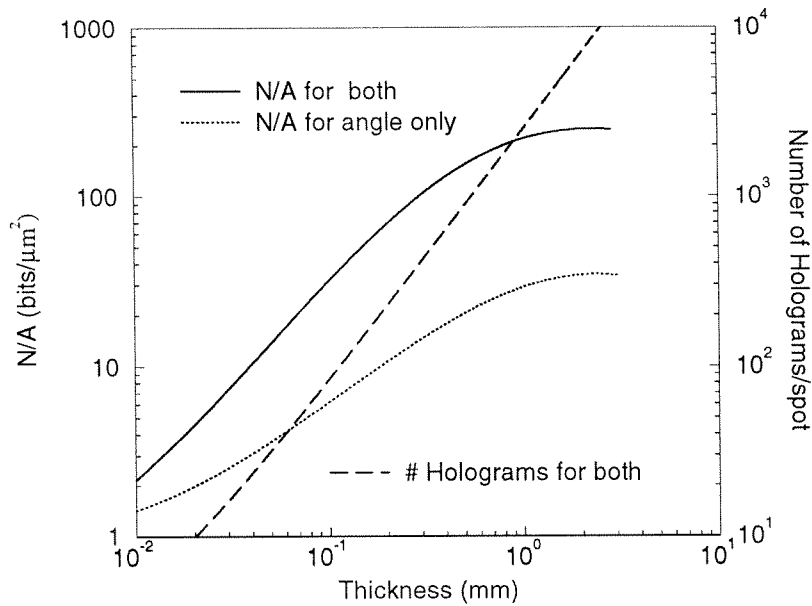


Figure 5.14: Geometrical storage limit of peristrophic and angle multiplexed HD.

can be done on the other side of the signal beam and in reflection geometry which adds another factor of four to the storage densities. Peristrophic multiplexing can help alleviate the bandwidth limitation inherent in multiplexing holograms in thin films and increase the density.  $N_r$  may also be further increased by making the signal beam incident at an angle. Thus, the density of 3-D disks using these techniques is not geometry limited but is limited by the dynamic range of the material used to fabricate the disk.

### 5.2.5 Read-out Rates

Data read-out rates are determined by the power of the light source, diffraction efficiency of the holograms, the size of the SLM, noise sources in the system, and the required signal-to-noise ratio. Li [226] has shown that with Fourier transformed holograms and a disk rotation rate of 3,600 rpm for a 6cm radius disk the maximum

$M_e$	$\eta$	$I_{inc}$	$\tau$ ( $\mu\text{sec}$ )	Read-out Rate
1,000	$10^{-5}$	20 W	2	500 Gbits/sec
100	$10^{-5}$	2 W	2	500 Gbits/sec
100	$10^{-4}$	200 mW	2	500 Gbits/sec
100	$10^{-4}$	400 mW	1	1 Tbits/sec

Table 5.2: Read-out Rate Parameters

alignment limited integration time is about  $2\mu\text{sec}$ . Thus, the transfer rate for a  $1,000 \times 1,000$  pixel SLM, assuming binary value for the pixels, is  $\sim 5 \times 10^{11}$  bits/second. Unfortunately, pixel registration is not the only factor that must be considered. A sufficient number of photons must be incident on the detector to ensure a reasonable SNR. The required incident intensity for a given number of electrons is expressed as

$$I_{inc} = \frac{N_p^2 M h c}{\tau \lambda \eta}, \quad (5.26)$$

where  $M$  is the number of required photons,  $\eta$  is the diffraction efficiency,  $N_p^2$  is the total number of pixels per hologram,  $\tau$  is the integration time,  $h$  is Plank's constant,  $c$  is the speed of light, and  $\lambda$  is the wavelength of the light. In order to get a feeling for the read-out rates and the required power for these devices. Table 5.2 shows read-out rates calculated with various diffraction efficiencies and photo-electrons with  $N_p^2 = 10^6$ , and  $\lambda = 500\text{nm}$ . The diffraction efficiency scales as one of the total number of holograms stored per spot squared. Thus, the capacity can be traded off for an increase in the data rate or for a lower power requirement as shown in the last two lines of Table 5.2. Since detector noise is typically between 20 to 60 photons for commercial detectors,  $M_e$  equal to 1,000 is probably too high while 100 is too low for a practical system. Clearly, with their inherent parallel read-out, 3-D memories can have fast read-out rate but requires relatively large light powers to achieve these rates.

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